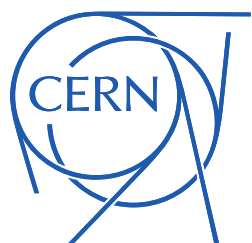


# Quo Vadis, Sterile Neutrino?

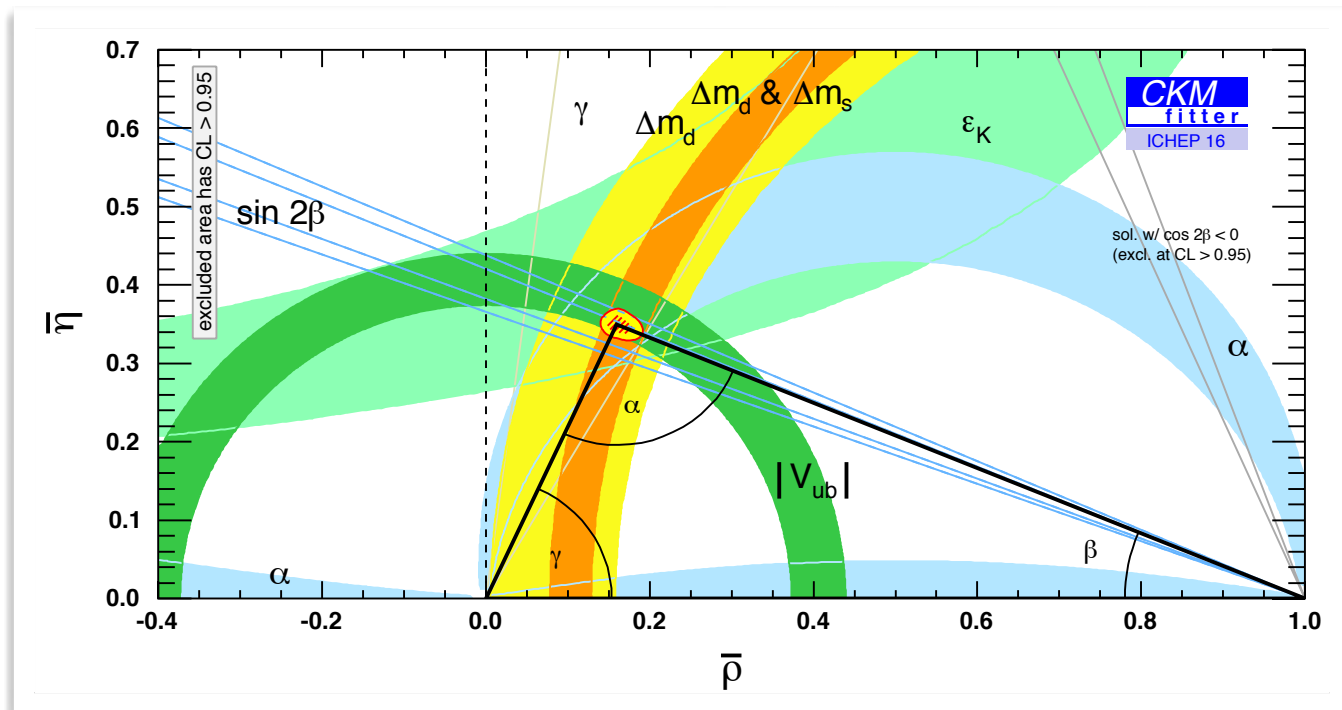
The Current Status of Searches for a Fourth Neutrino

Joachim Kopp (CERN & Uni Mainz)  
Neutrino 2022 | 30 May 2022



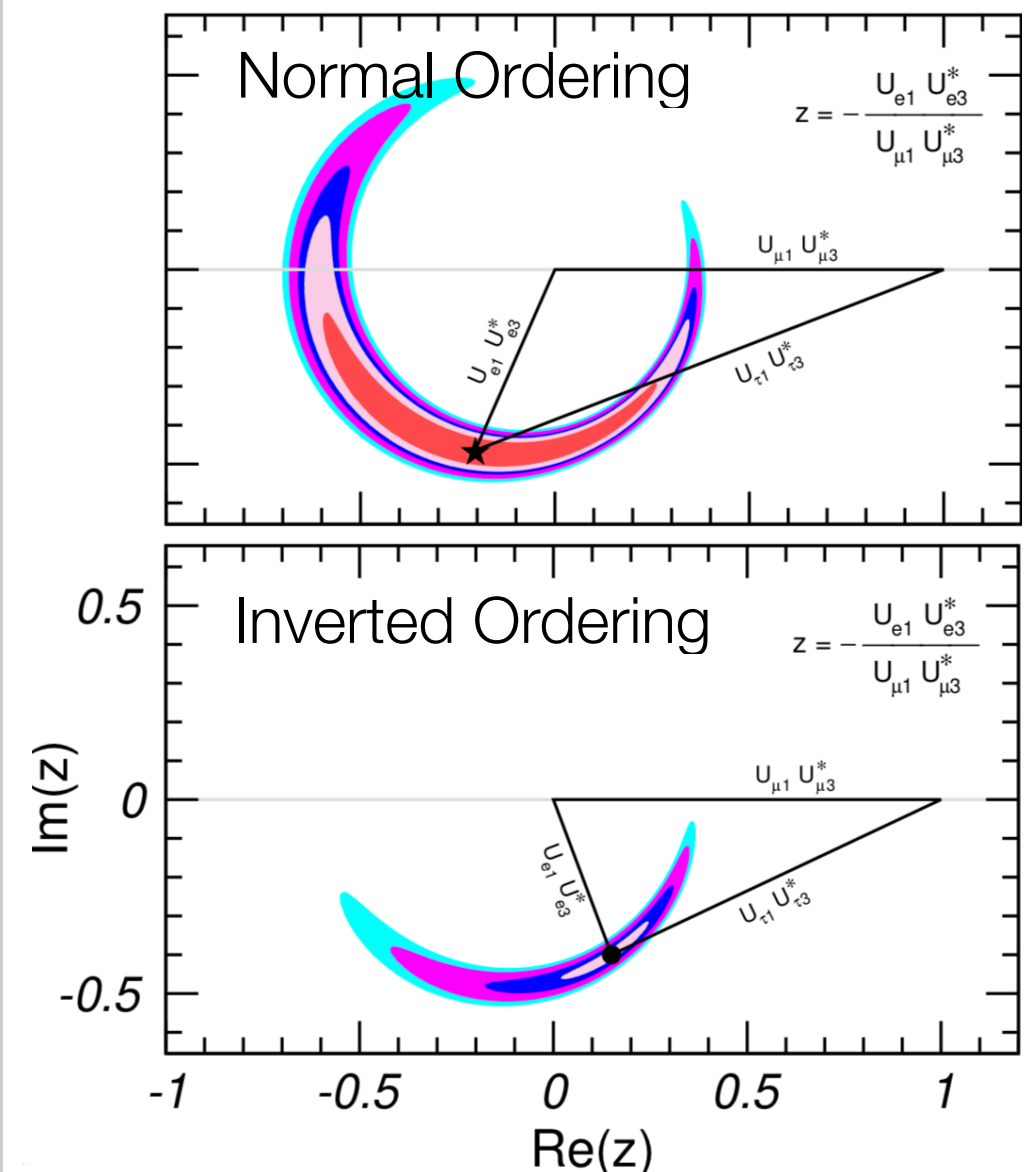
# Precision Flavour Physics

## Quarks



## Leptons

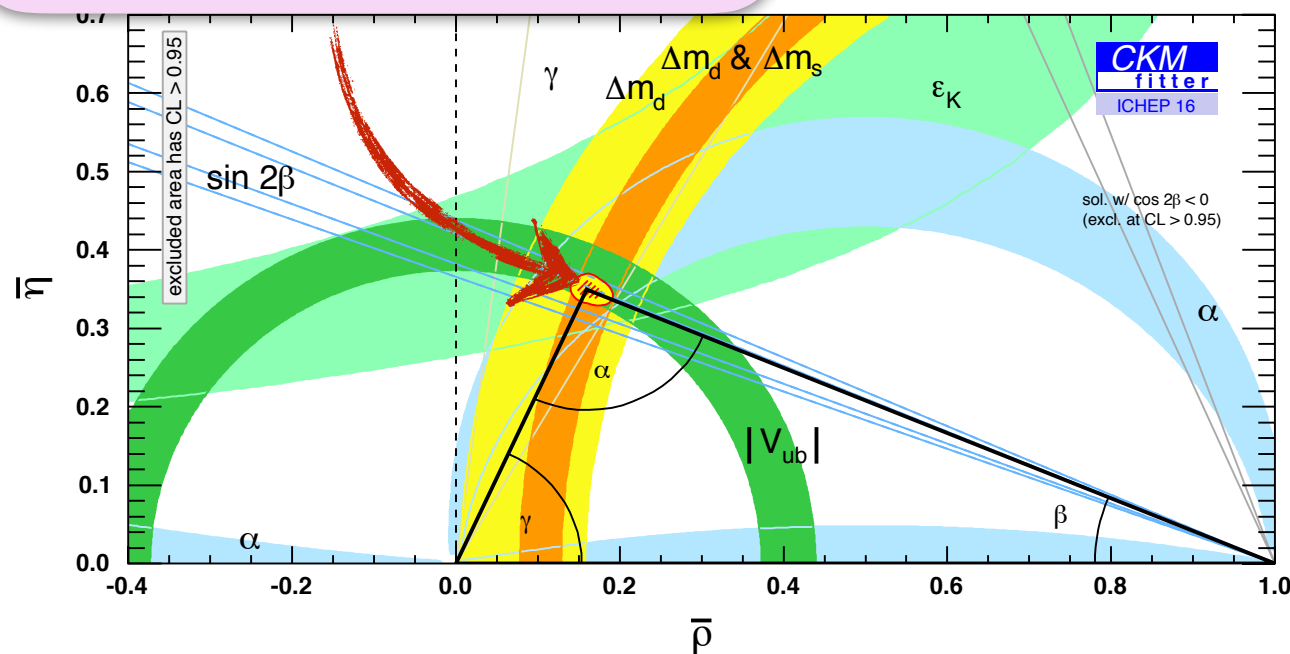
NuFIT 3.2 (2018)



# Precision Flavour Physics

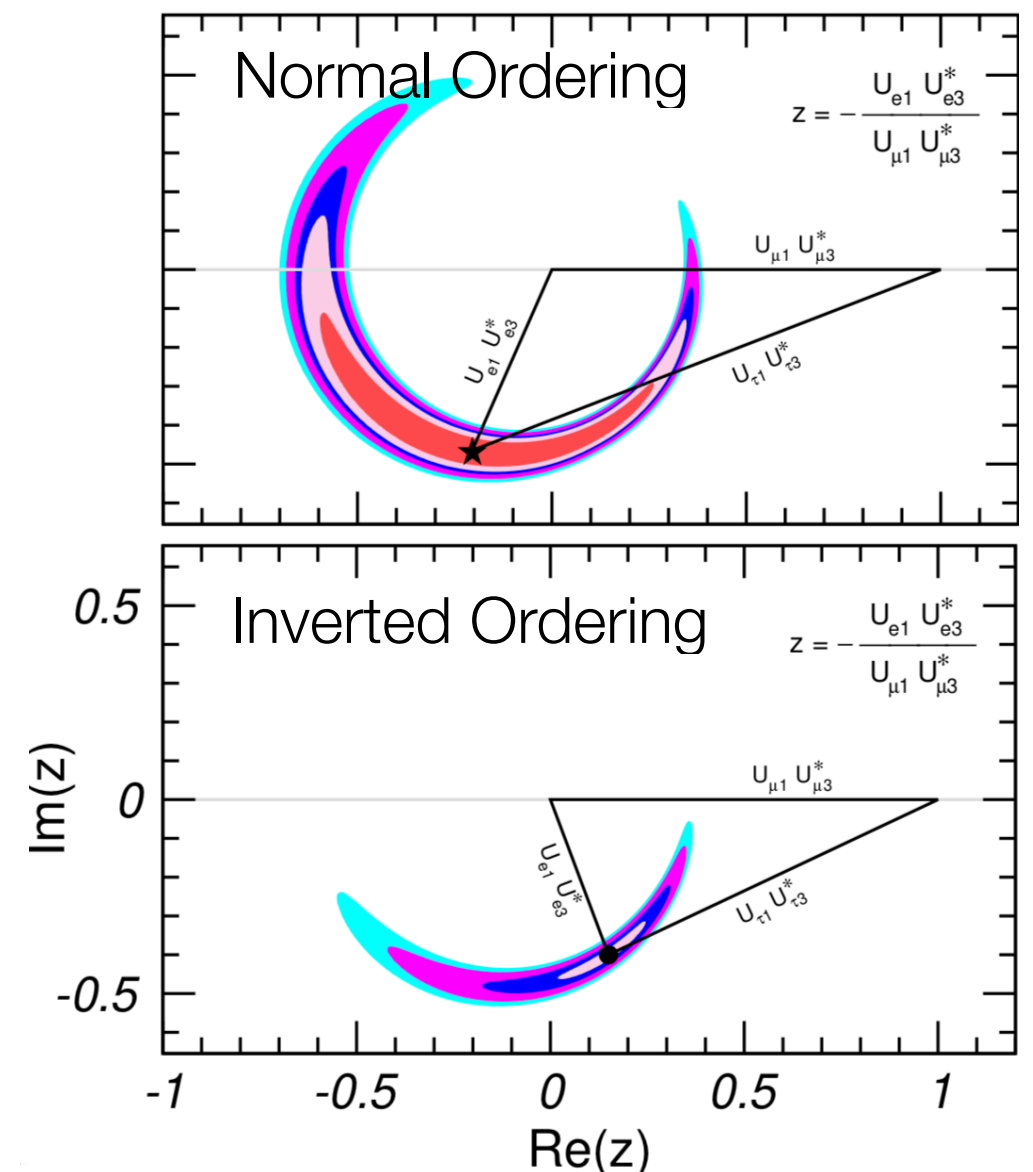
## Quarks

- ✓ superb precision
- ✓ excellent consistency



## Leptons

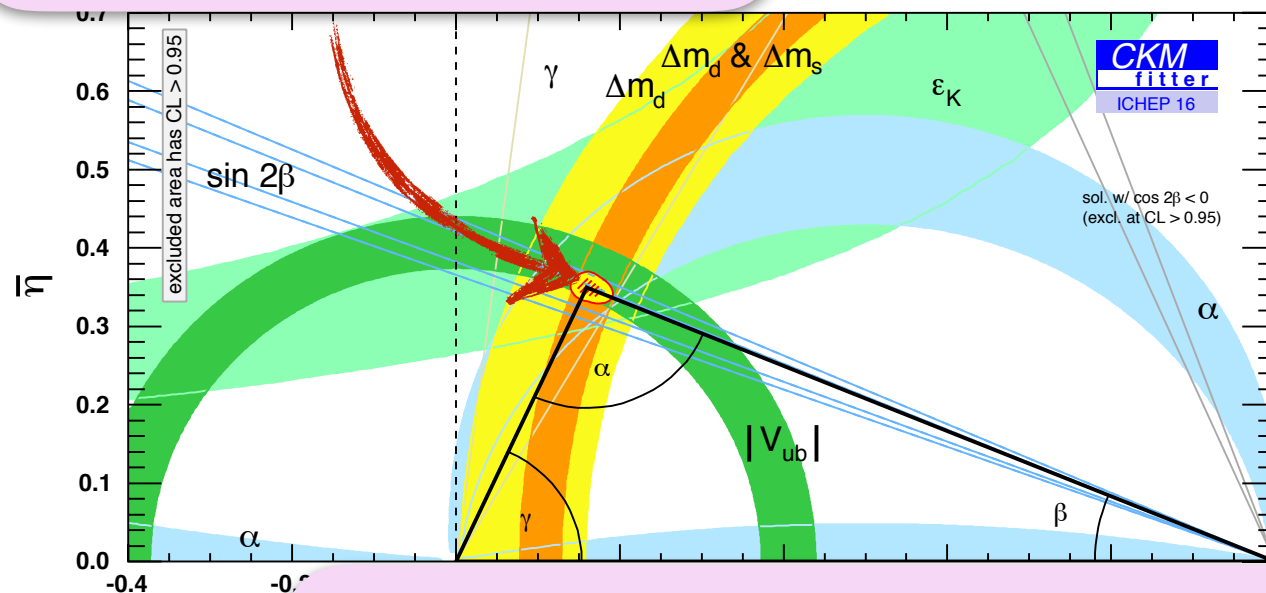
NuFIT 3.2 (2018)



# Precision Flavour Physics

## Quarks

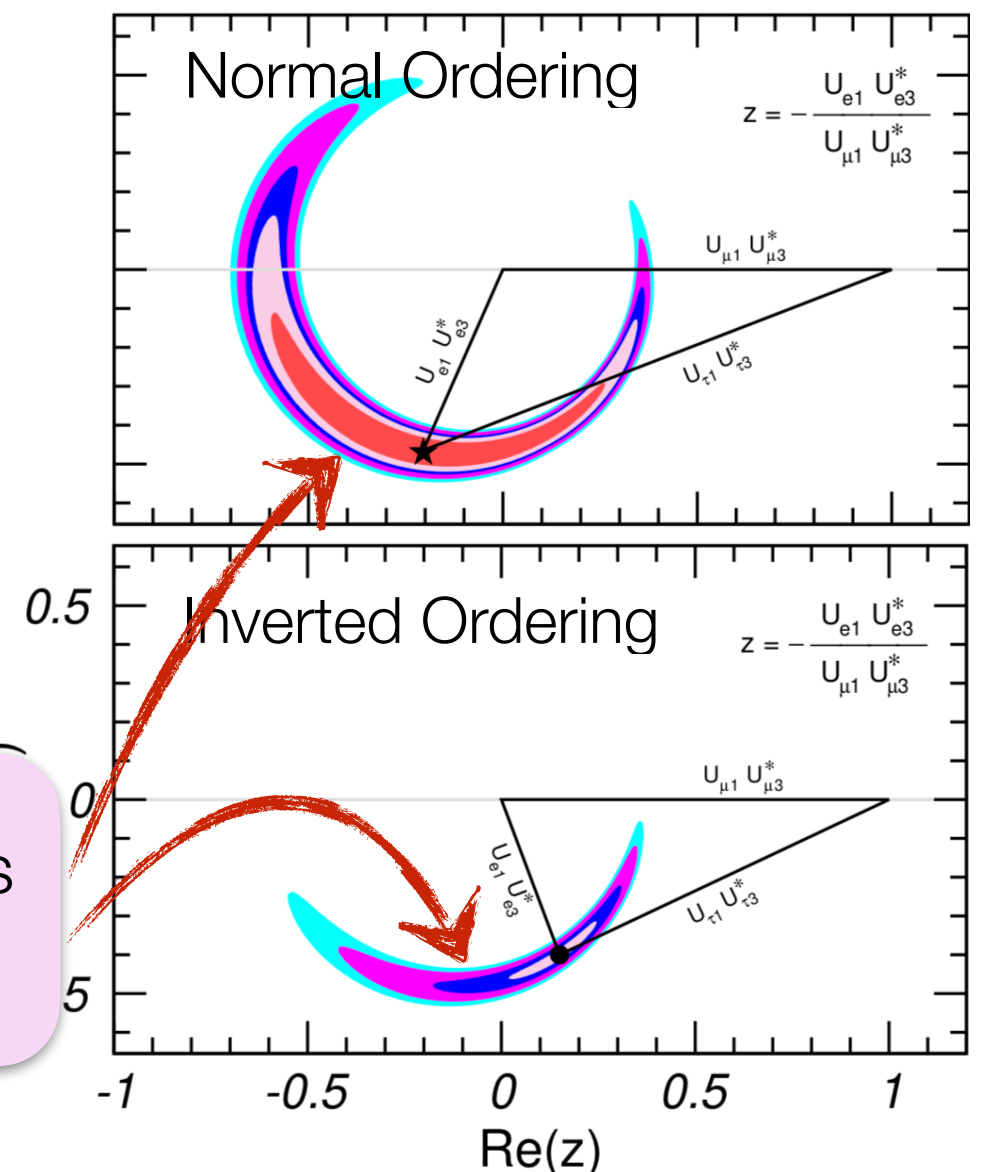
- ✓ superb precision
- ✓ excellent consistency



- ✓ < 10% accuracy in spite of tiny x-secs
- ✓ some parameters still unconstrained

## Leptons

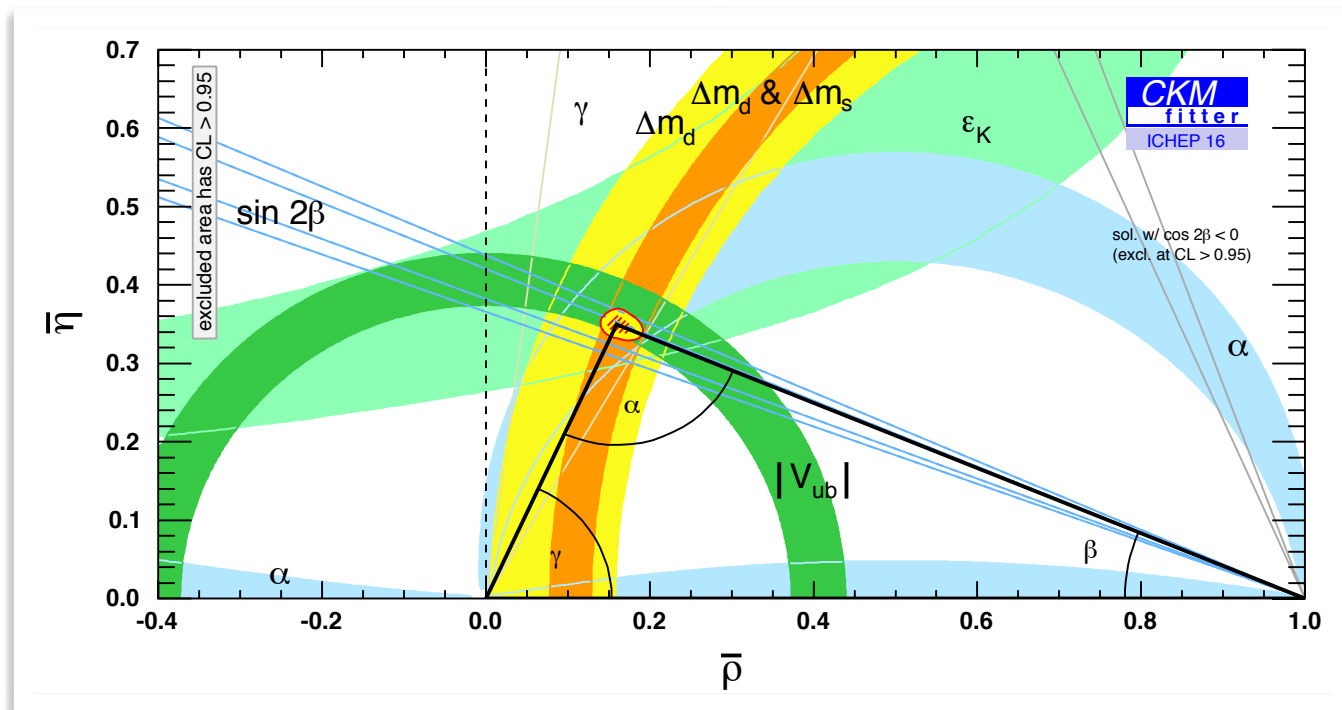
NuFIT 3.2 (2018)





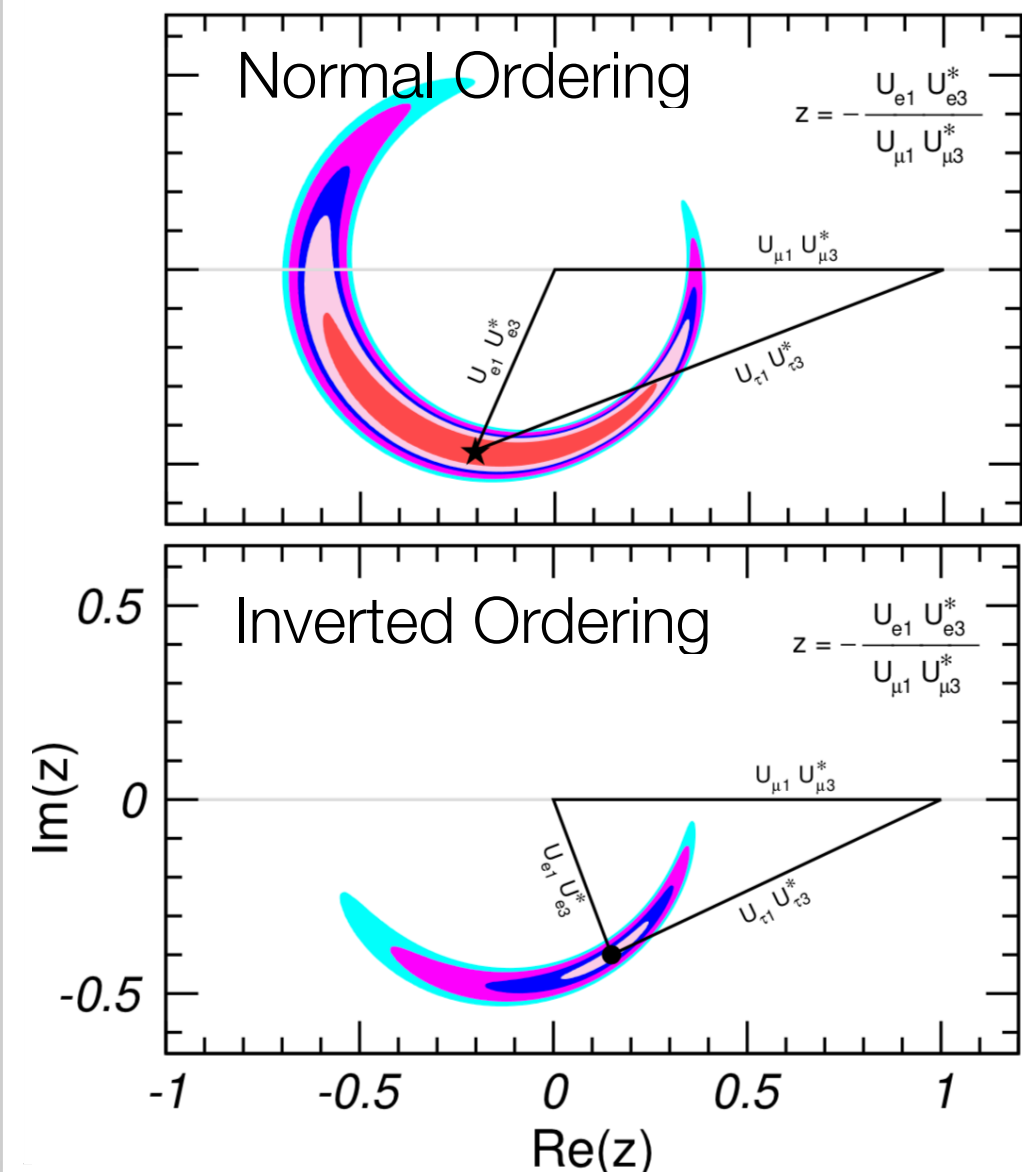
# Precision Flavour Physics

## Quarks



## Leptons

NuFIT 3.2 (2018)



Quarks

Leptons

**in both sectors:**

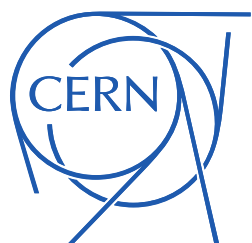
broadly consistent picture marred by  
intriguing anomalies

▮▮▮▮▮ ➡ this session

# Outline

- ☑ Short-Baseline Anomalies
- ☑ Standard Model Explanations for the MiniBooNE Anomaly?
- ☑ Sterile Neutrinos?

# Short-Baseline Anomalies

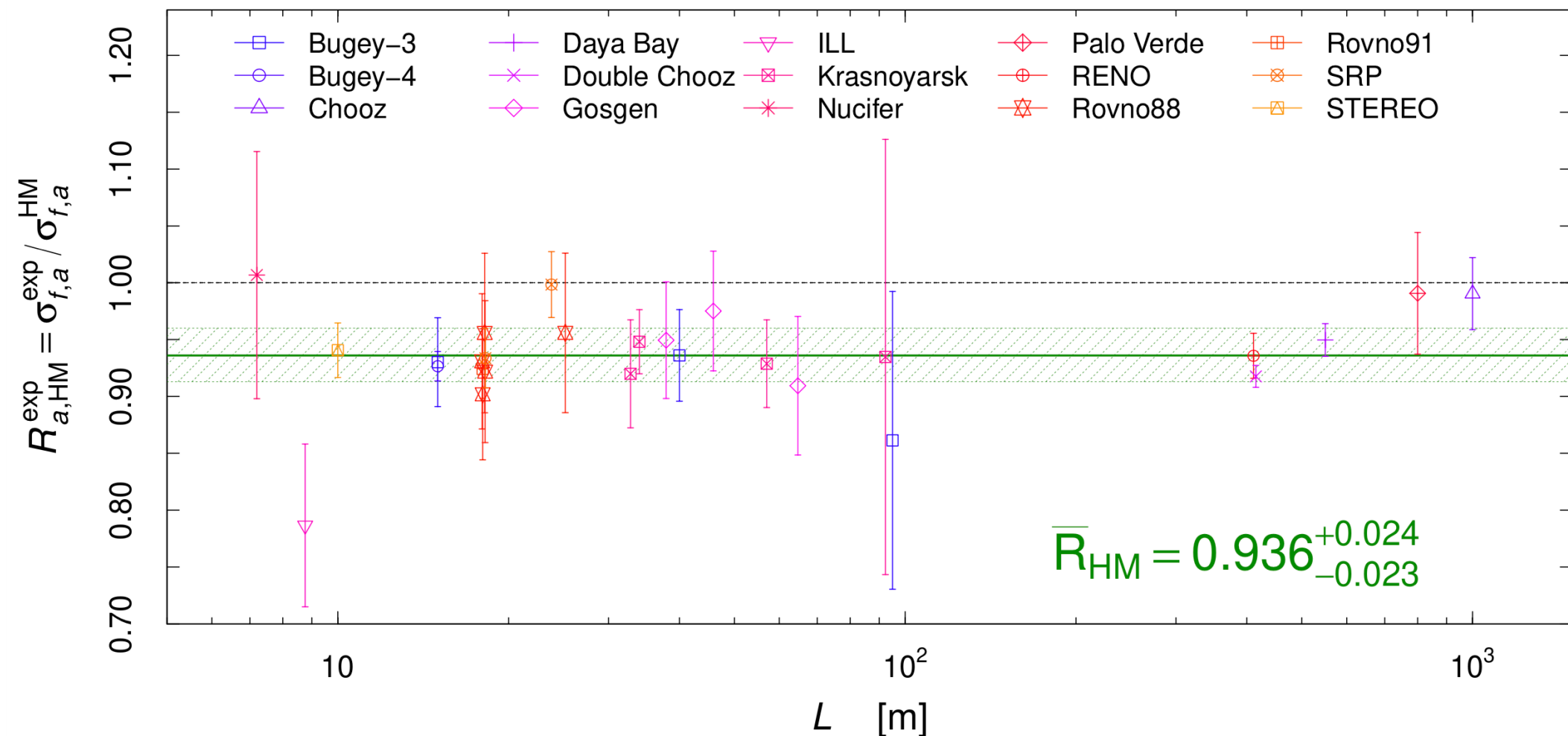


# Anomaly #1: Reactor Neutrino Fluxes

$\bar{\nu}_e$  flux from nuclear reactors  $\sim 3.5\%$  ( $\sim 3\sigma$ ) below prediction

⇒ oscillations of  $\bar{\nu}_e$  into **sterile neutrinos**  $\bar{\nu}_s$ ?  
(L/E too small for standard oscillations)

Mueller et al. [1101.2663](#), Huber [1106.0687](#), Giunti et al. [2110.06820](#)



Definition: sterile neutrino = SM singlet fermion

- ☑ Very generic extension of SM
    - can be leftover of extended gauge multiplet
  - ☑ Useful phenomenological tool
    - can explain  $\nu$  masses (seesaw mechanism,  $m \sim \text{TeV} \dots M_{\text{Pl}}$ )
    - can explain cosmic baryon asymmetry (leptogenesis,  $m \gg 100 \text{ GeV}$ )
    - can explain dark matter ( $m \sim \text{keV}$ )
    - can explain oscillation anomalies ( $m \sim \text{eV}$ )
- Promote mixing matrix to  $4 \times 4$ , oscillation formula unchanged:

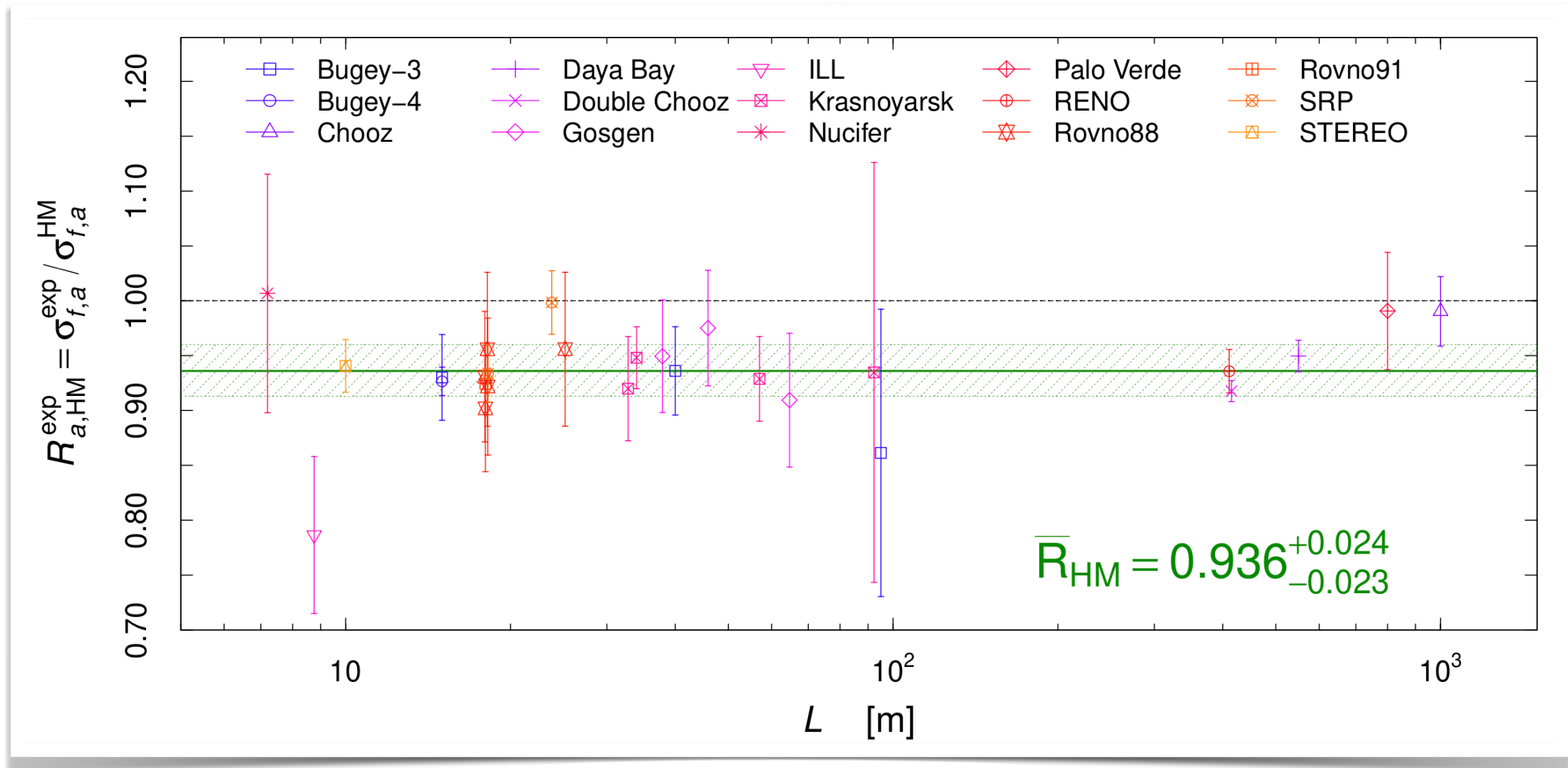
$$P_{\alpha \rightarrow \beta} = \sum_{j,k} U_{\alpha j}^* U_{\beta j} U_{\alpha k} U_{\beta k}^* \exp \left[ -i(E_j - E_k)T \right]$$





# Anomaly #1: Reactor Neutrino Fluxes

With updated input data to flux calculation  
(new  $\beta$  spectra from  $^{235}\text{U}$  fission)



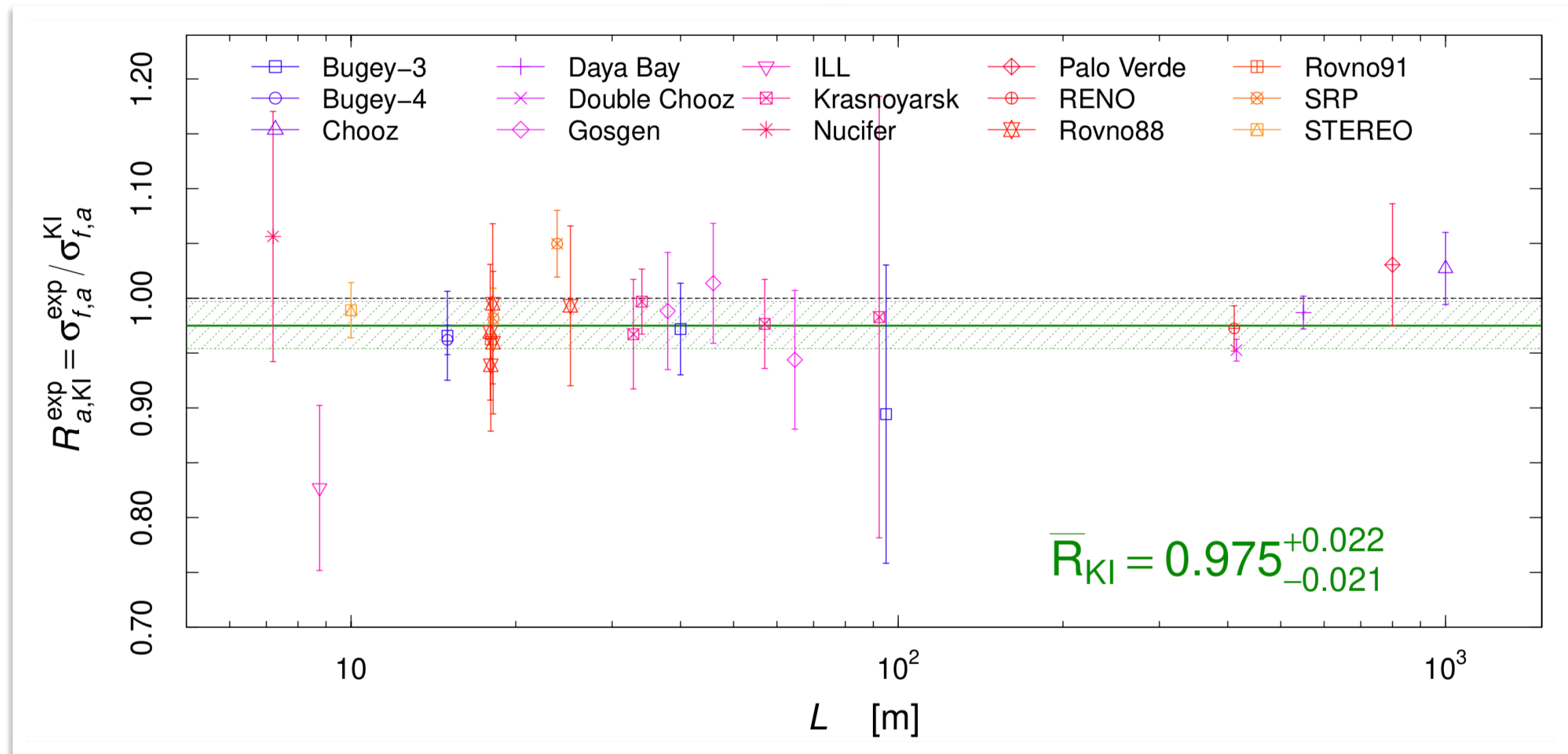
Kopeikin Skorokhvatov Titov [arXiv:2103.01684](https://arxiv.org/abs/2103.01684)

Berryman Huber [arXiv:2005.01756](https://arxiv.org/abs/2005.01756)

Giunti Li Ternes Xin [arXiv:2110.06820](https://arxiv.org/abs/2110.06820)

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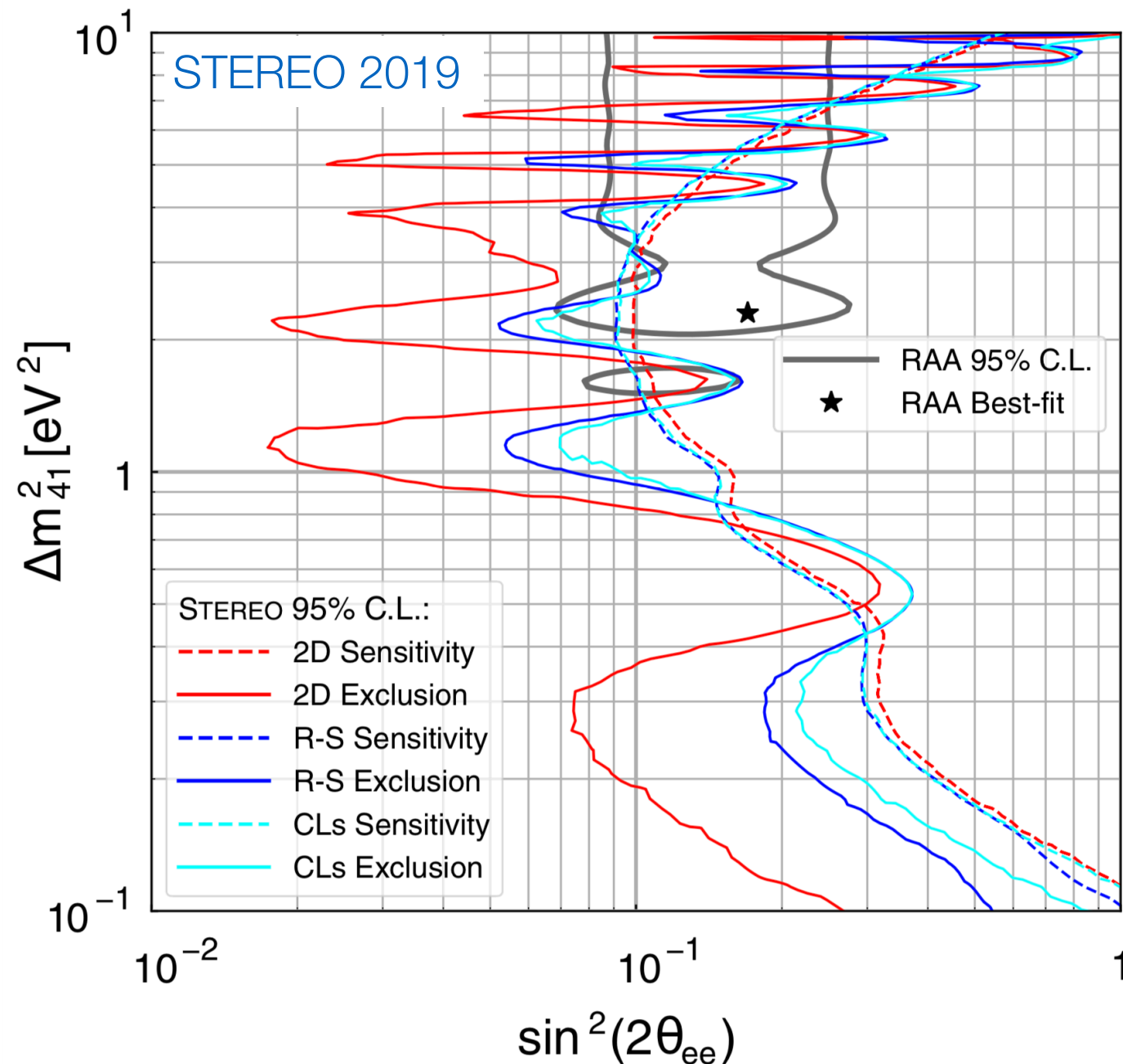


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# Anomaly #1: Reactor Neutrino Fluxes

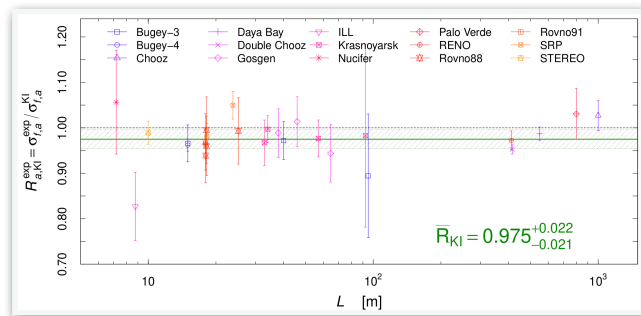


Moreover:

strong constraints  
from experiments  
comparing rates at  
different baselines

- ➡ independent of flux prediction
- ➡ see next talk  
(by Matthieu Licciardi)  
for updates

# Short-Baseline Anomalies

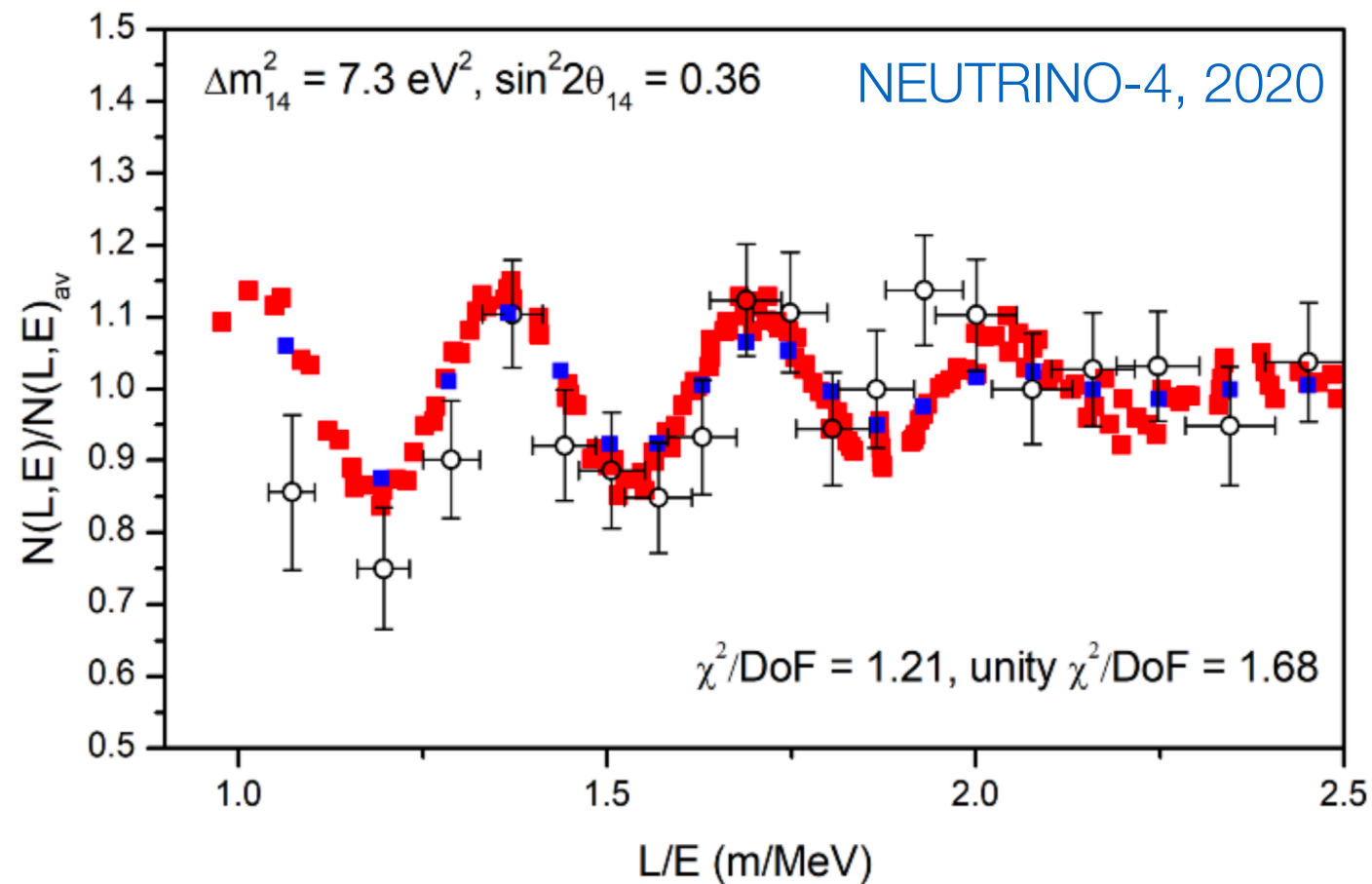


reactor flux anomaly  
resolved with new input data  
to flux calculation



Spectra

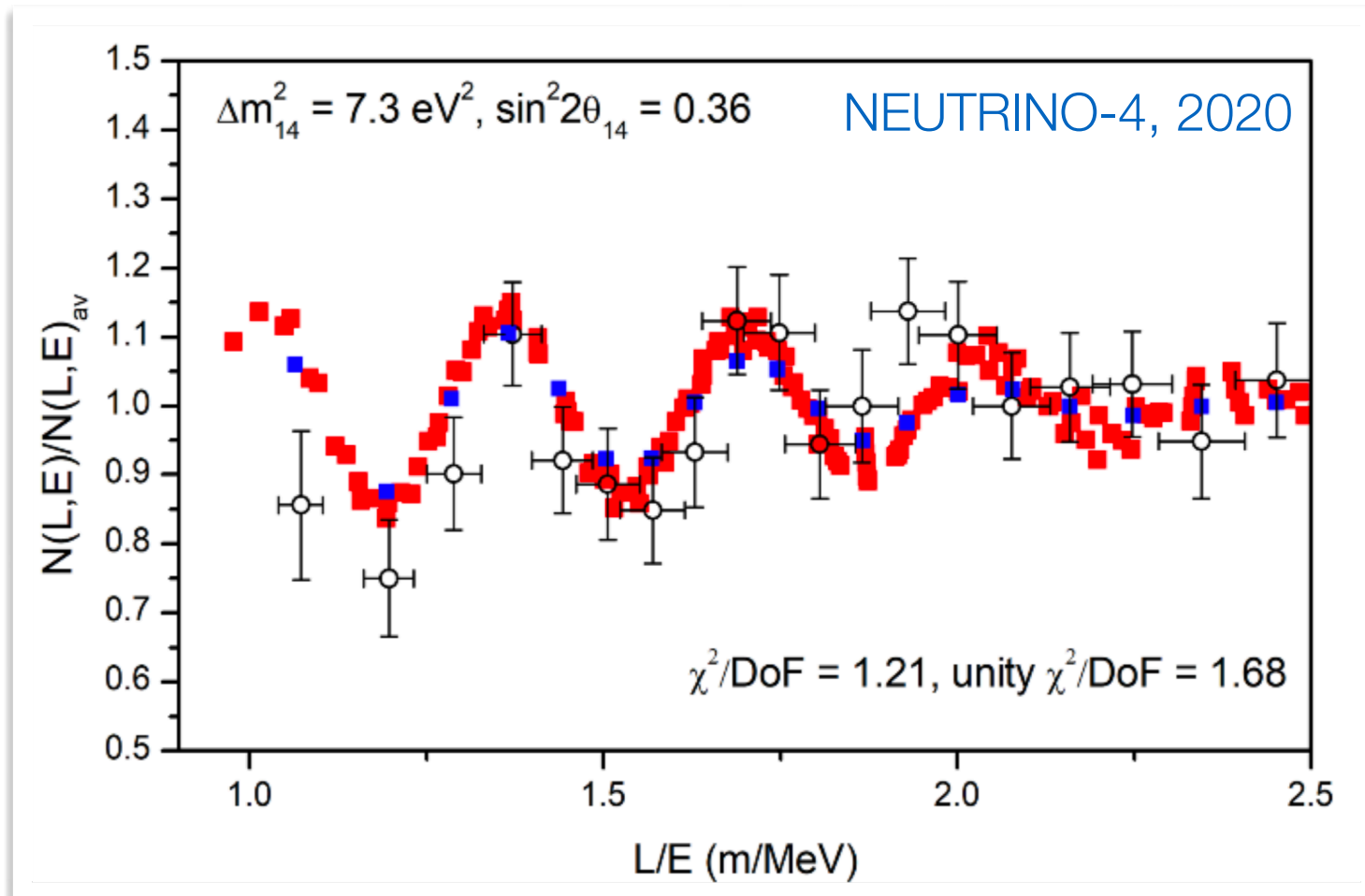
# Anomaly #2: Reactor Spectra



Almazán *et al.*, [arXiv:2006.13147](https://arxiv.org/abs/2006.13147)  
Berryman Coloma Huber Schwetz Zhou  
[arXiv:2111.12530](https://arxiv.org/abs/2111.12530)



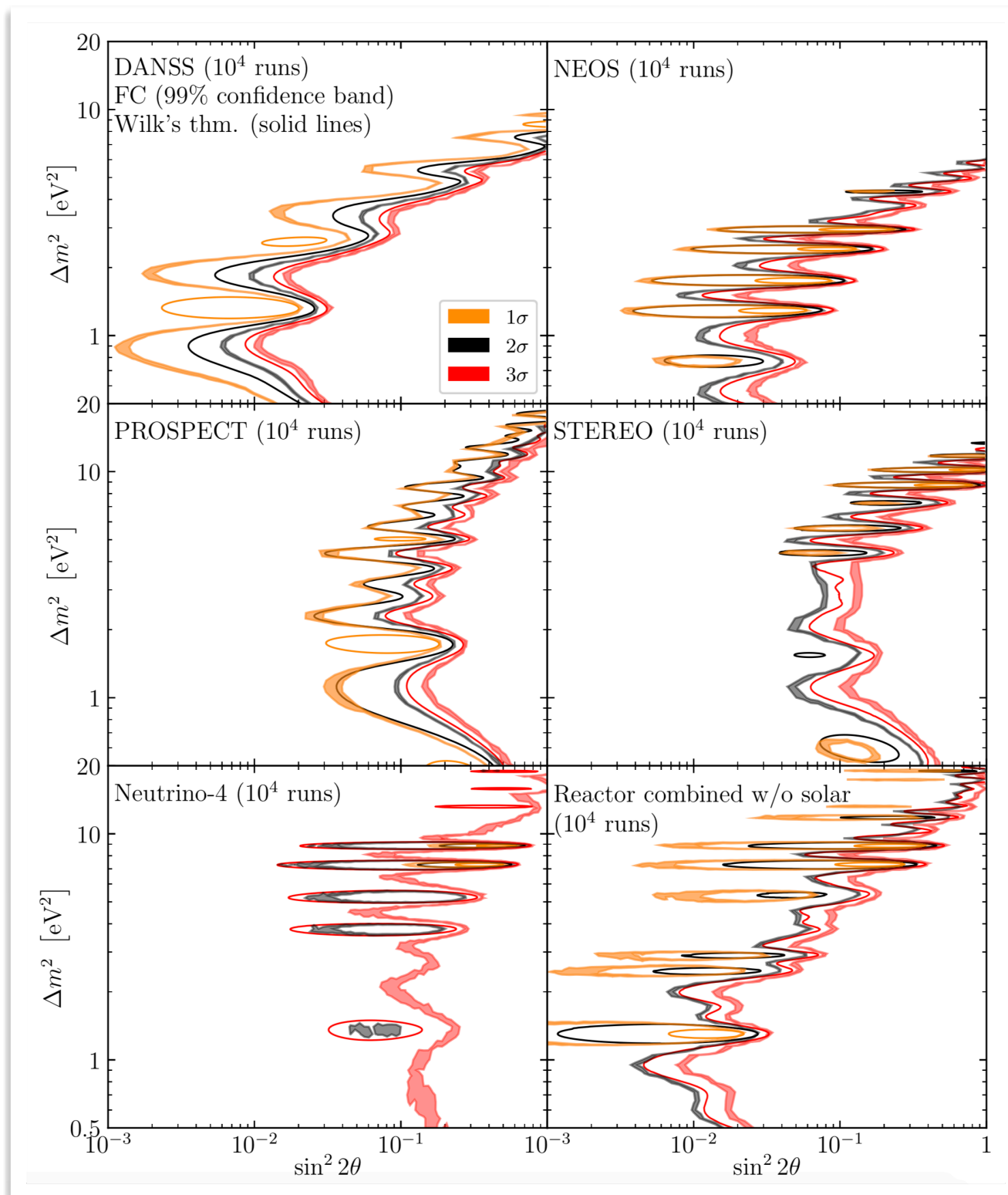
# Anomaly #2: Reactor Spectra



- ☑ spectral “wiggles” in several experiments
  - can be interpreted as neutrino oscillation signal
  - see talks by [Matthieu Licciardi](#) and [Jinyu Kim](#)
- ☑ Use ratios of spectra at different baselines
  - results independent of flux predictions

[Almazán et al., arXiv:2006.13147](#)  
[Berryman Coloma Huber Schwetz Zhou](#)  
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- ✓ But low statistical significance
  - Wilk’s theorem often fails!

*Almazán et al.*, [arXiv:2006.13147](https://arxiv.org/abs/2006.13147)  
Berryman Coloma Huber Schwetz Zhou  
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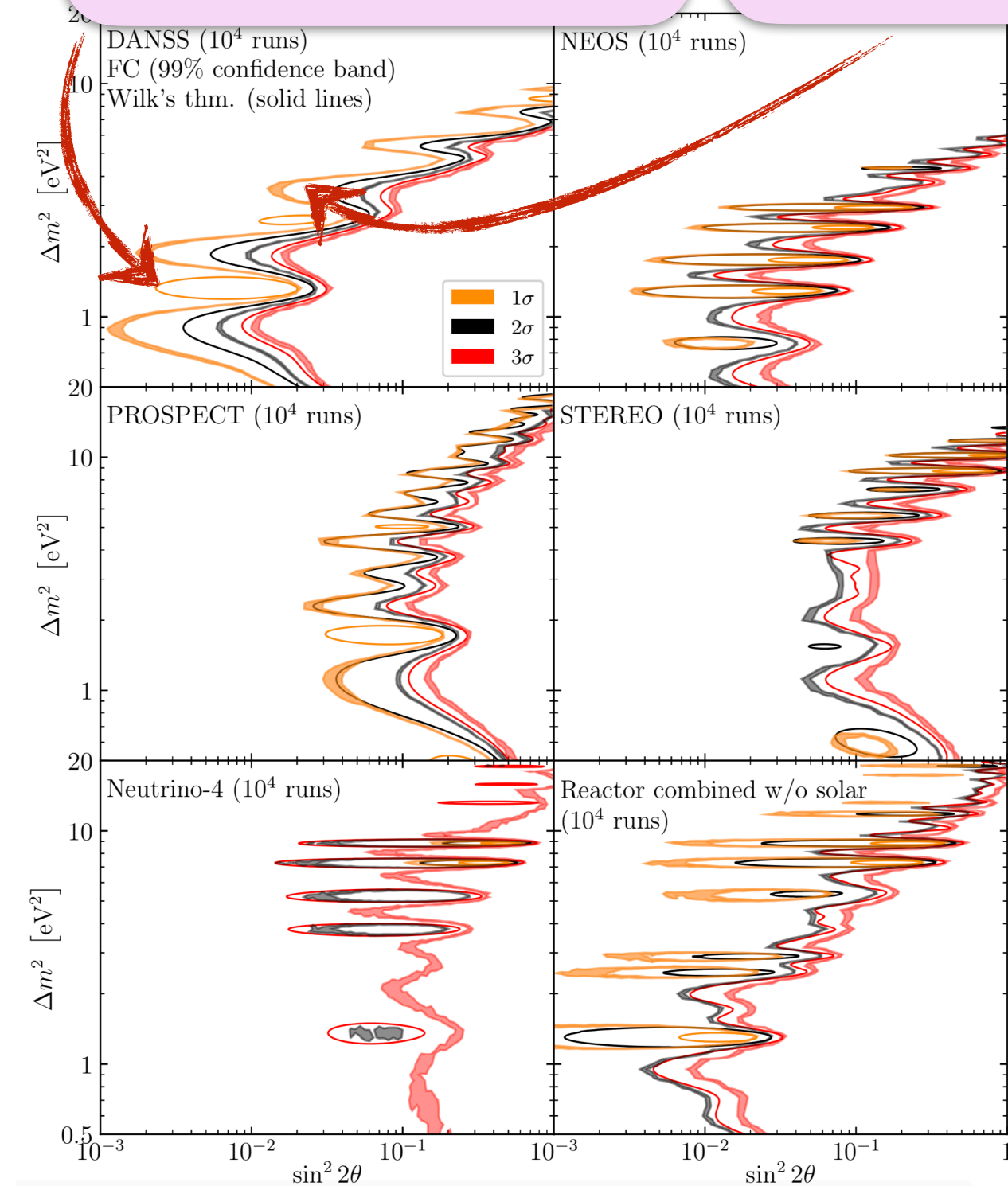
# Anomaly #2: Reactor Spectra

contours: Wilk's theorem

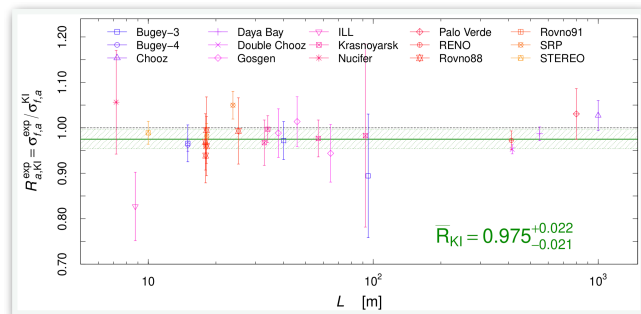
bands: Feldman–Cousins

- ✓ spectral “wiggles” in several experiments
  - can be interpreted as neutrino oscillation signal
  - see talks by **Matthieu Licciardi** and **Jinyu Kim**
- ✓ Use ratios of spectra at different baselines
  - results independent of flux predictions
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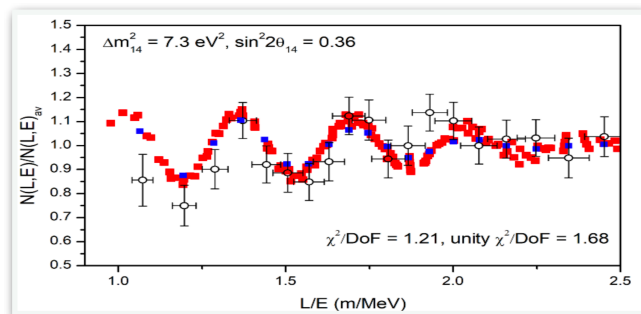
*Almazán et al.*, [arXiv:2006.13147](https://arxiv.org/abs/2006.13147)  
Berryman Coloma Huber Schwetz Zhou  
[arXiv:2111.12530](https://arxiv.org/abs/2111.12530)



# Short-Baseline Anomalies



reactor flux anomaly  
resolved with new input data  
to flux calculation



reactor spectra  
is there really an anomaly?



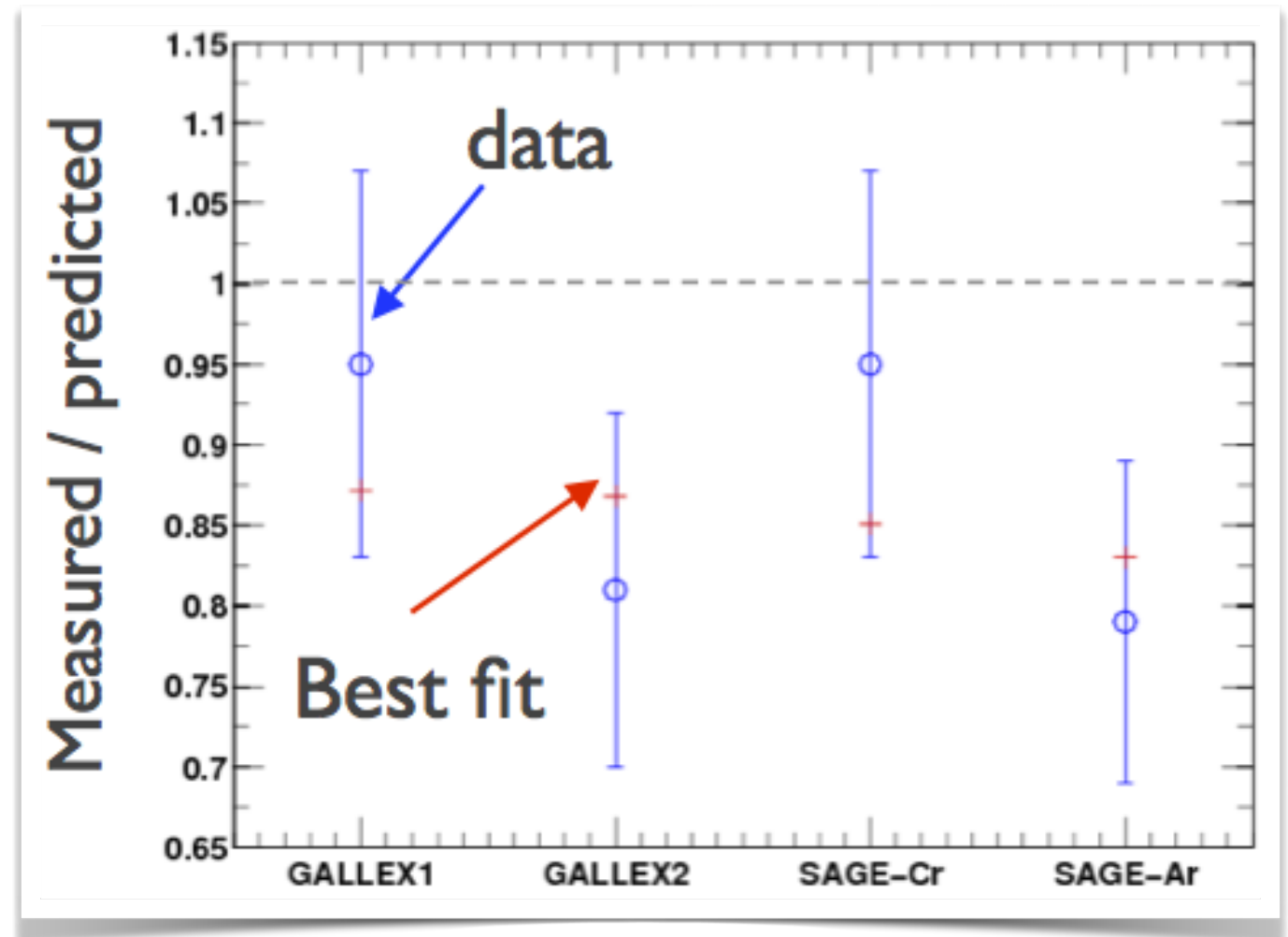
# The Gallium Anomaly

Once Upon a Time  
in the B.E.S.T.



# Anomaly #3: the Gallium Anomaly

- ✓ Experiments with intense radioactive sources
- ✓ Neutrino detection via
$${}^{71}\text{Ga} + \nu_e \rightarrow {}^{71}\text{Ge} + e^{-}$$
- ✓  $\sim 3\sigma$  deficit
- ✓  $\nu_e$  disappearance into sterile state?
- ✓ would require very large mixing (conflict with reactor observations)



Giunti Laveder [1006.3244](#)

# Anomaly #3: the Gallium Anomaly

- ☑ recently confirmed by BEST ( $\sim 4\sigma$ )

BEST [arXiv:2109.11482](https://arxiv.org/abs/2109.11482)

Barinov Gorbunov [arXiv:2109.14654](https://arxiv.org/abs/2109.14654)

- ☑ two independent target volumes (hoping to see oscillation pattern)

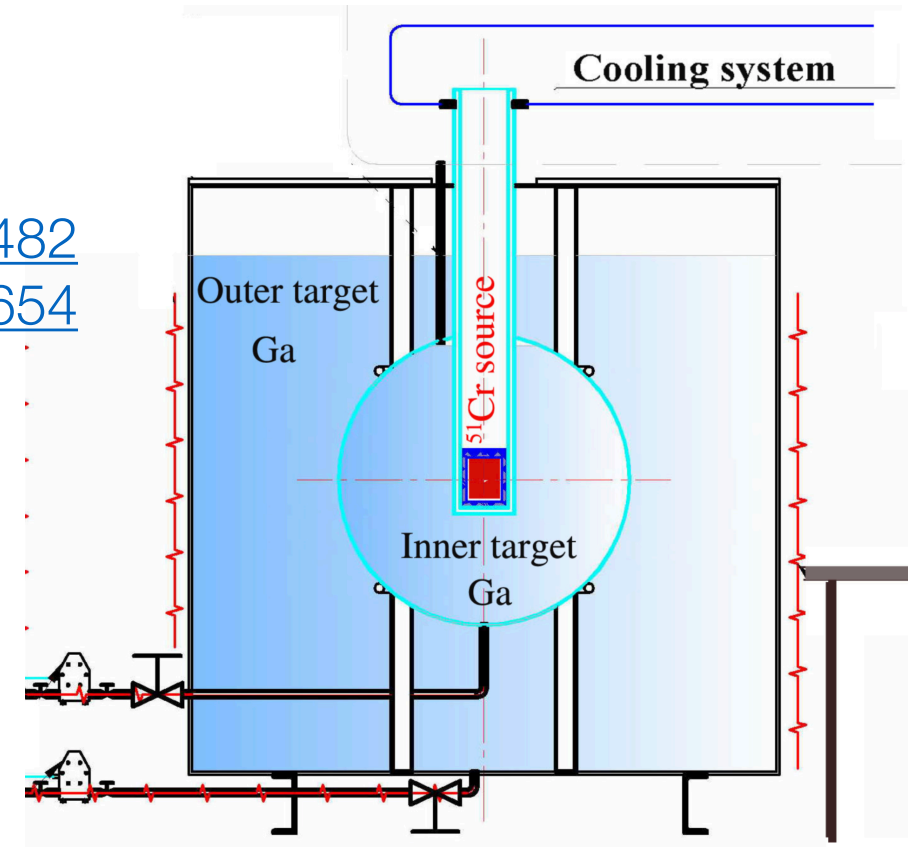
- ☑ radiochemistry similar to other gallium experiments (correlated systematics?)

- ☑ but robustness proven by

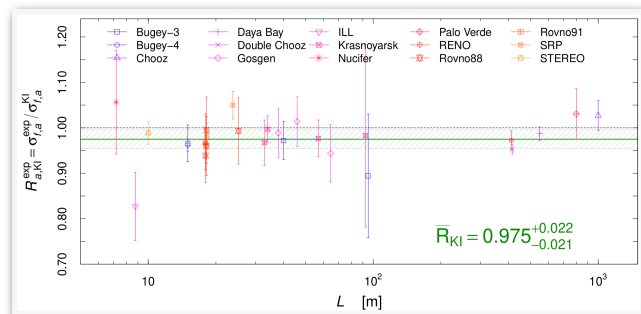
- sophisticated calibration [Cleveland et al. 2015](#)
- solar  $\nu$  measurements in agreement with other experiments

- ☑ uncertainty in x-sec to excited states of Ge-71

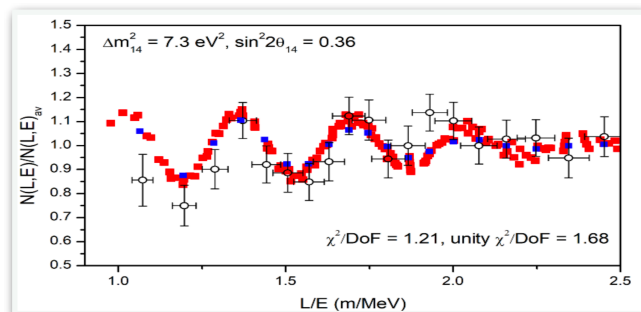
- but even ignoring excited states altogether, anomaly persists



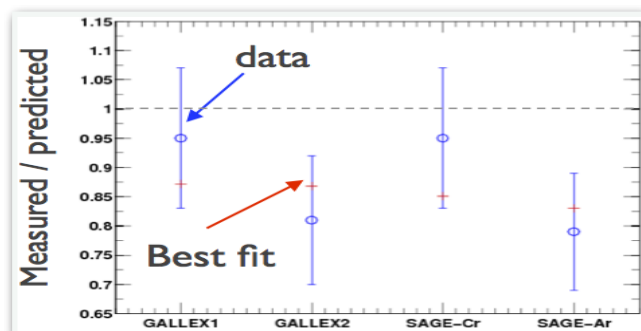
# Short-Baseline Anomalies



reactor flux anomaly  
resolved with new input data  
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reactor spectra  
is there really an anomaly?



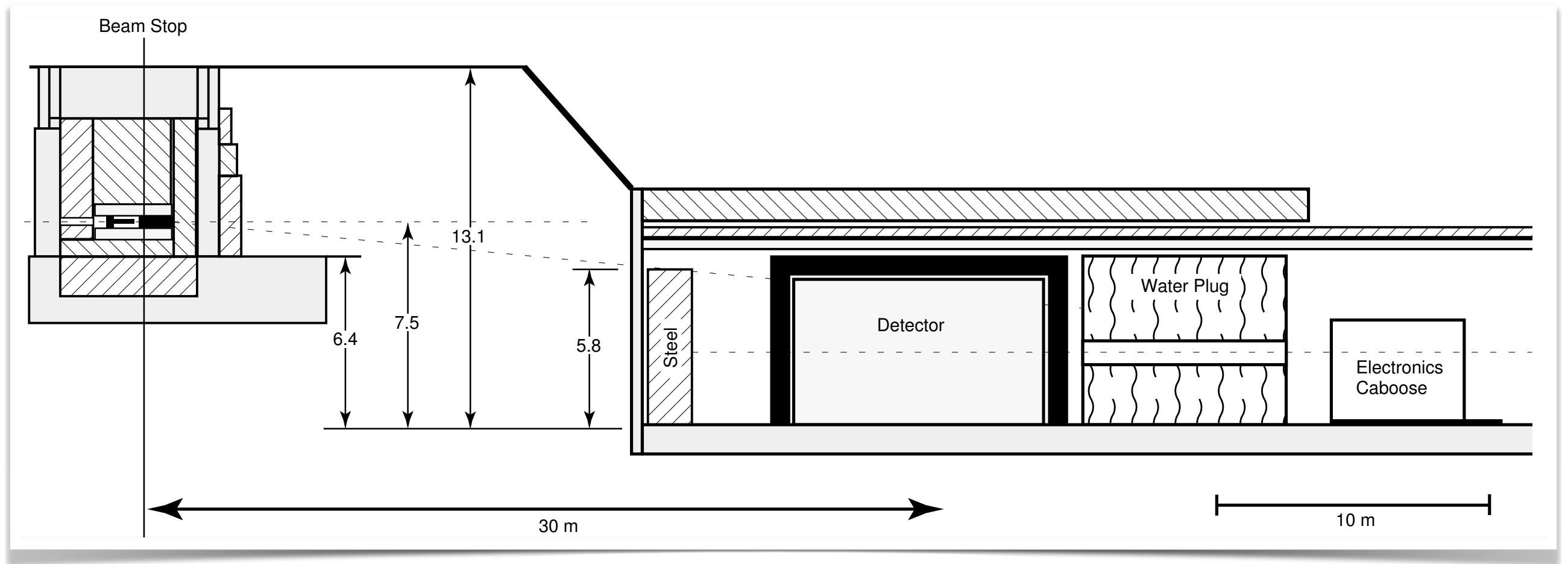
gallium anomaly  
unresolved, recently reinforced



LSND

# The Los Alamos Chamber of Secrets

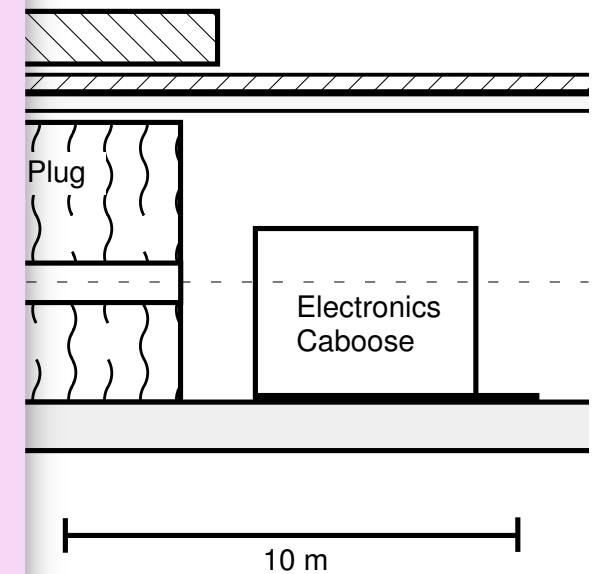
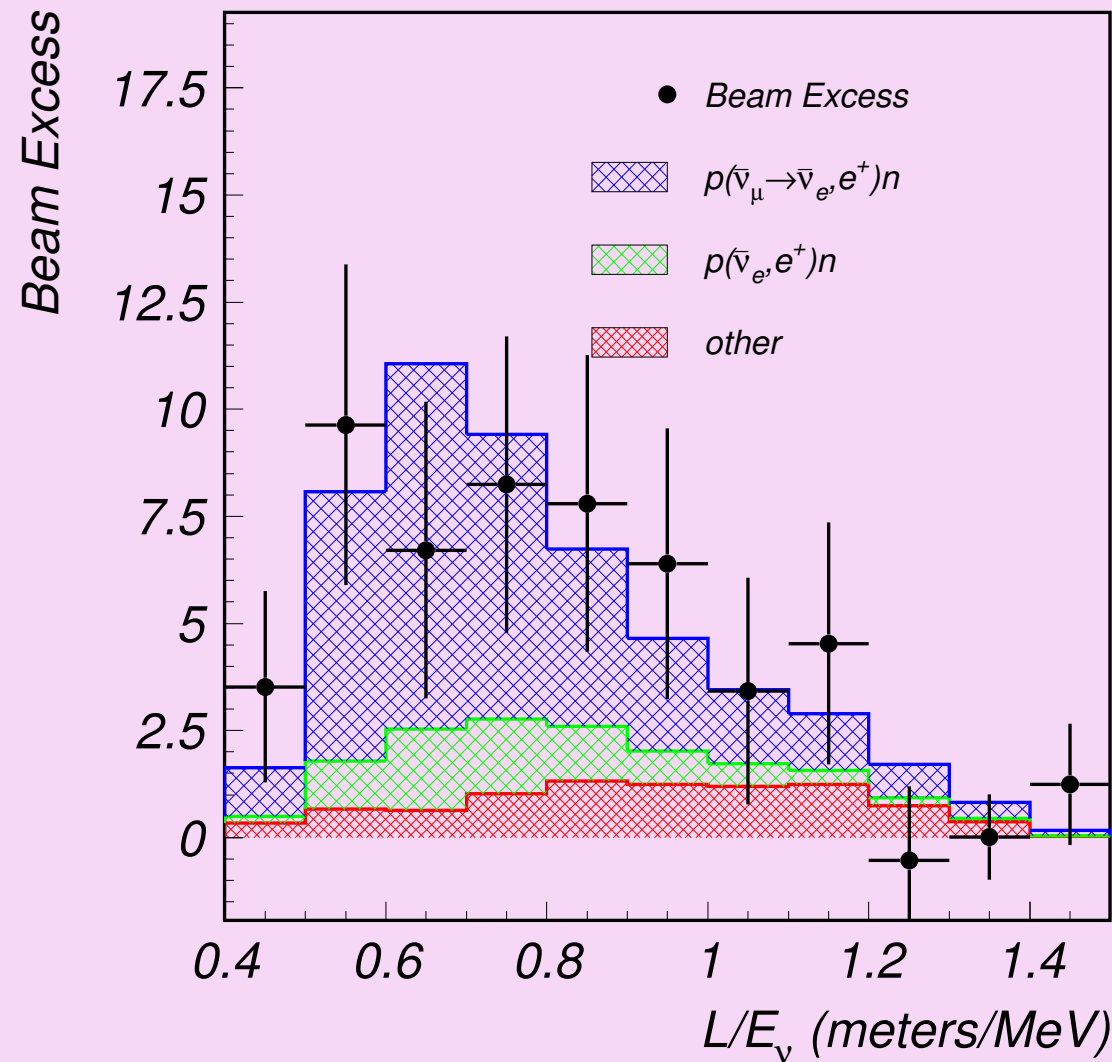
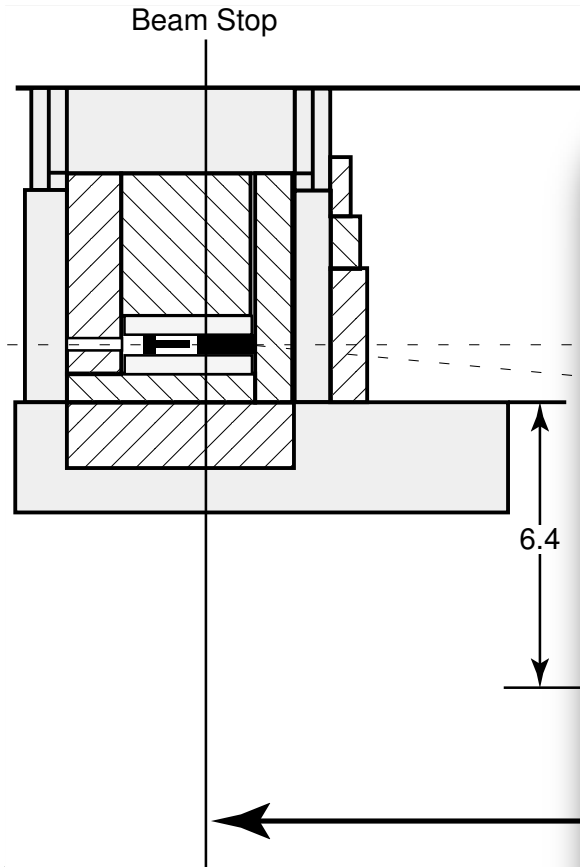
# Anomaly #4: LSND



- ✓  $\bar{\nu}_e$  appearance in a  $\bar{\nu}_\mu$  beam ( $\sim 3\sigma$ )
- ✓ Source—detector distance (“baseline”)  $\sim 30$  m
- ✓  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations?



# Anomaly #4: LSND



✓  $\bar{\nu}_e$  appearance

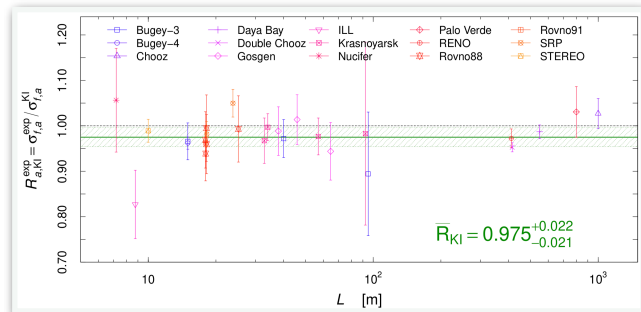
✓ Source—d

✓  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations?

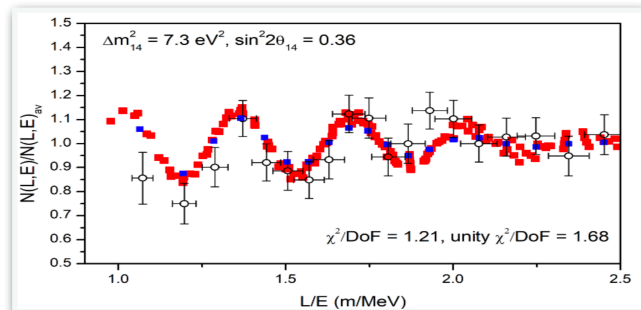
LSND Collaboration, [hep-ex/0104049](https://arxiv.org/abs/hep-ex/0104049)

m

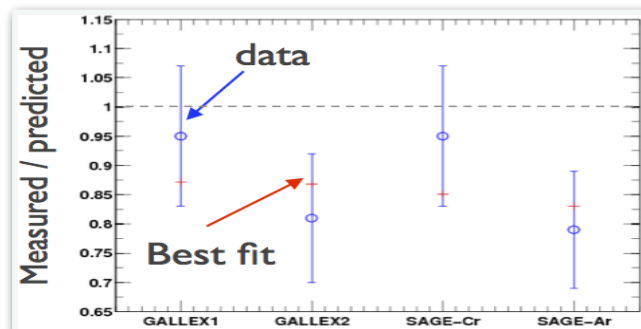
# Short-Baseline Anomalies



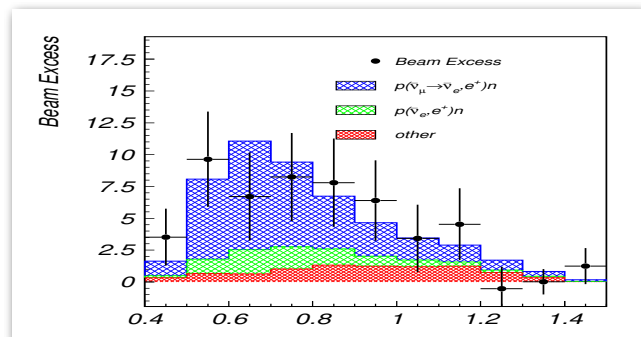
reactor flux anomaly  
resolved with new input data  
to flux calculation



reactor spectra  
is there really an anomaly?



gallium anomaly  
unresolved, recently reinforced

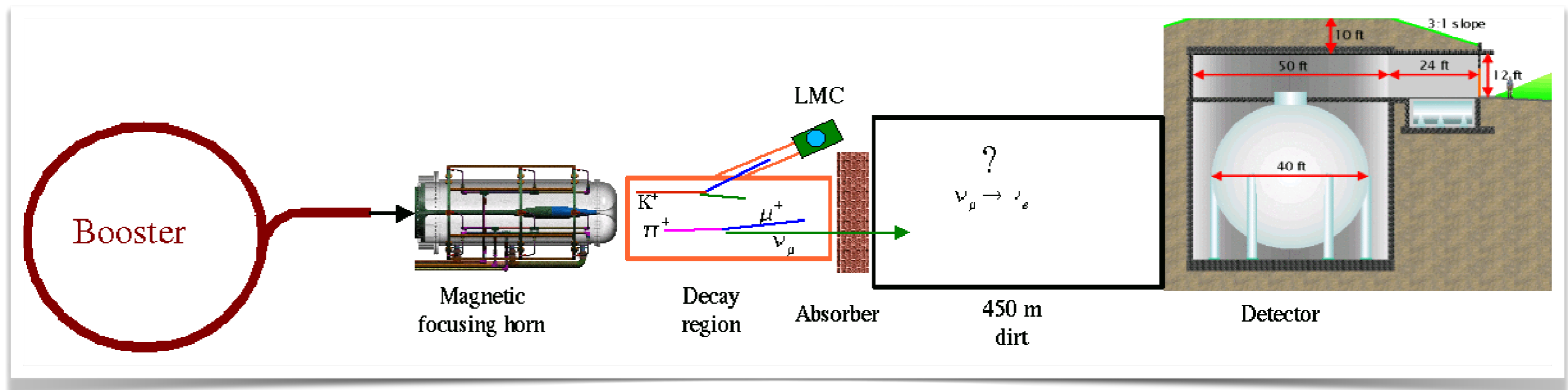


LSND  
unresolved



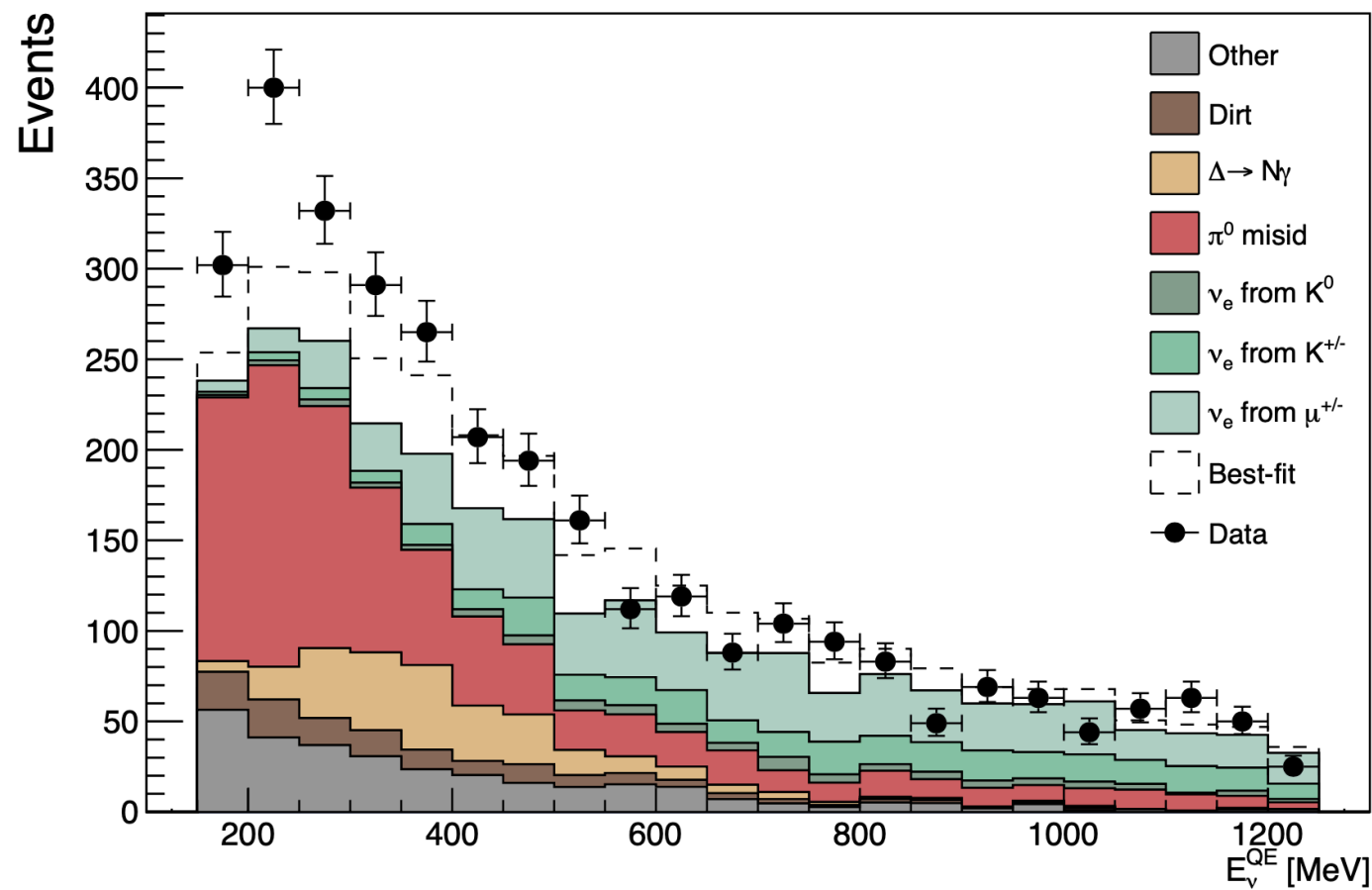
High BoONE

# Anomaly #5: MiniBooNE

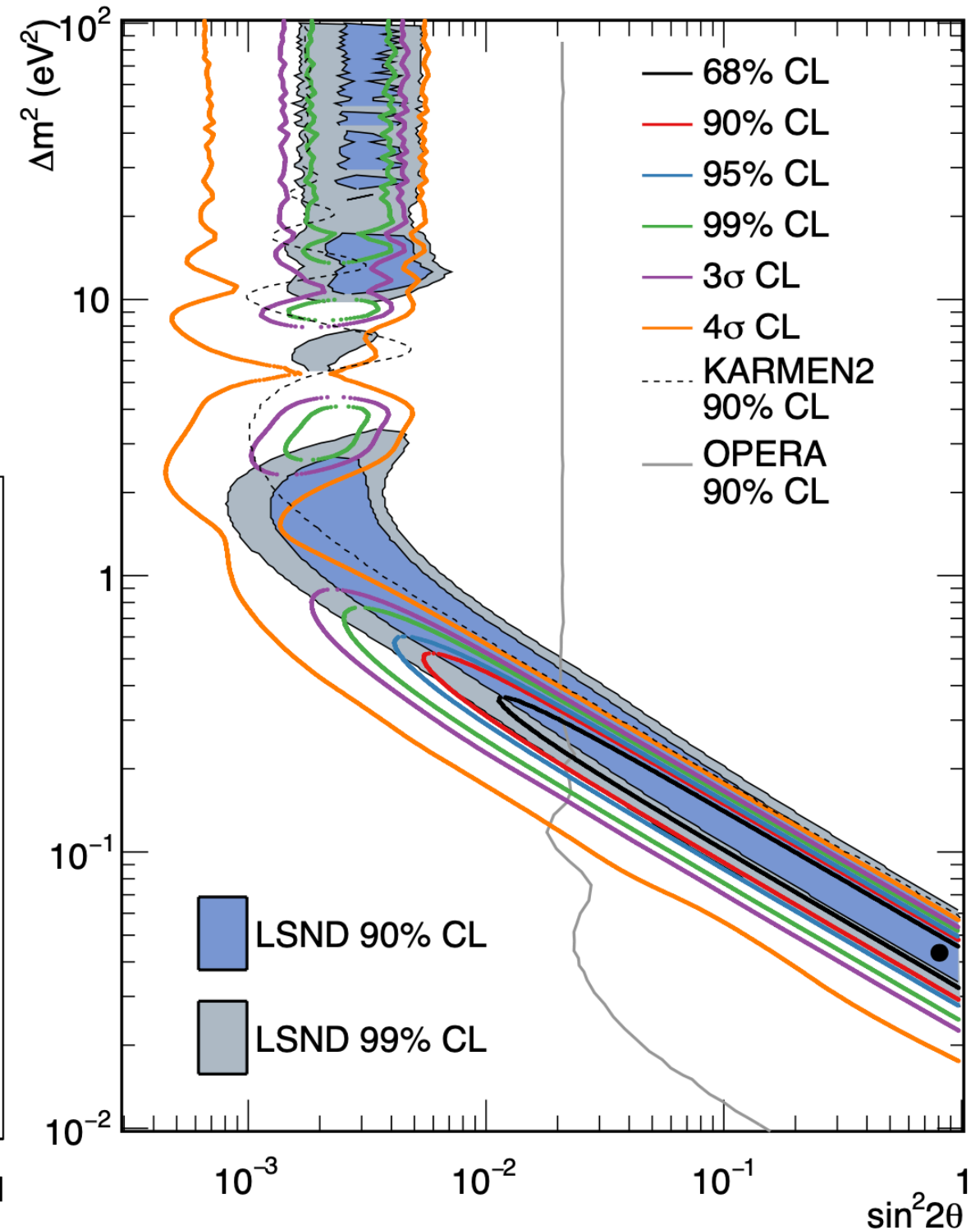


# Anomaly #5: MiniBooNE

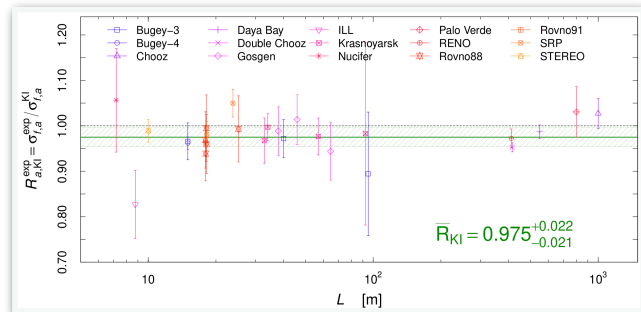
- ☑ Unexplained **low- $E$  excess** ( $4.8\sigma$ )
- ☑ Consistent with LSND
- ☑ **L/ $E$**  too small for std. oscillations (**wrong  $\Delta m^2$** )



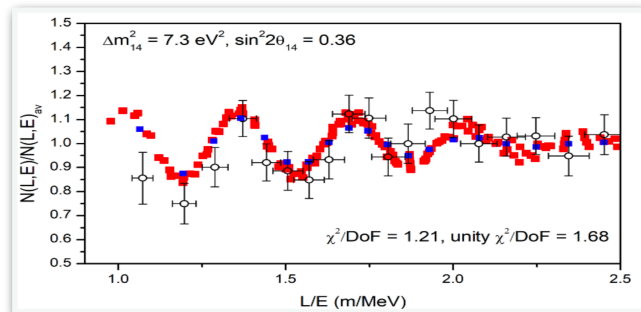
MiniBooNE Collaboration arXiv:2006.16883



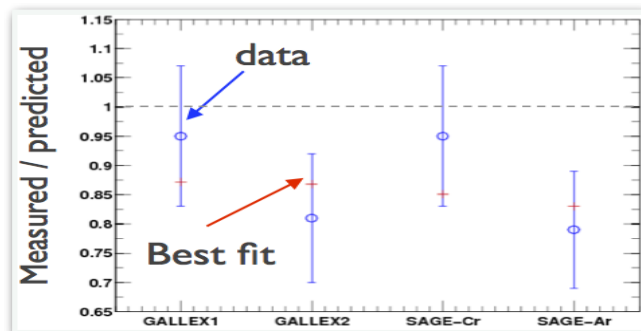
# Short-Baseline Anomalies



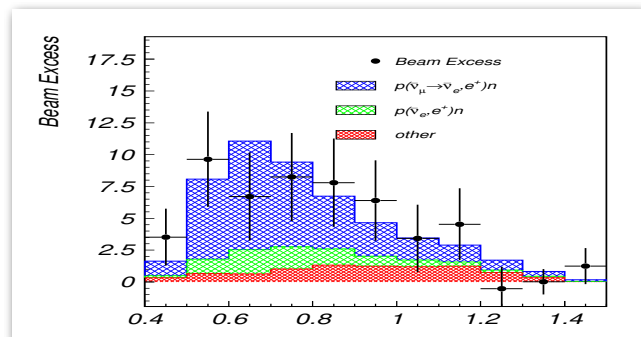
reactor flux anomaly  
resolved with new input data  
to flux calculation



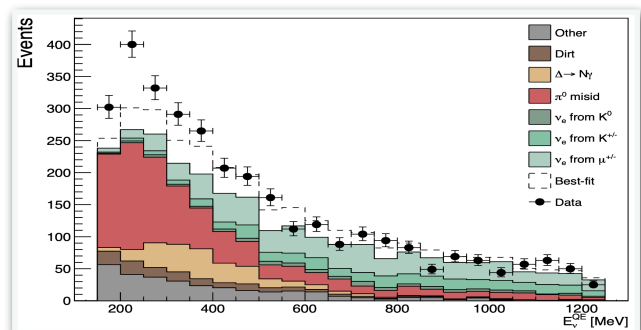
reactor spectra  
is there really an anomaly?



gallium anomaly  
unresolved, recently reinforced



LSND  
unresolved



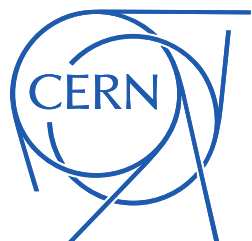
MiniBooNE  
unresolved



“With great precision comes great responsibility.”

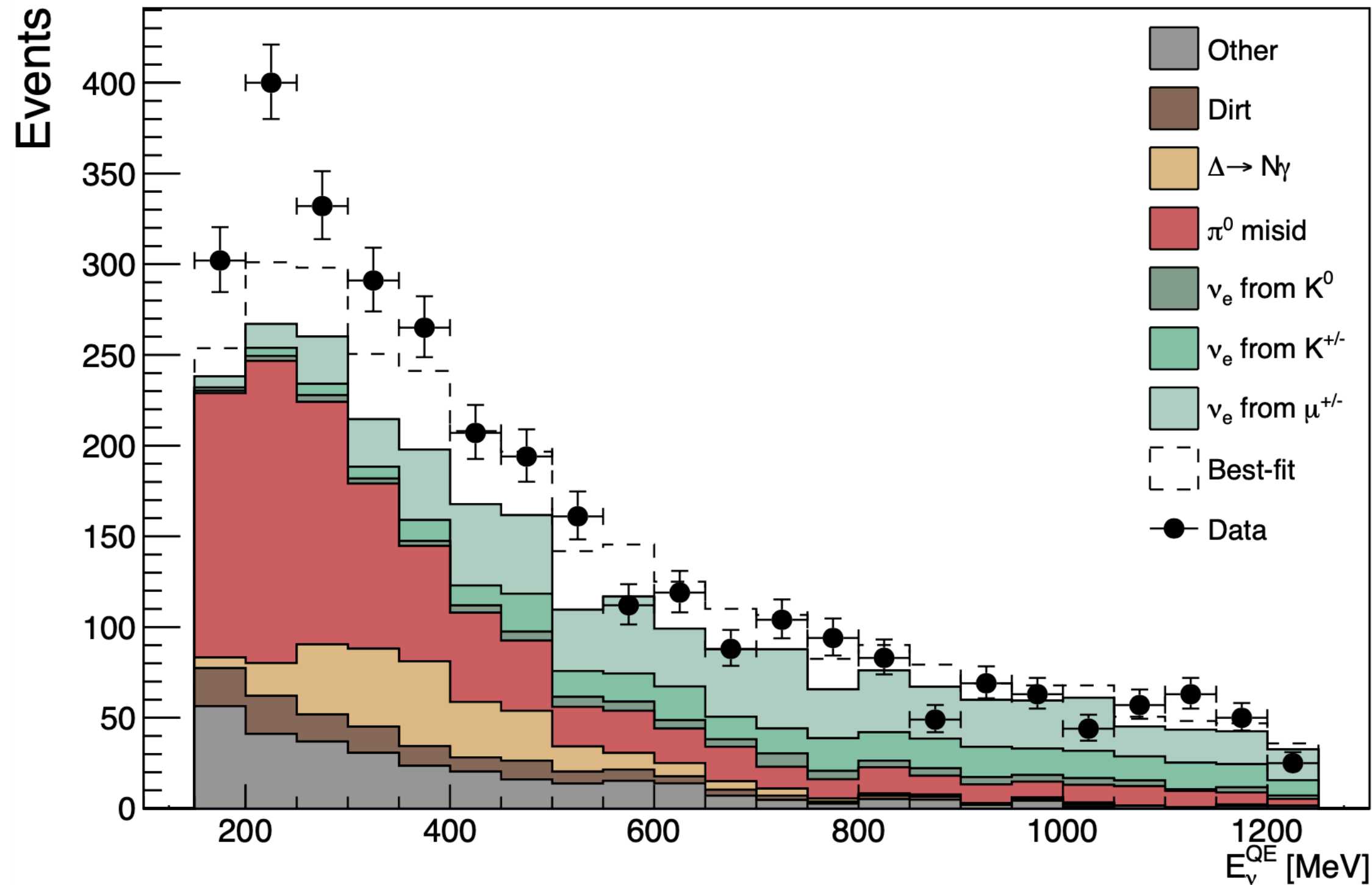
*Tim Linden, WIN 2021*

# Standard Model Explanations for the MiniBooNE Anomaly?



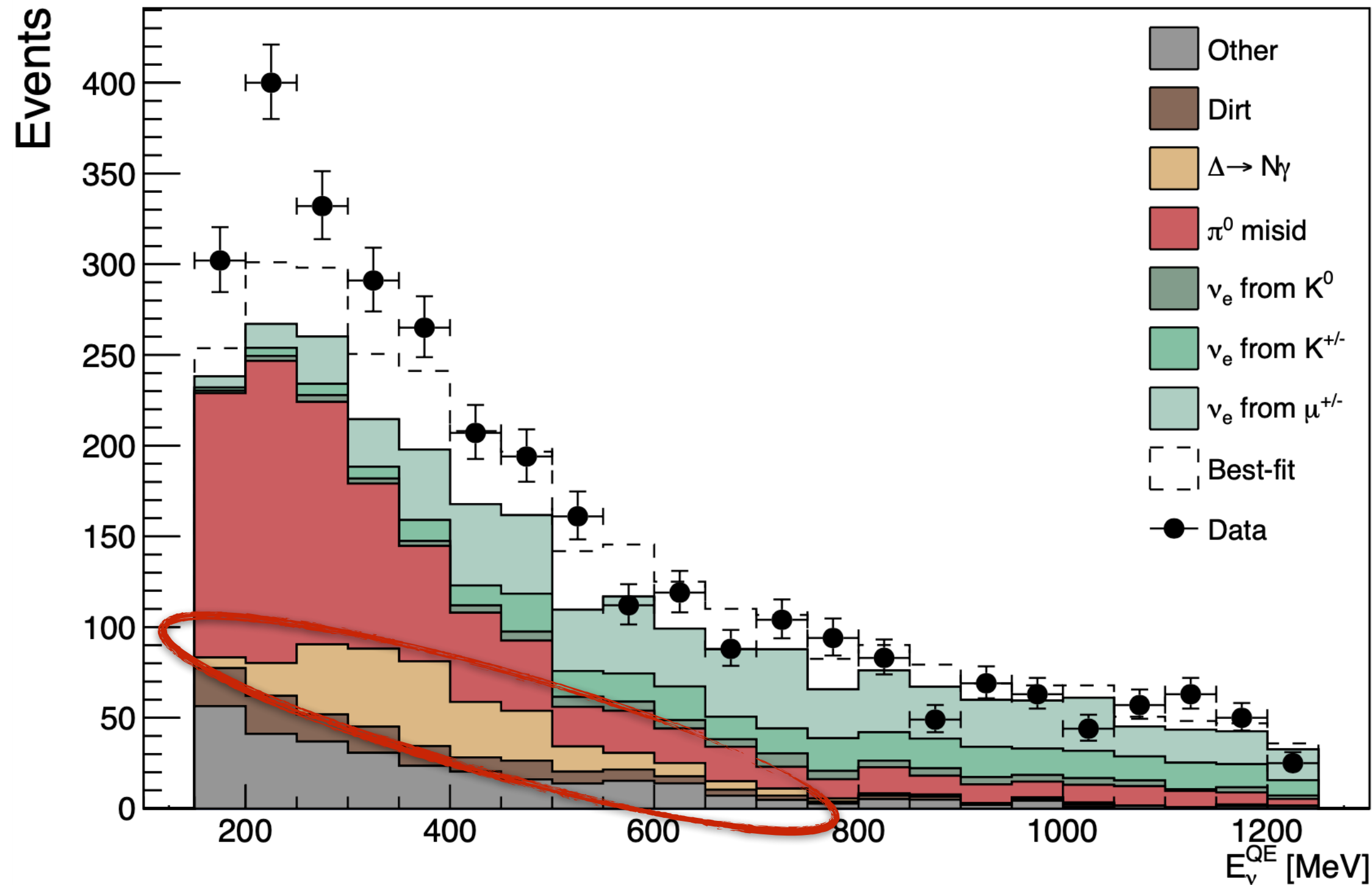


# MiniBooNE Backgrounds



MiniBooNE 2020

# MiniBooNE Backgrounds



MiniBooNE 2020

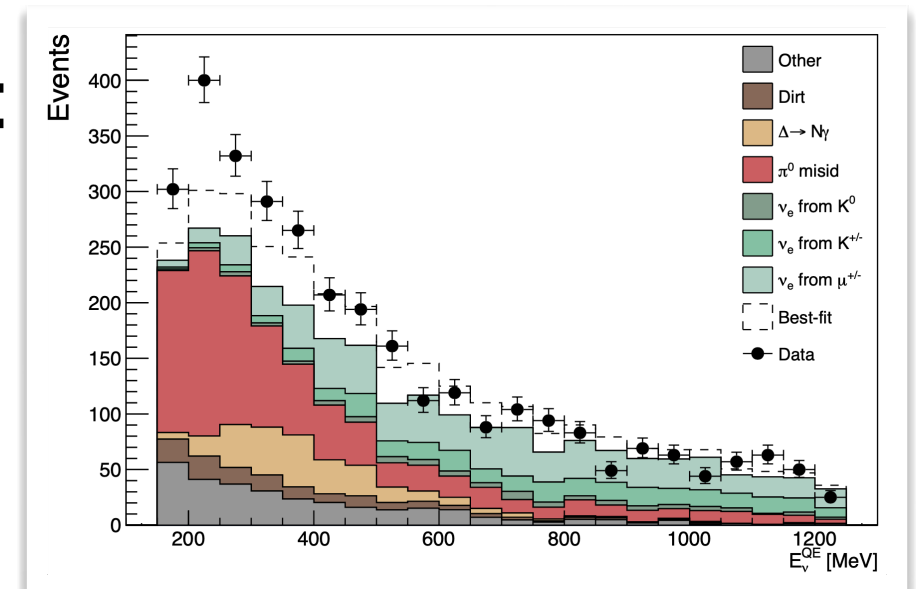
$$\Delta \rightarrow \gamma N$$

☑ Neutral current neutrino interaction:  
 $\nu + N \rightarrow \nu + \Delta(1232)$

☑  $\Delta(1232)$  mostly decays to  $\pi + N$

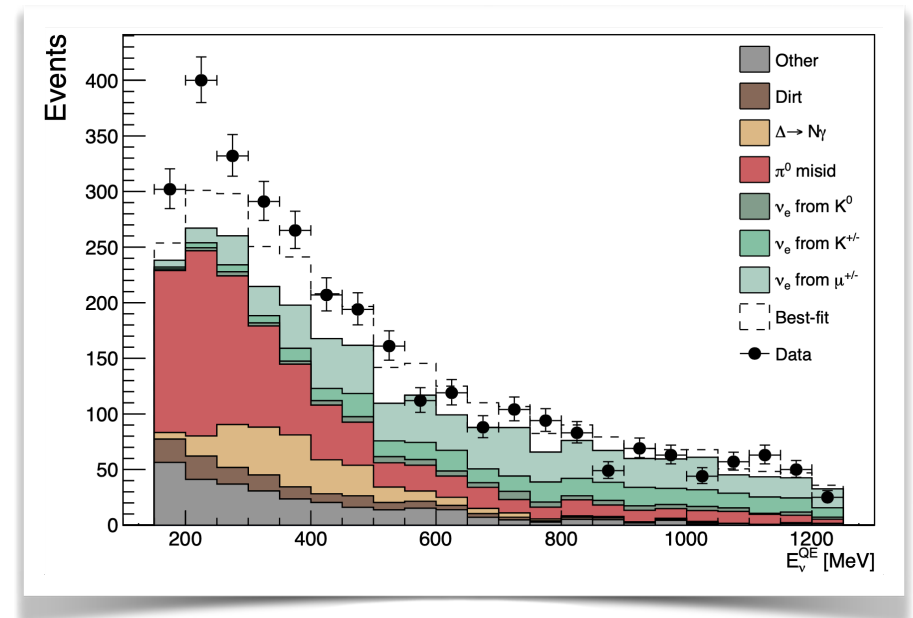
☑ But a rare decay exists to  $\gamma + N$

☑ MiniBooNE cannot distinguish  
 $\gamma$  from  $e^-$



$$\Delta \rightarrow \gamma N$$

- ☑  $\Delta$  production rate can be estimated from  $\Delta \rightarrow \pi N$
- ☑ Pions may be absorbed
  - may excite another  $\Delta$  resonance
    - ➡  $\Delta \rightarrow \gamma N$  enhanced by  $\sim$ factor 2
  - or may be absorbed
    - ➡ control region suppressed by  $\sim$ factor 2



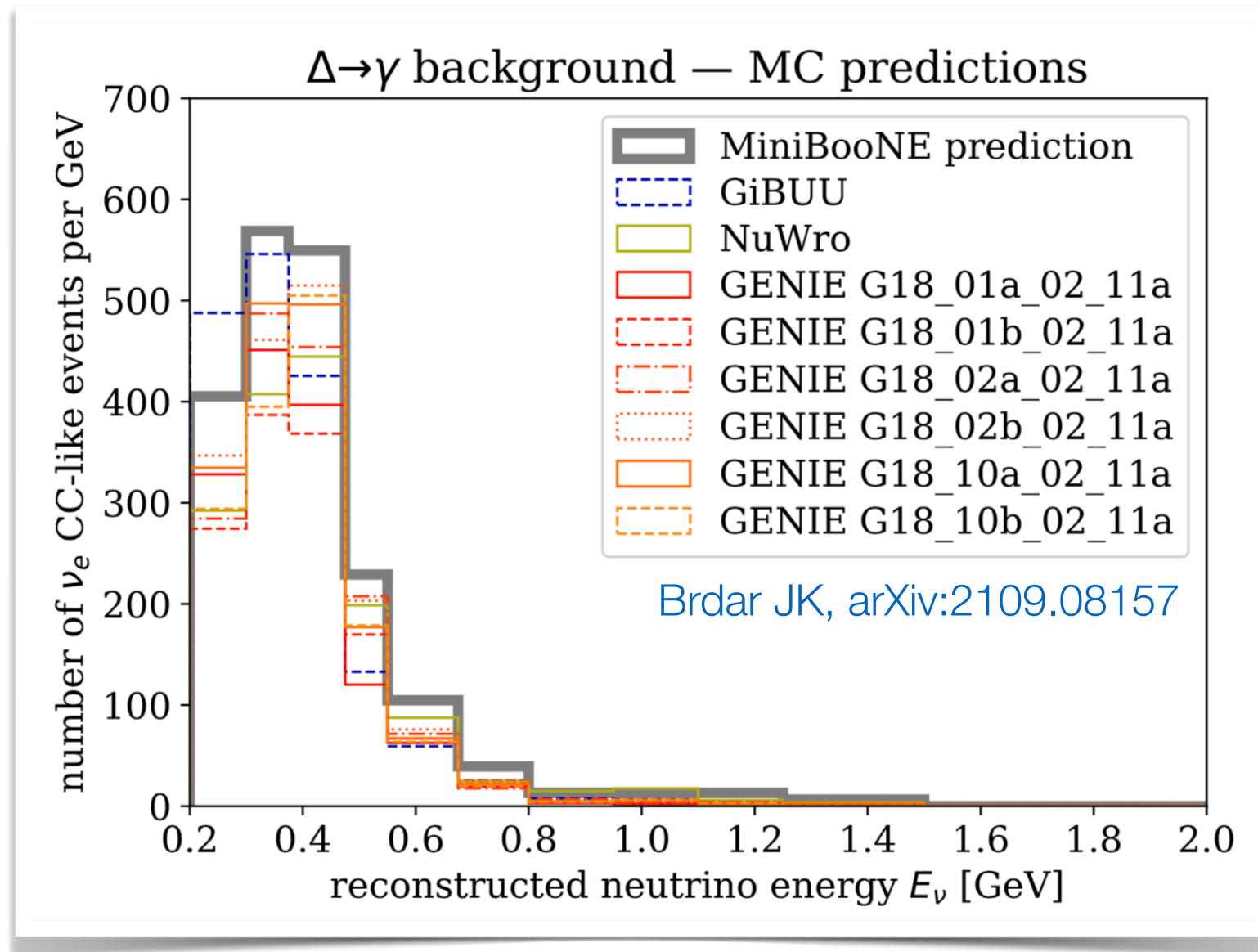
Ioannisian [1909.08571](#)

Giunti Ioannisian Ranucci [1912.01524](#)

- ☑ This factor 2 **has been taken into account** by MiniBooNE

MiniBooNe, [arXiv:2006.16883](#)

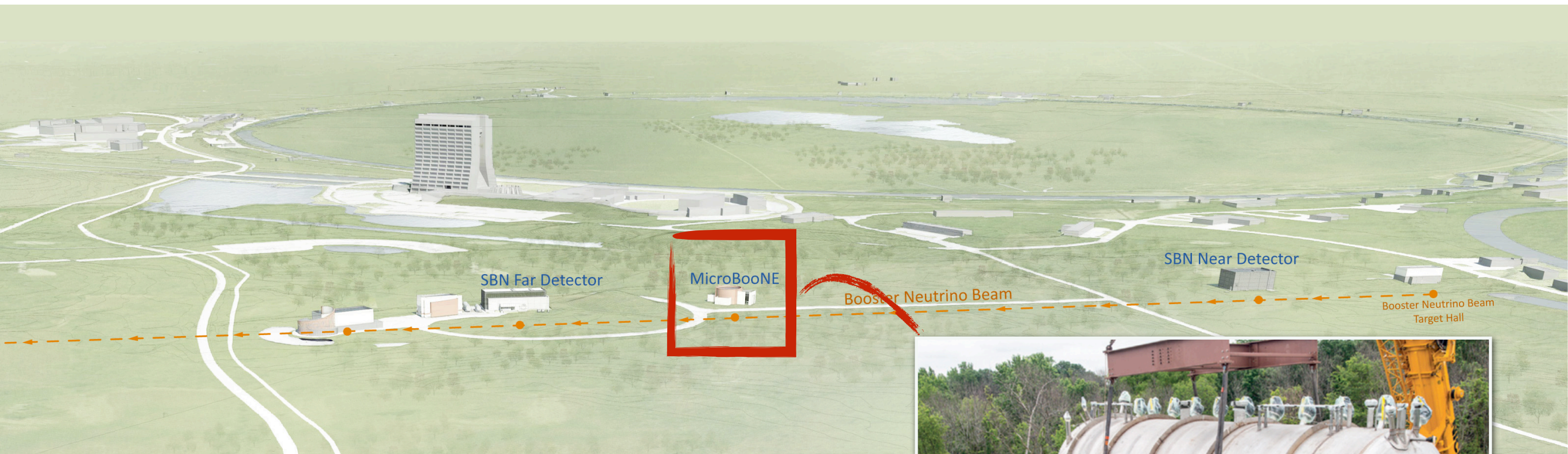
# $\Delta \rightarrow \gamma N$ : Comparison of Generators



- ✓ histograms calibrated to NUANCE  
(the generator used by MiniBooNE)
- ✓ using our own implementation of radiative resonance decays  
in GiBUU, NuWro, NUANCE



# MicroBooNE

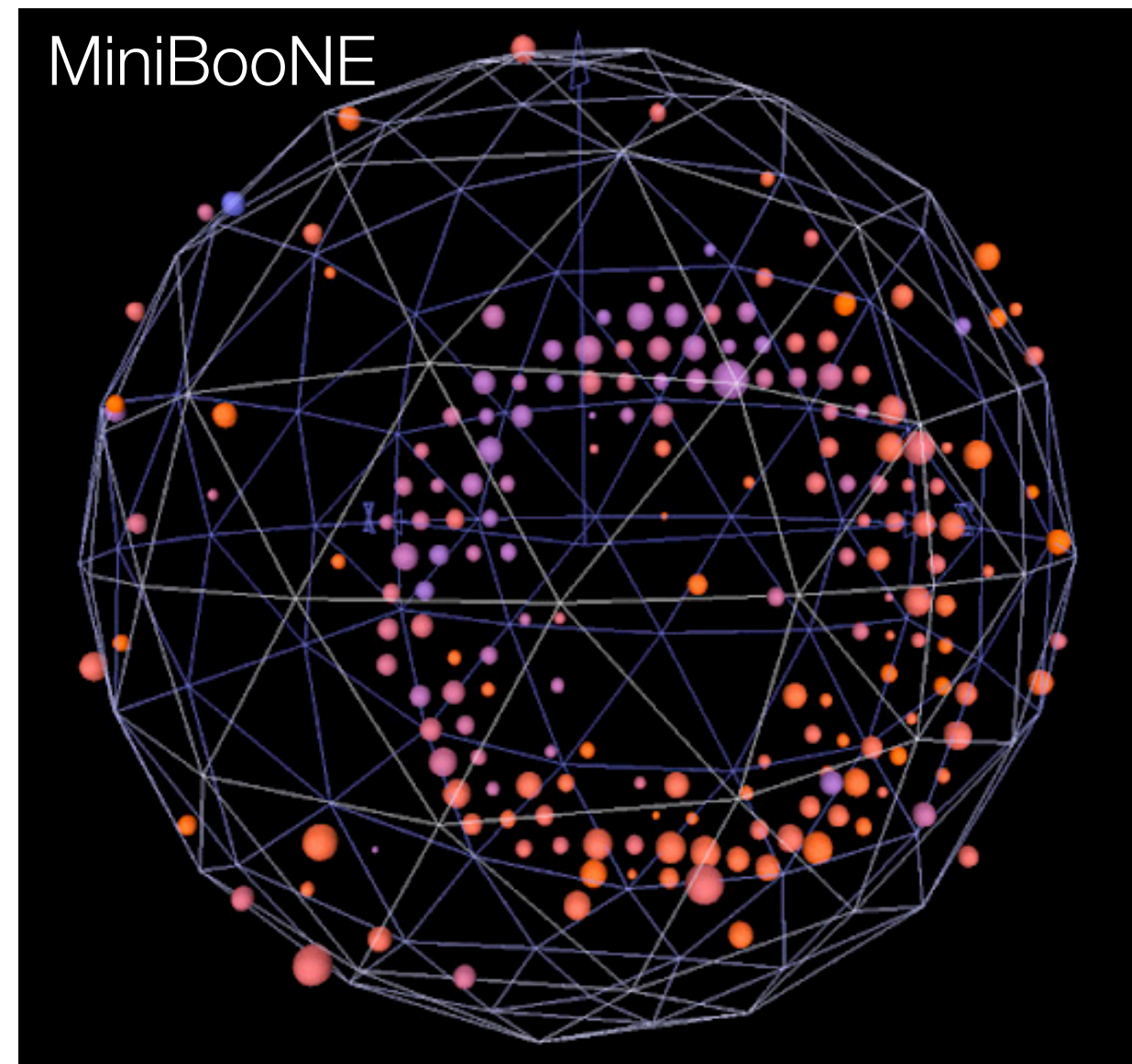
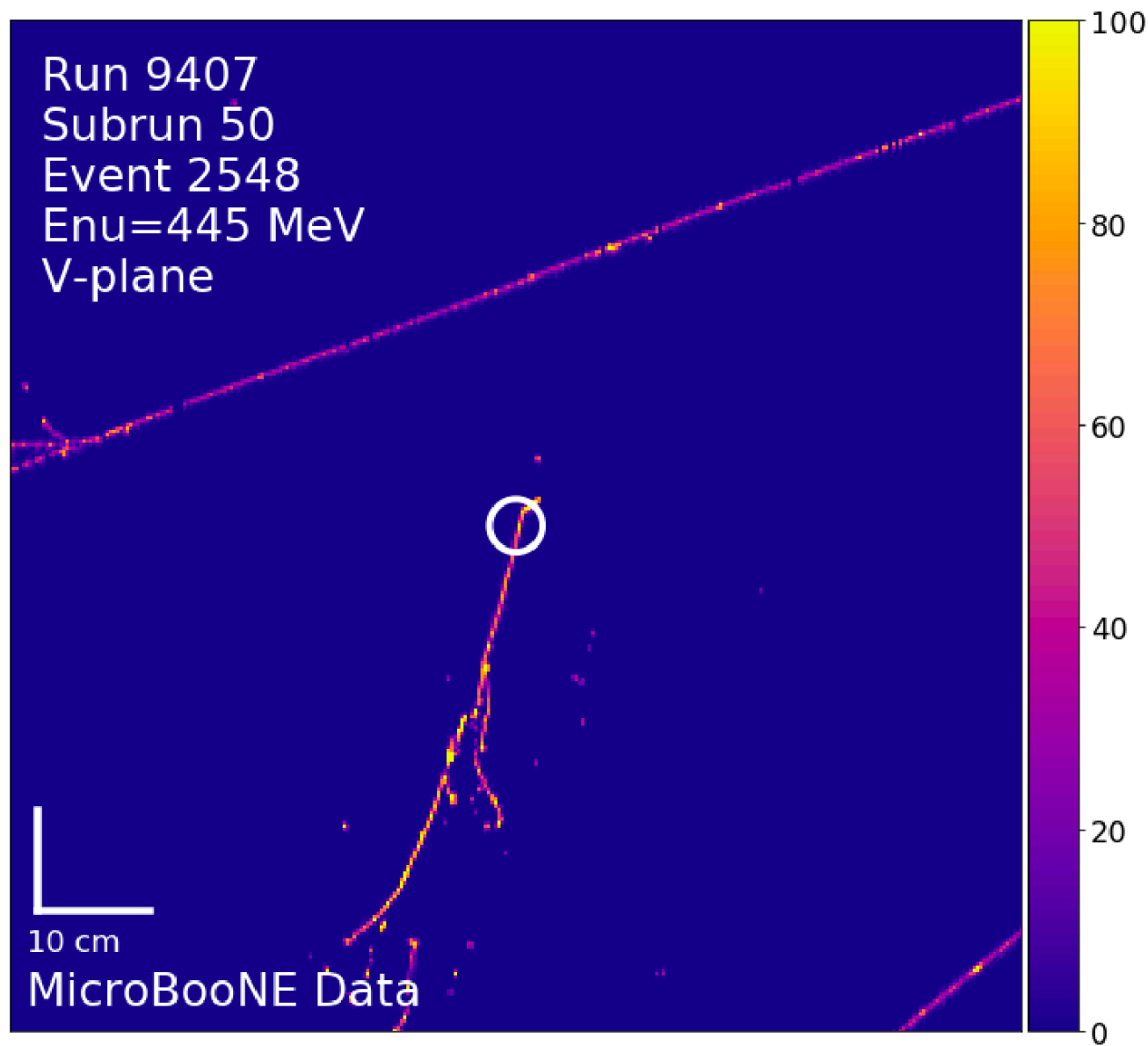


- ✓ 80 ton LAr TPC
- ✓ Very good event reconstruction capabilities
- can distinguish  $e^\pm$  from  $\gamma$

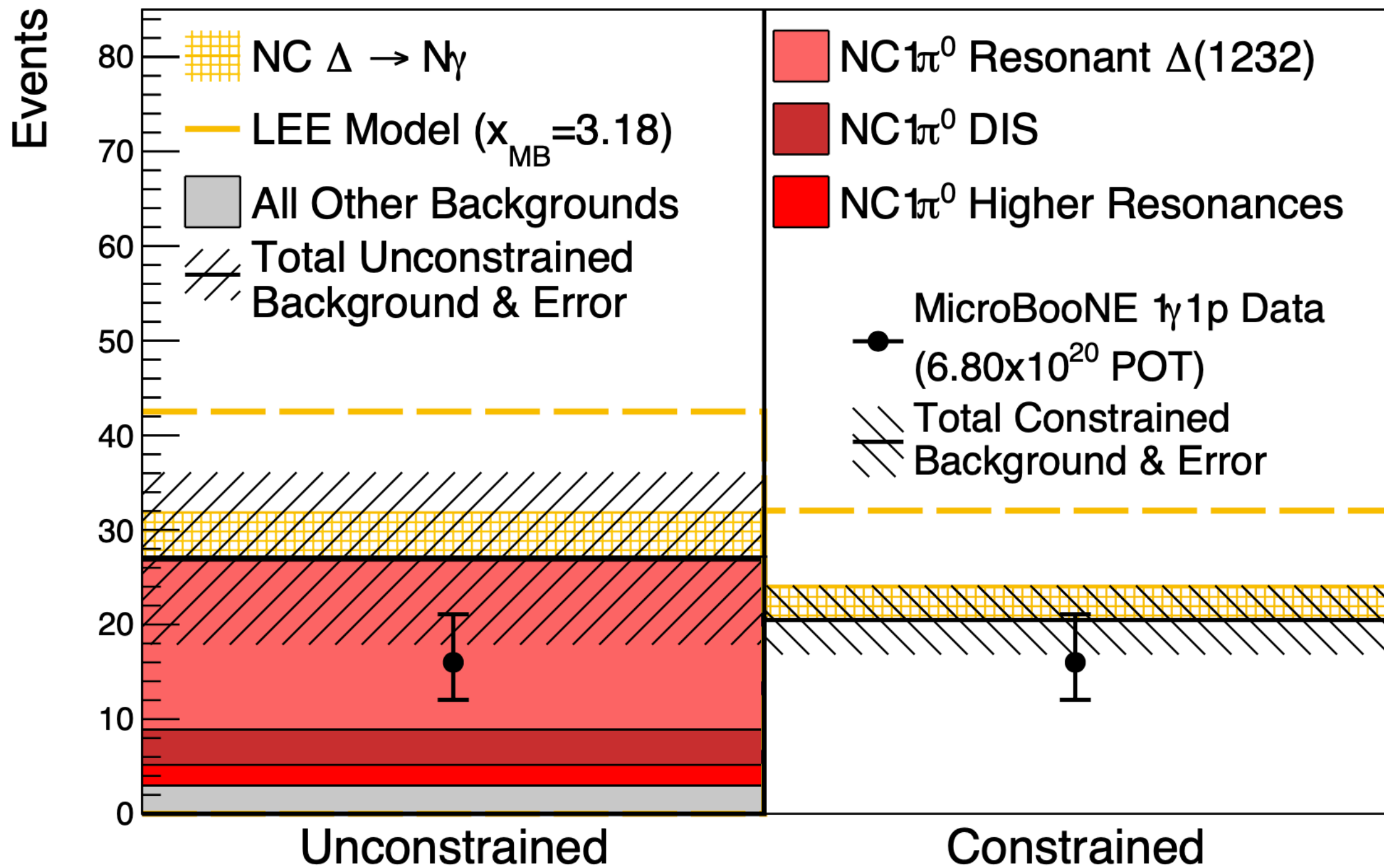




# MiniBooNE vs. MicroBooNE



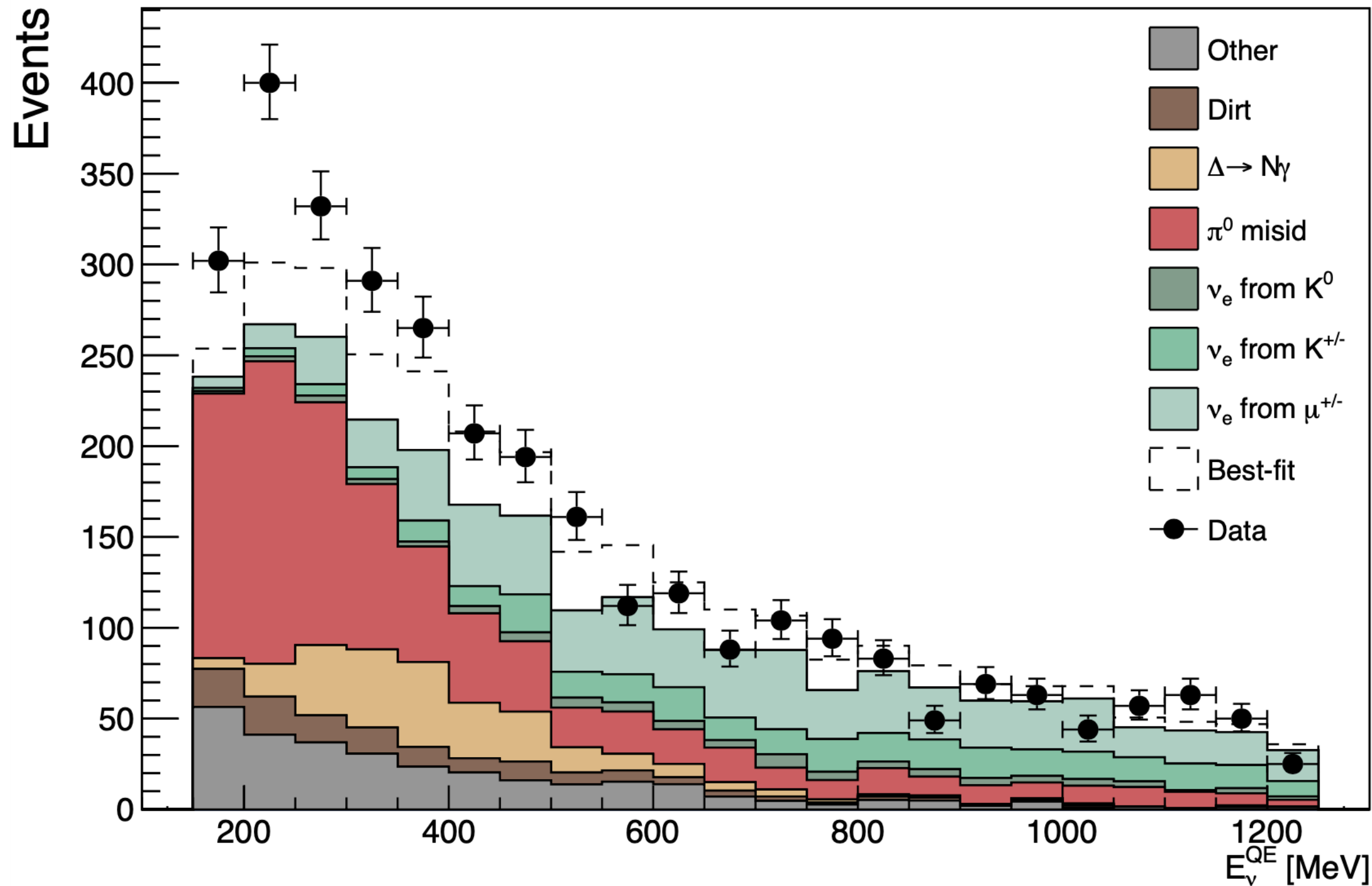
# $\Delta \rightarrow \gamma N$ : MicroBooNE



MicroBooNE arXiv:2110.00409

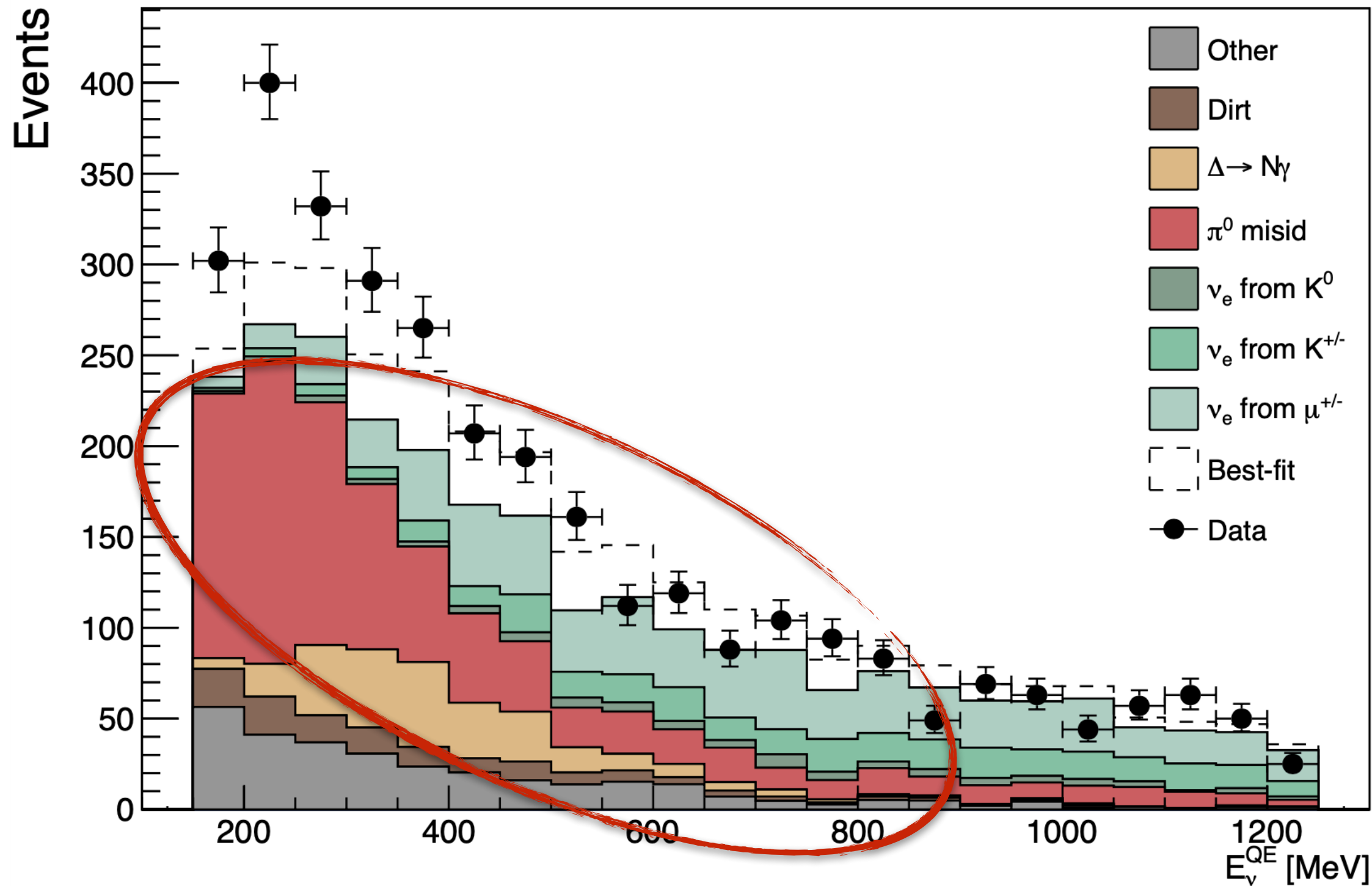


# NC $\pi^0$ Production in MiniBooNE



MiniBooNE 2020

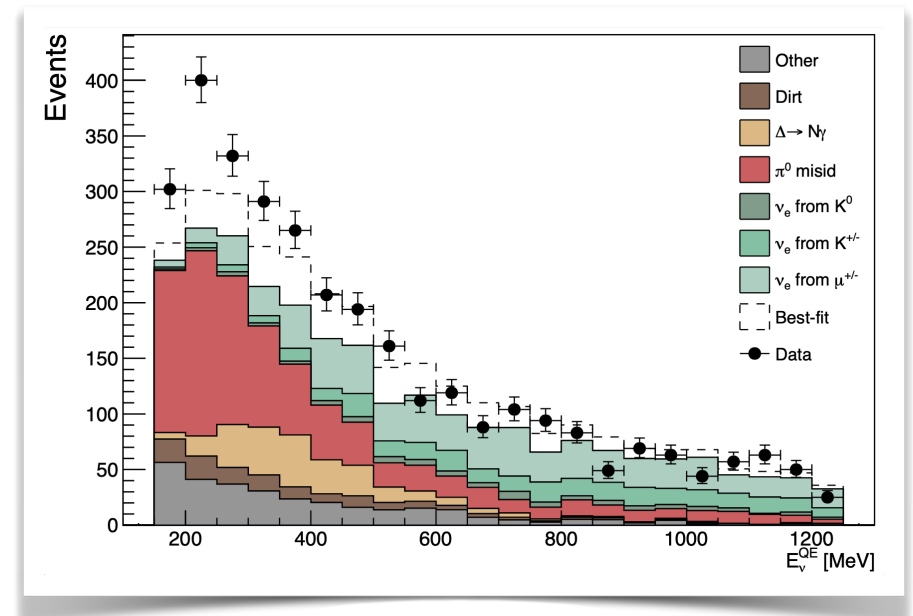
# NC $\pi^0$ Production in MiniBooNE



MiniBooNE 2020

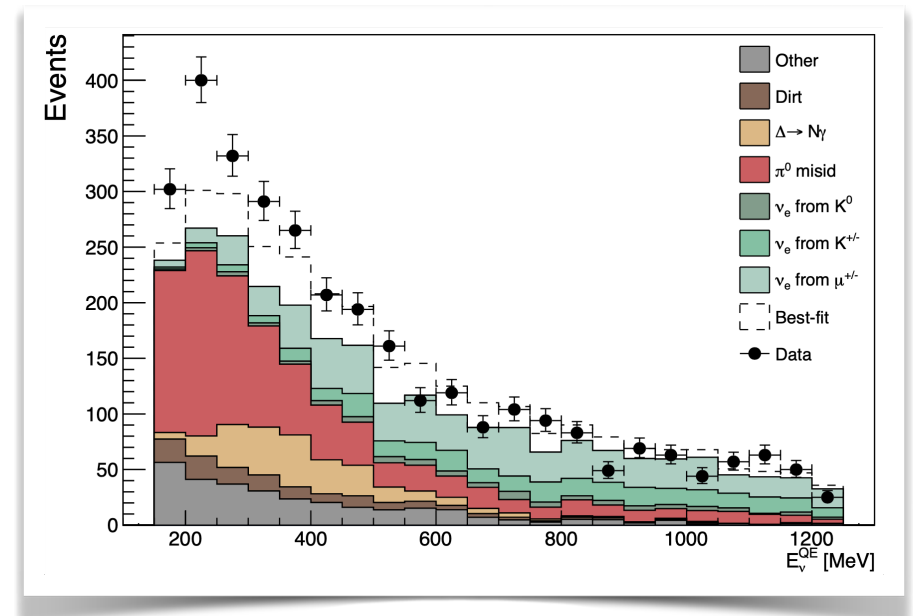
# NC $\pi^0$ Production in MiniBooNE

- ☑  $\nu + N \rightarrow \nu + N + \pi^0$
- ☑  $\pi^0 \rightarrow \gamma\gamma$  is observable  
→  $\pi^0$  production rate *measured*
- ☑ Rarely, one photon from  $\pi^0 \rightarrow \gamma\gamma$  is missed
  - conversion outside active volume
    - ▢ disavored by uniform distribution of excess events
  - decay may be asymmetric
  - boosted/collimated photons
    - ▢ not at low energies, where the excess is
  - photon absorbed on nucleus before converting



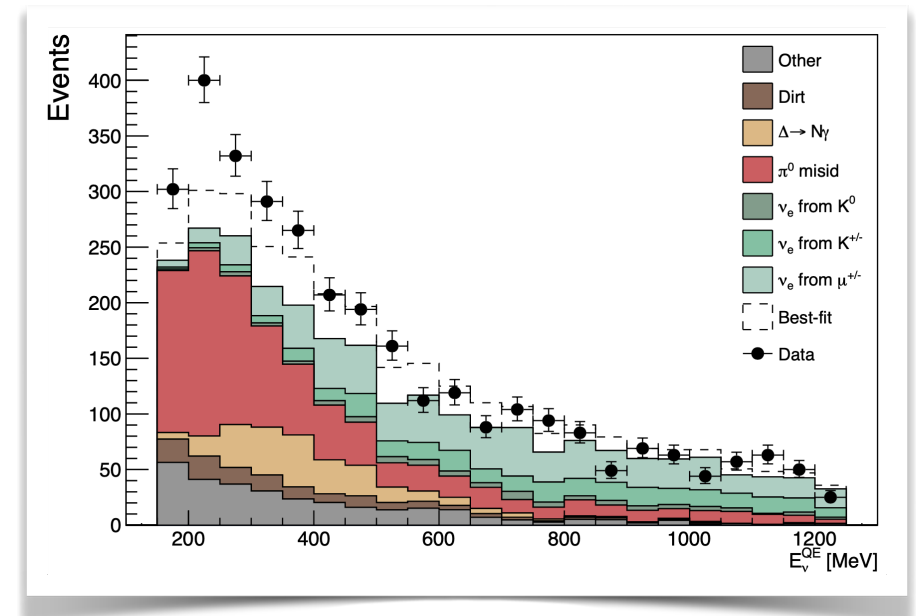
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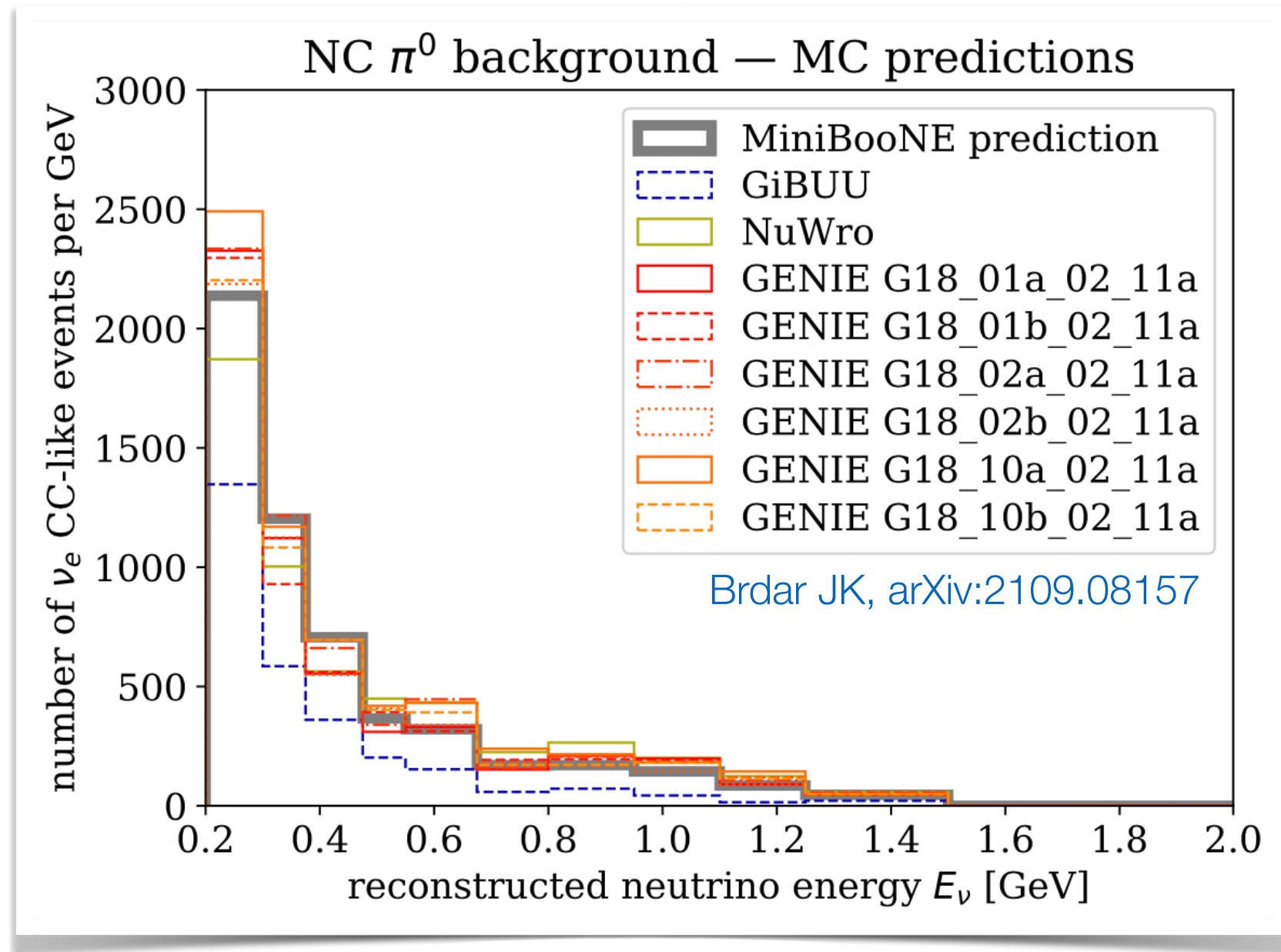


# NC $\pi^0$ Production in MiniBooNE

- ☑ Problem with translation from control sample to signal region?
  - unlikely after >10 years
  - but probability for missing one  $\gamma$  is determined based on MC only
- ☑ One possible cross-check:  
CC  $\nu_\mu + \pi^0$  production
  - compare rate of  $\mu + 1$  shower and  $\mu + 2$  shower events



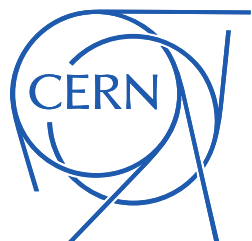
# $\pi^0$ in MiniBooNE: Comparison of Generators



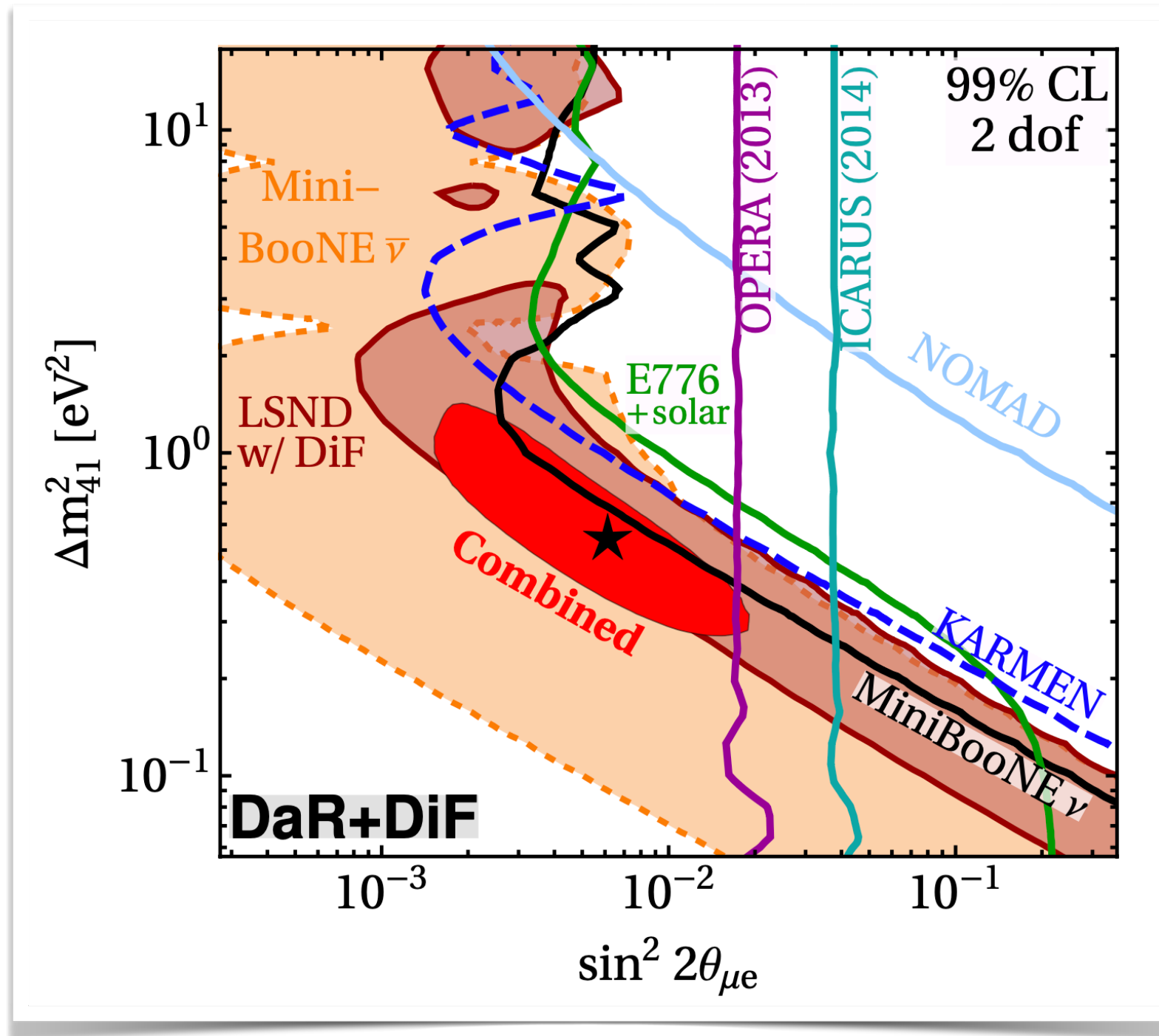
☑ histograms calibrated to NUANCE  
(the generator used by MiniBooNE)

# Sterile Neutrinos after all?

## Global Fits



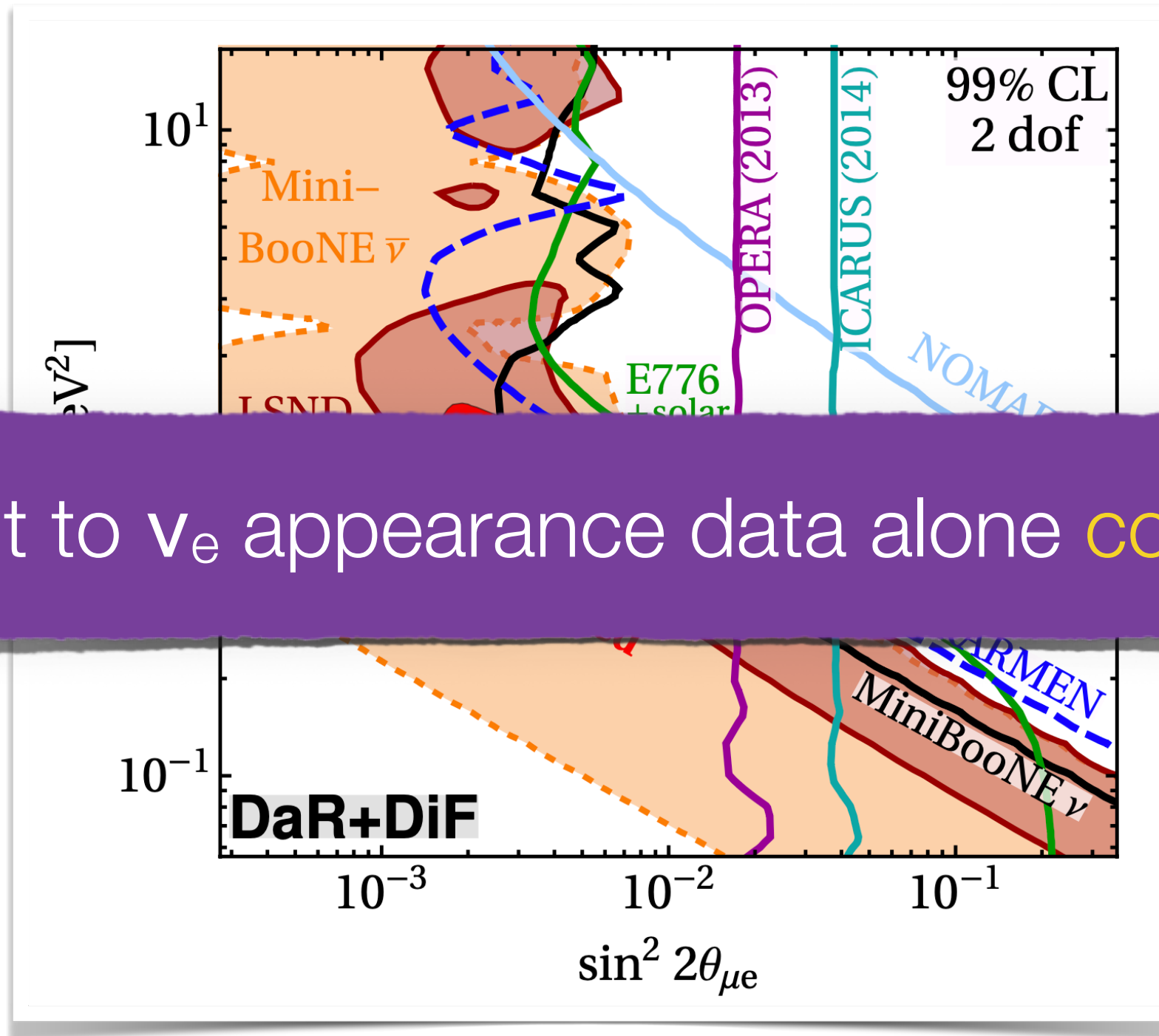
# Global Fit to $\nu_e$ Appearance Data



Dentler *et al.*, [1803.10661](https://arxiv.org/abs/1803.10661)

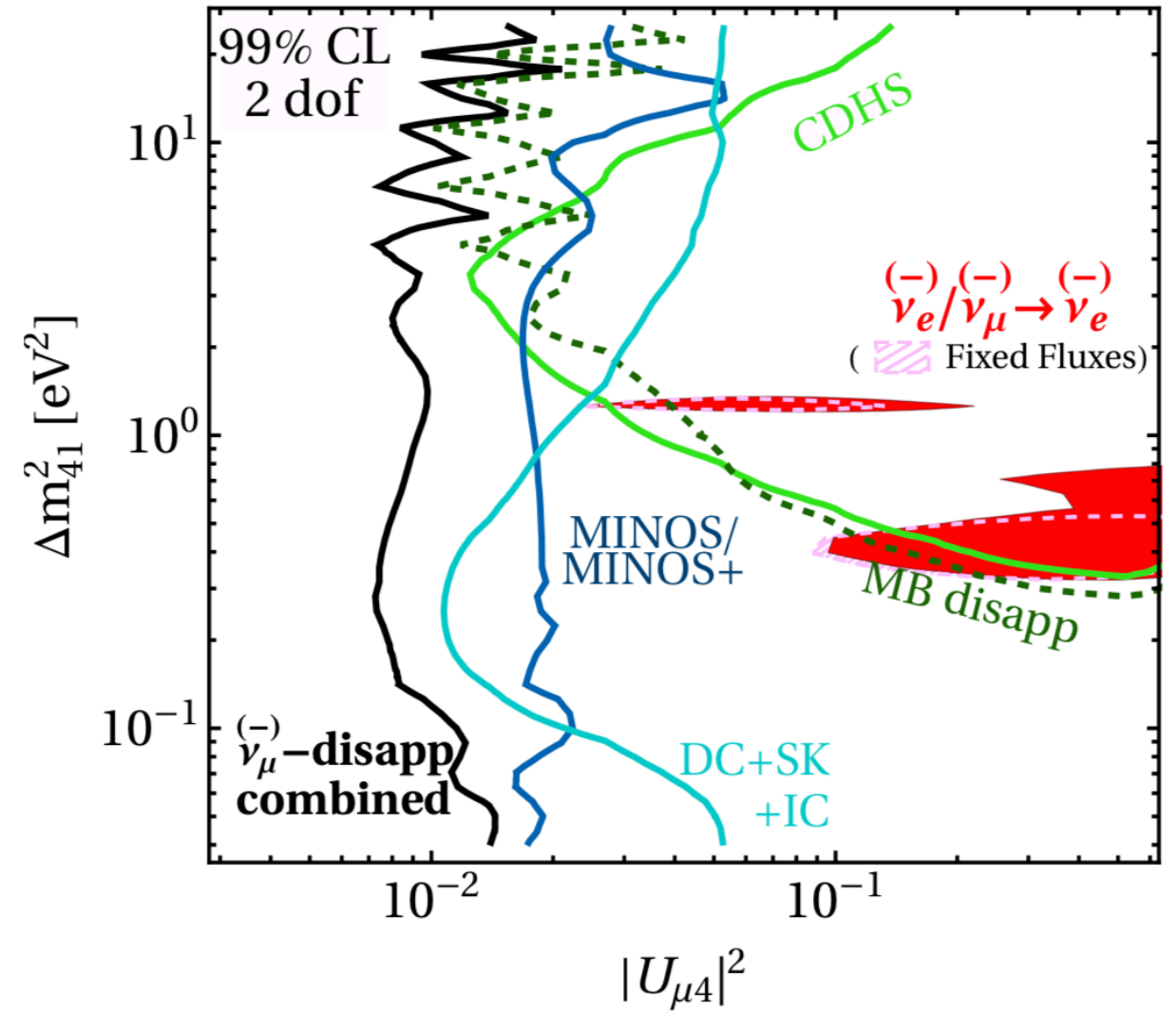
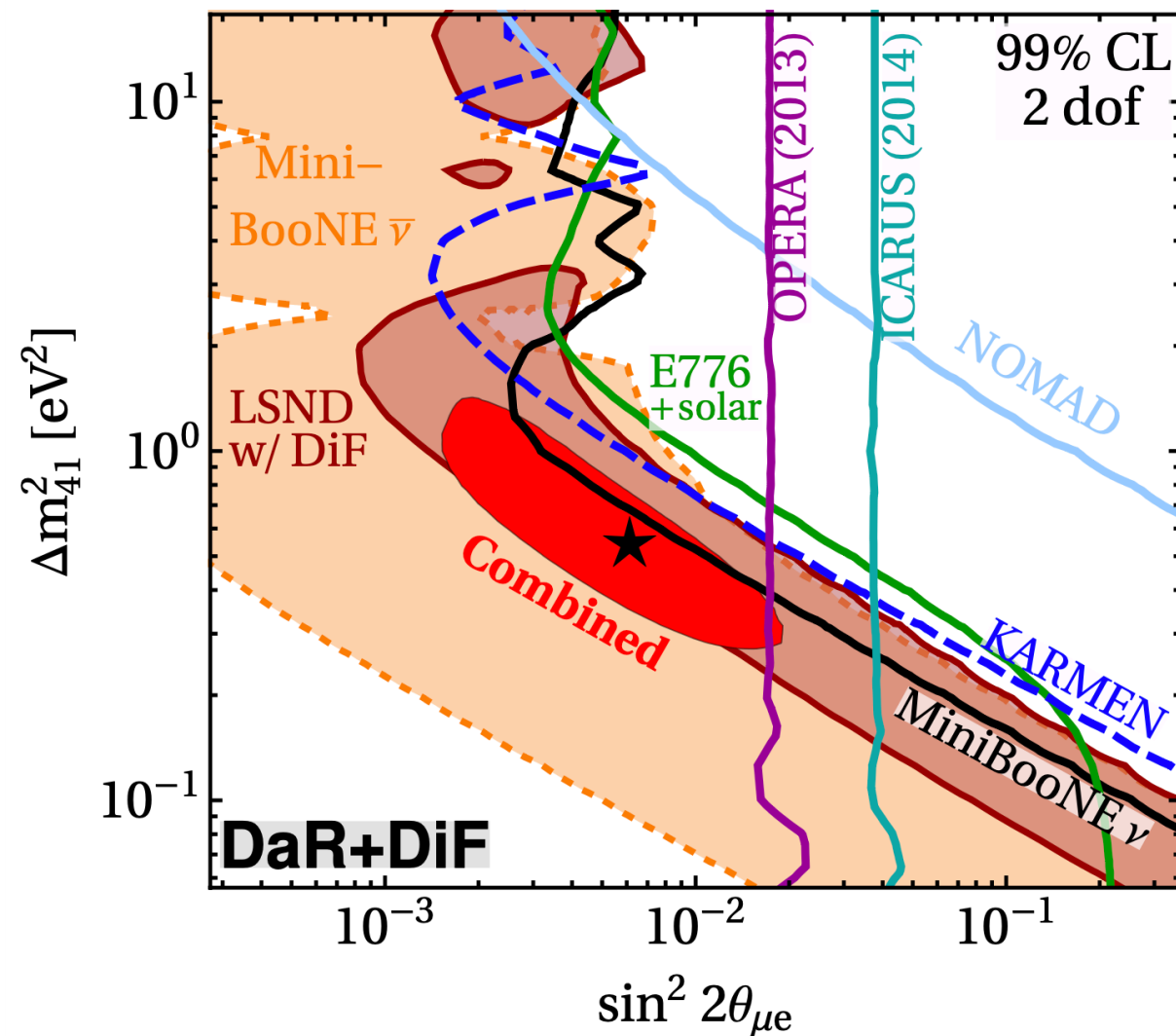


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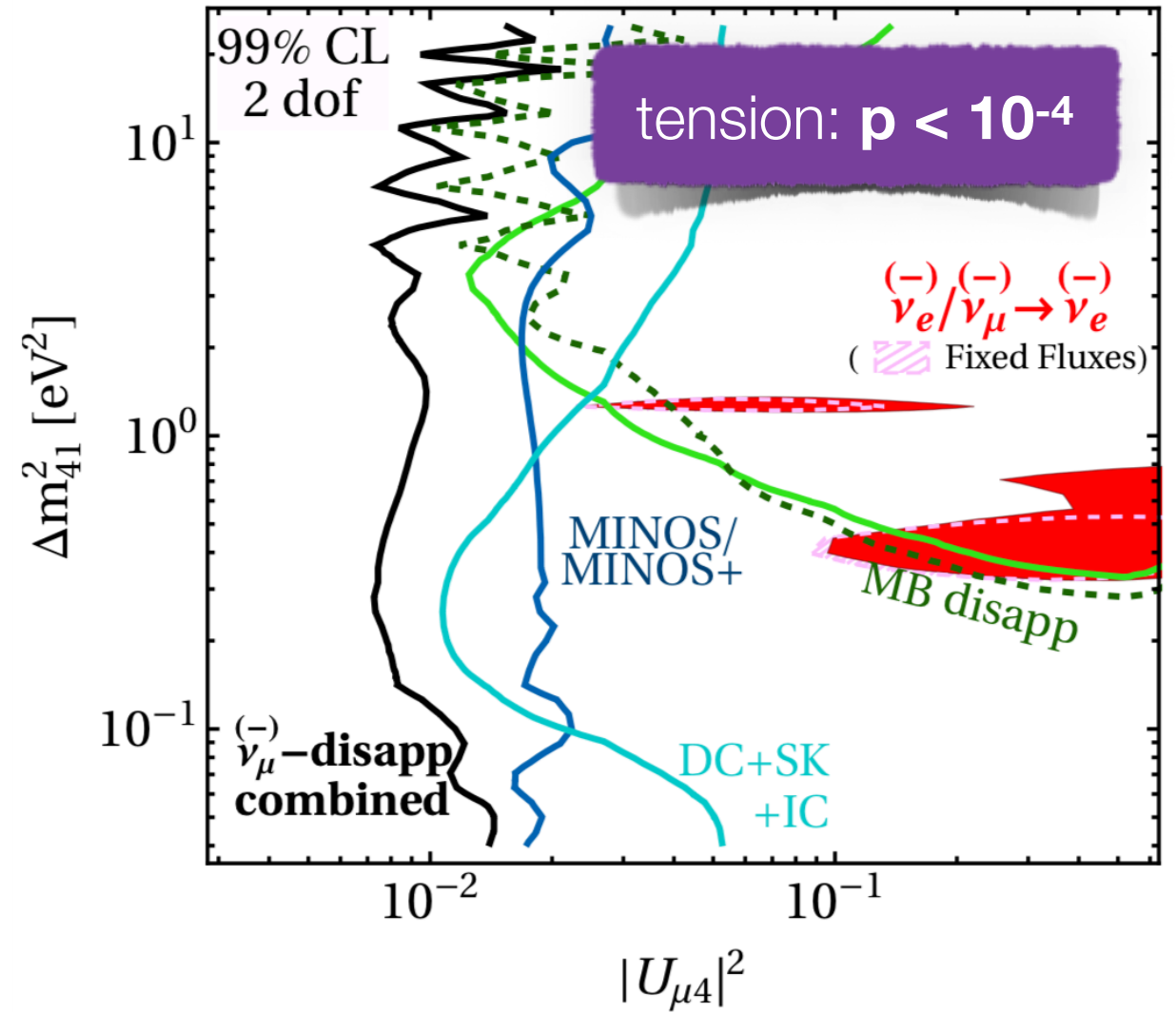
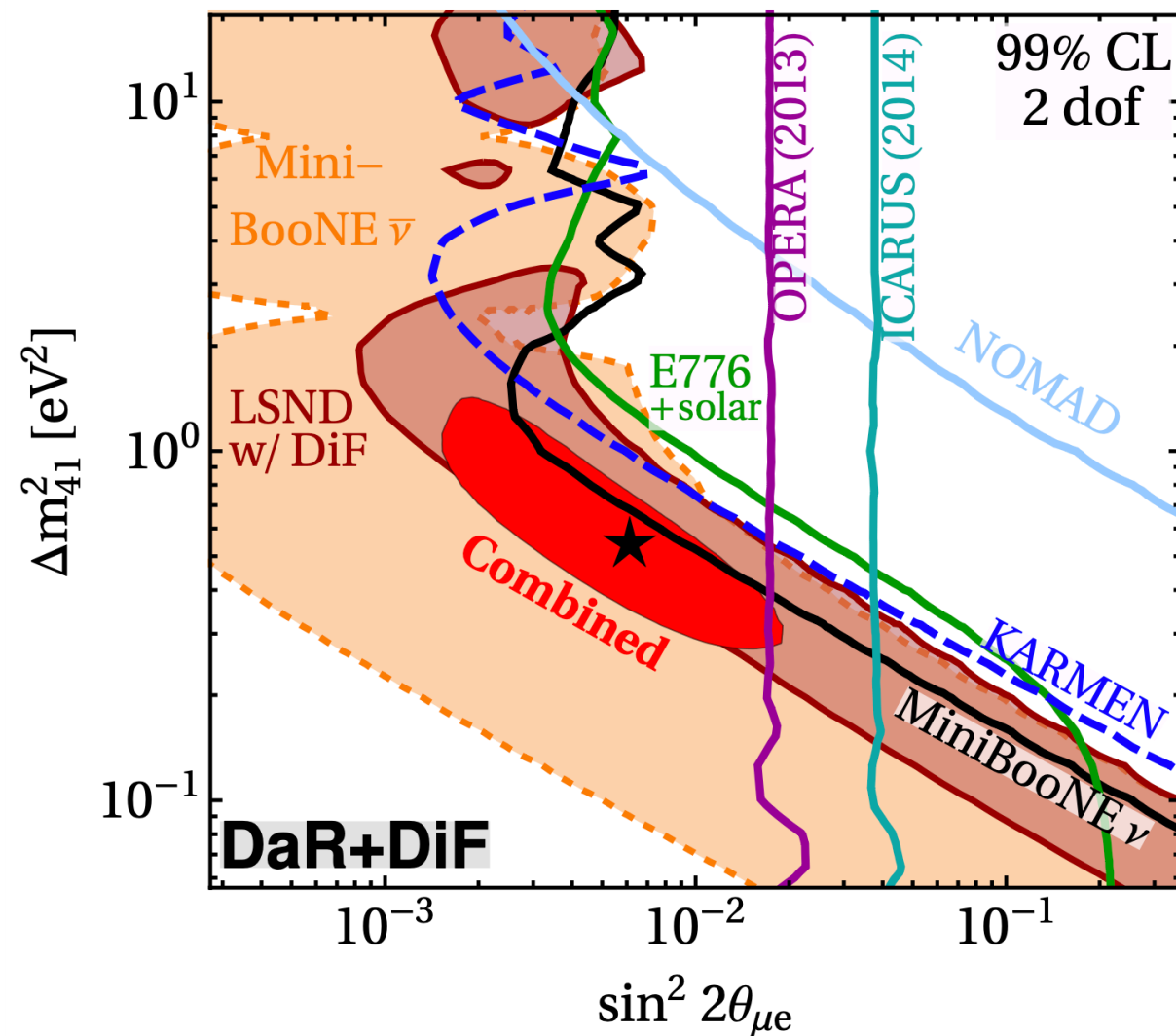
Dentler *et al.*, [1803.10661](#)

# Tension in Global Sterile Neutrino Fits



Dentler Hernandez JK Machado Maltoni Martinez Schwetz, [1803.10661](#)  
 see also works by Argüelles Collin Conrad Hardin Shaevitz; Gariazzo Giunti Laveder Li

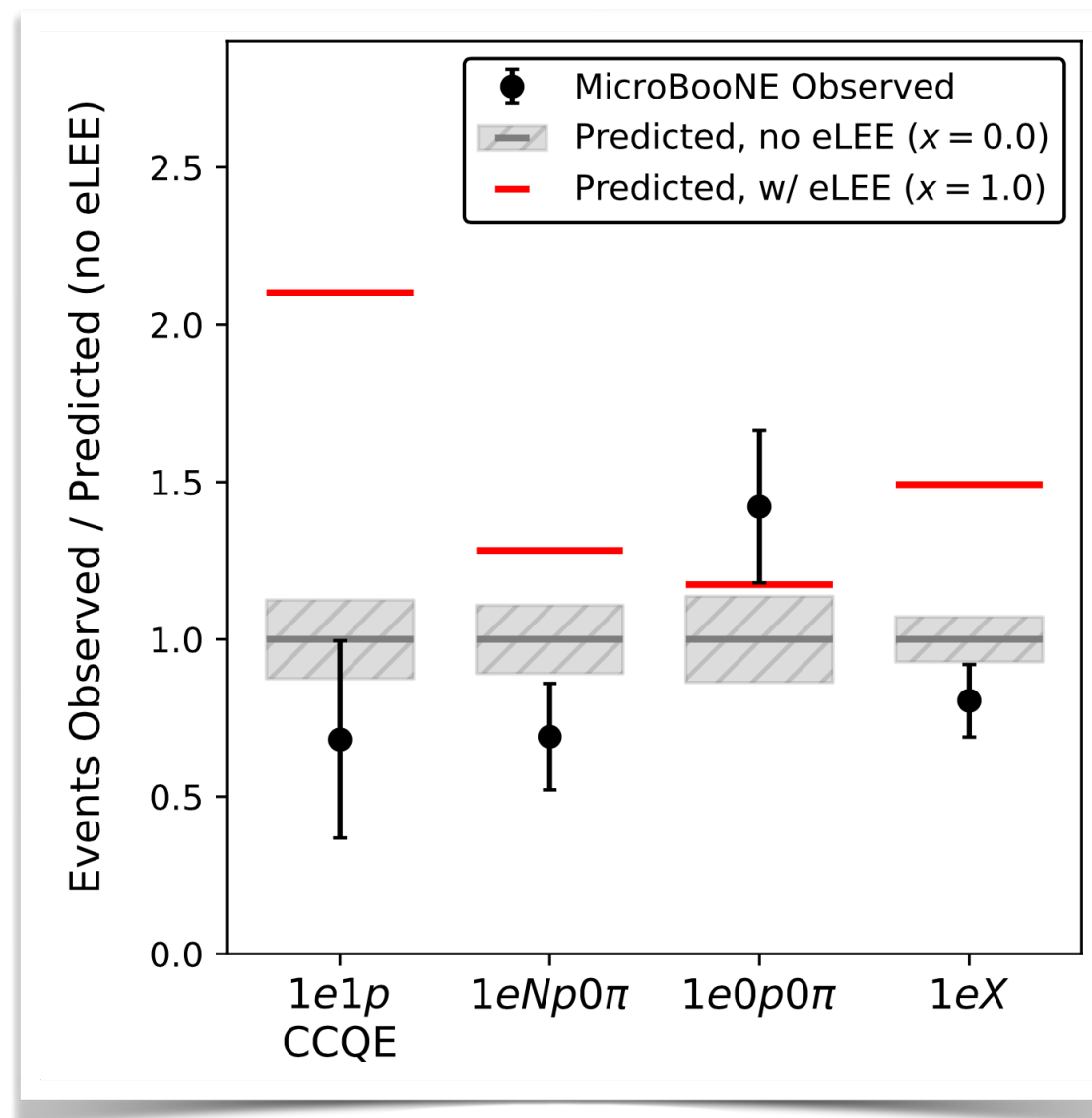
# Tension in Global Sterile Neutrino Fits



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# MicroBooNE weighs in

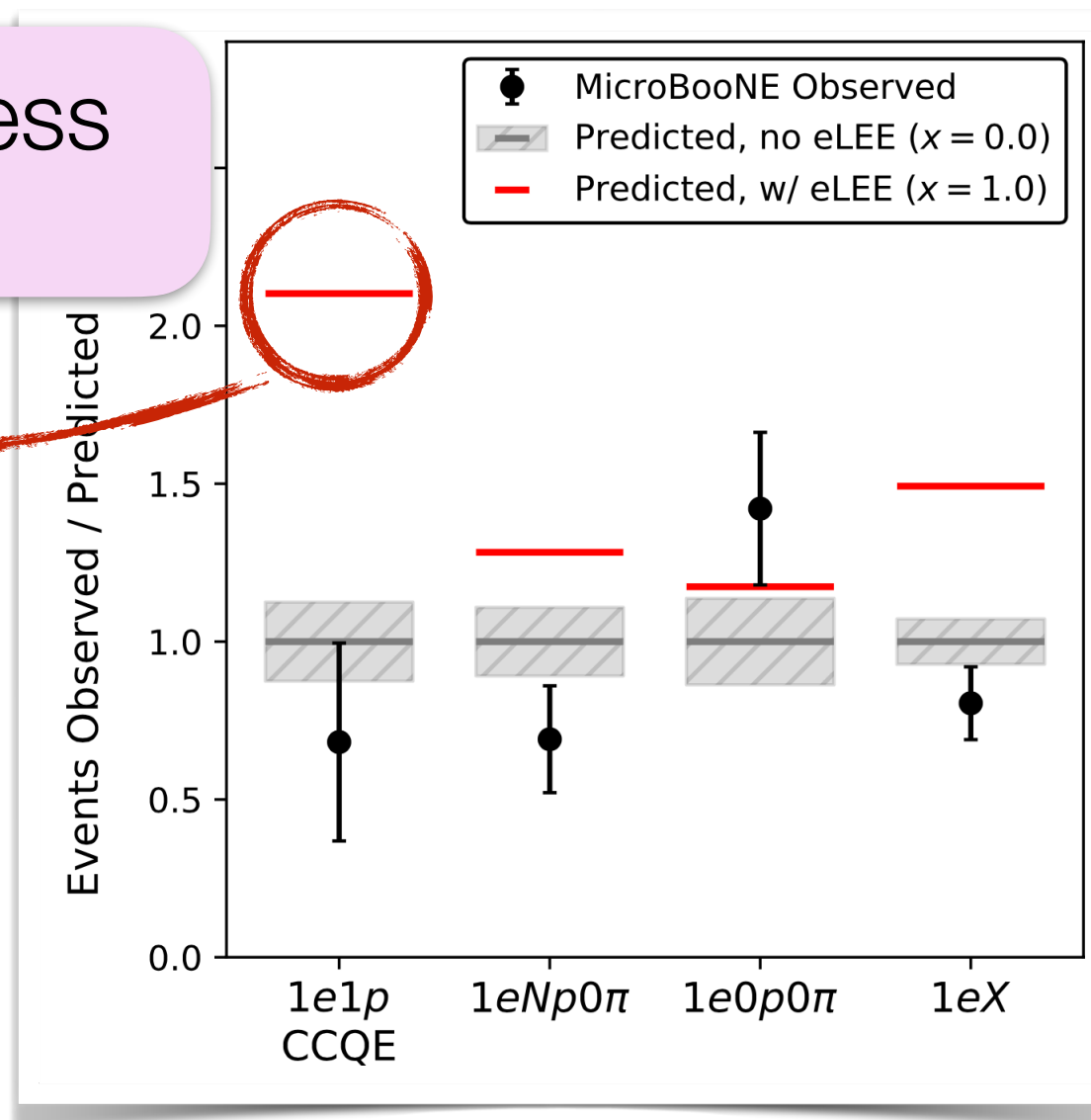
- ☑ anomalous  $\nu_e$  appearance tested by MicroBooNE in several channels
- ☑ No support for interpretation of MiniBooNE excess as  $\nu_e$



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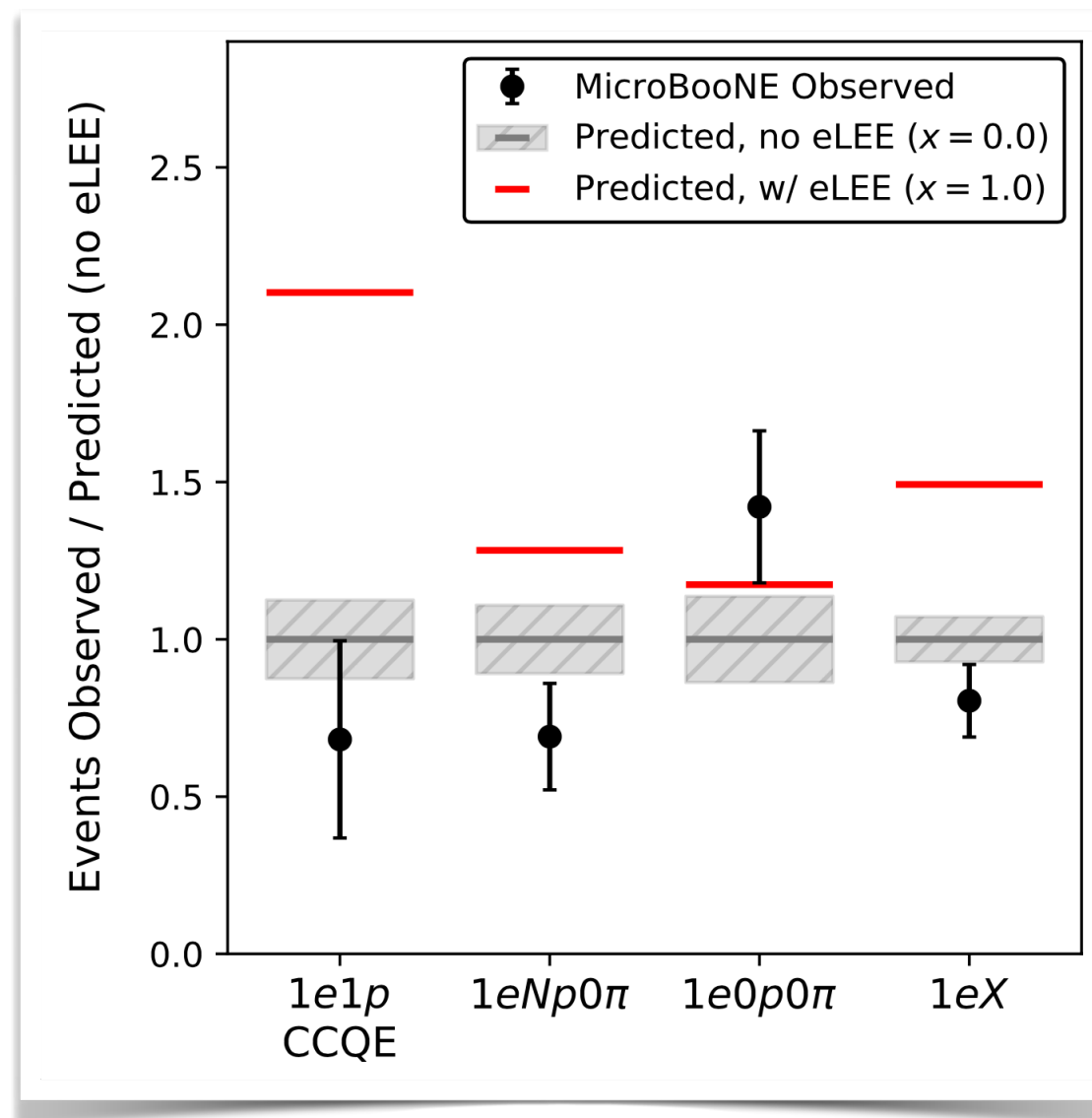
- ✓ anomalous  $\nu_e$  appearance tested by MicroBooNE in several channels
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MiniBooNE excess  
central value



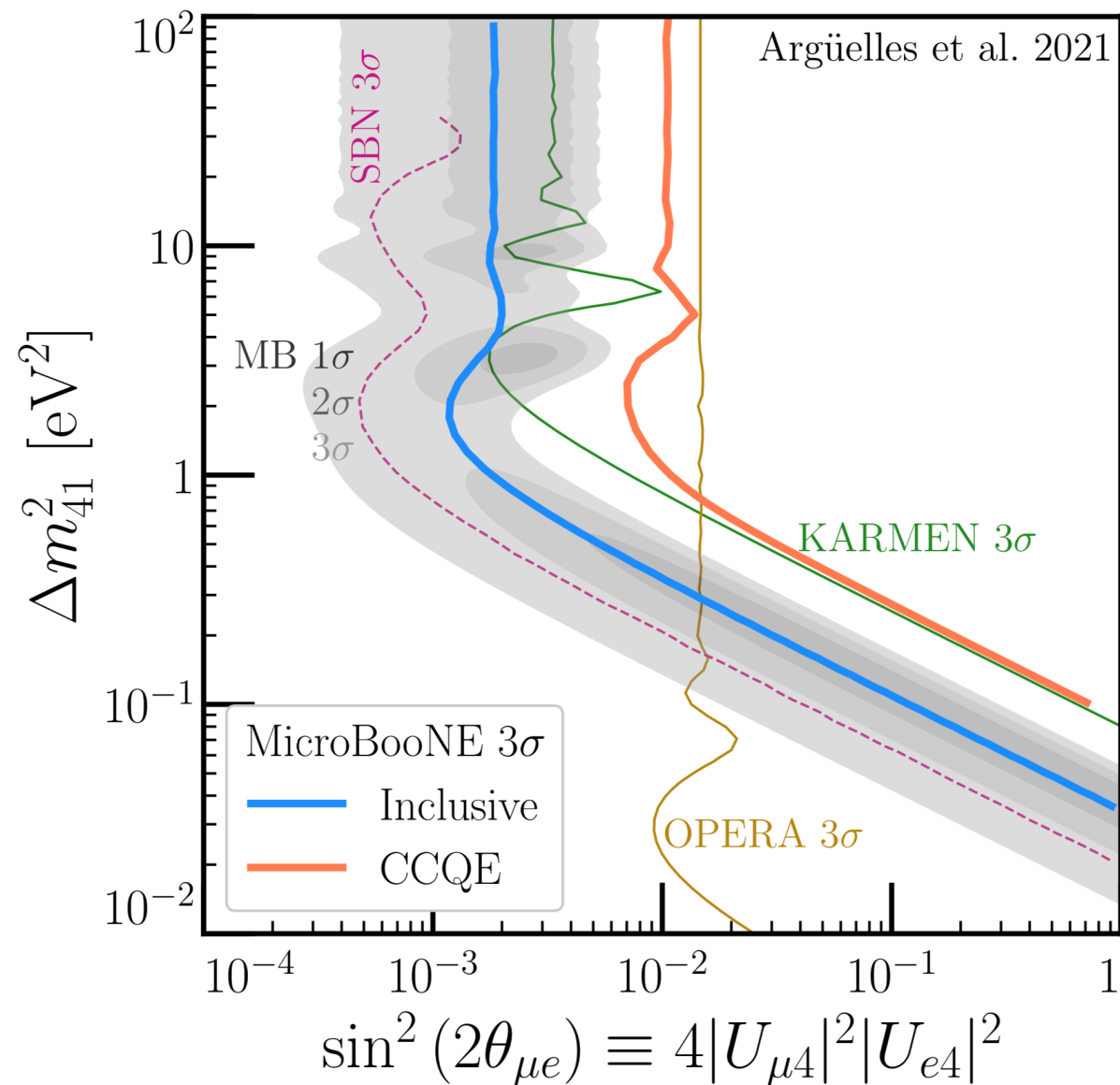
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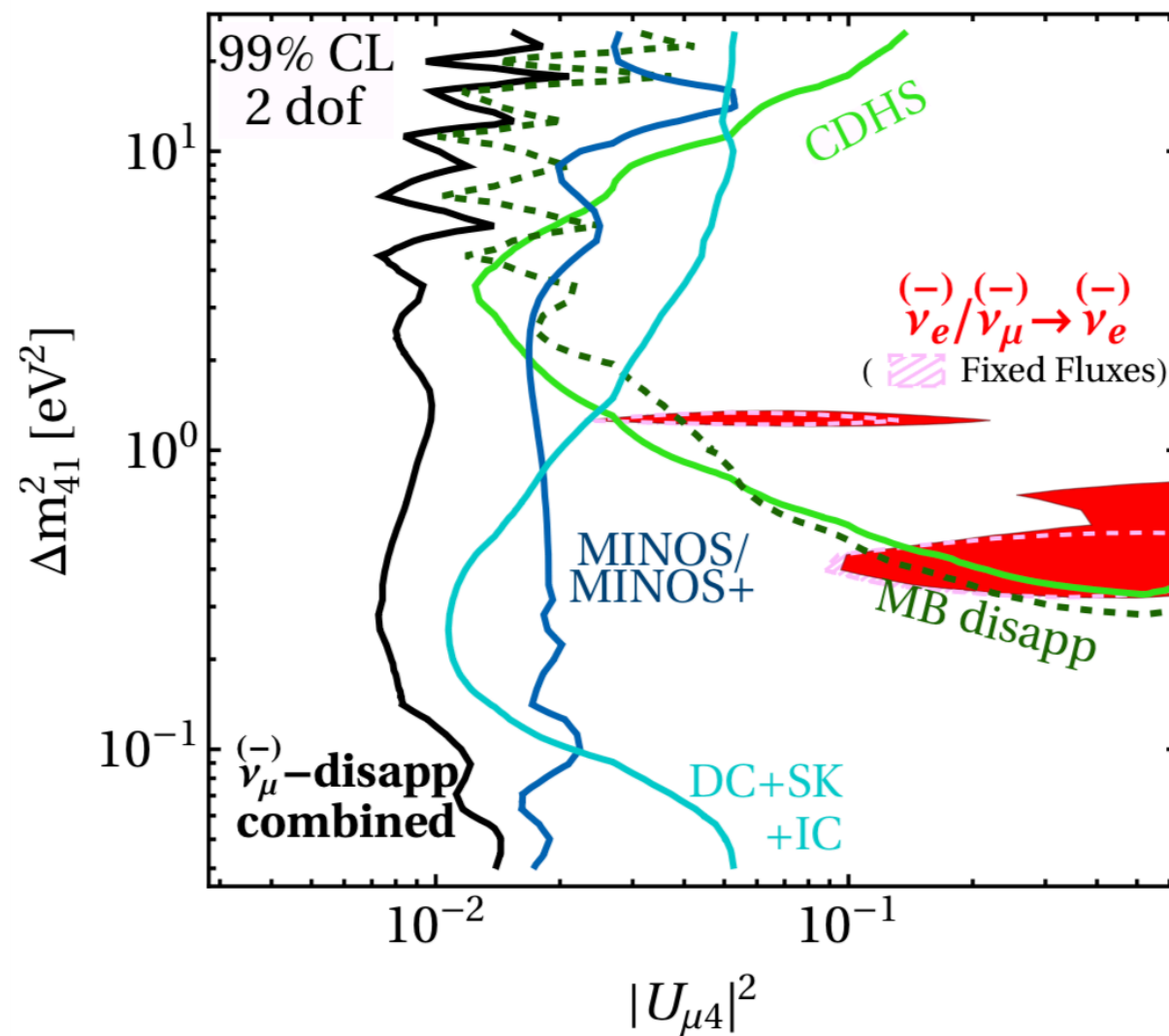
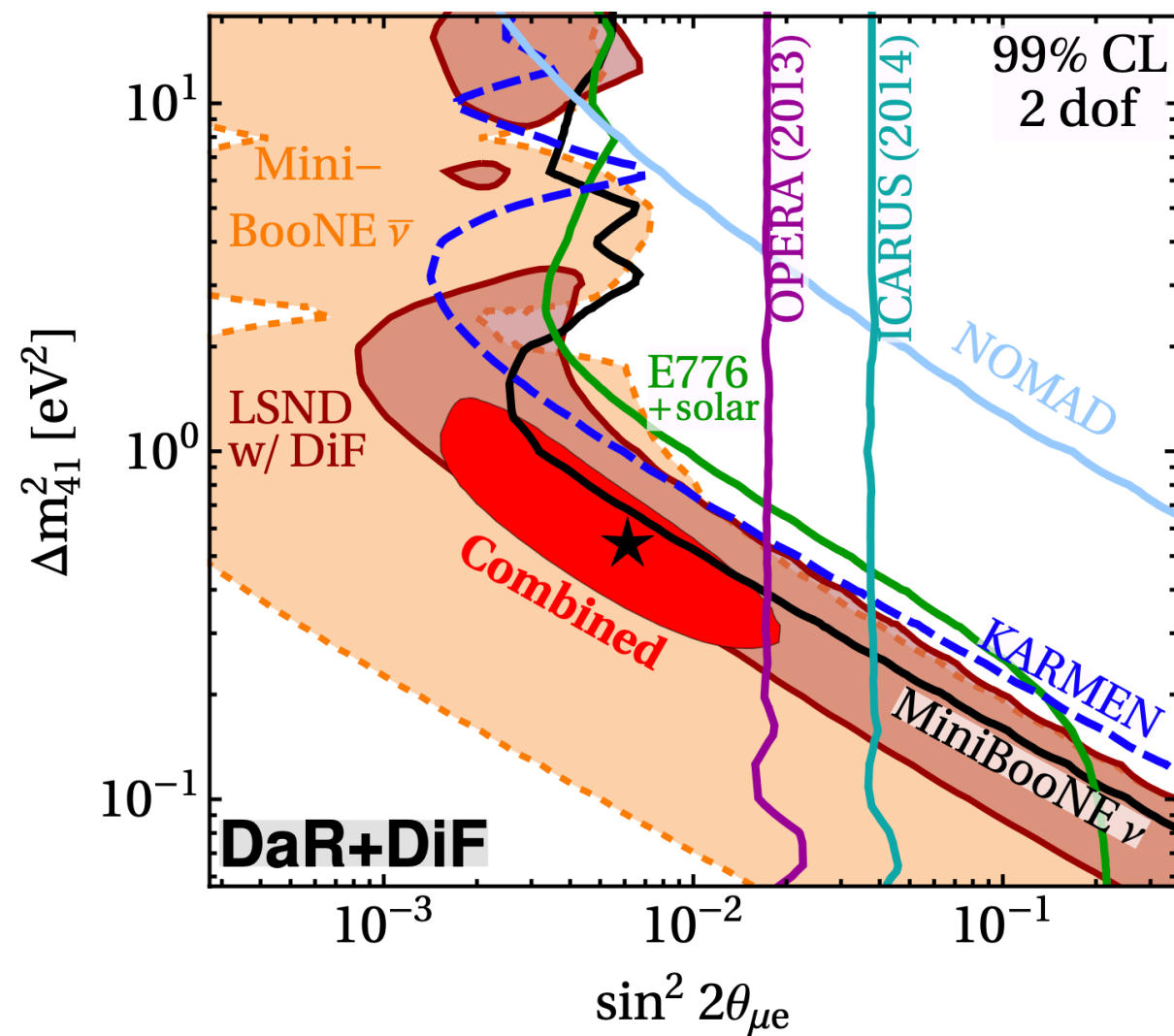
# Compatibility Between MiniBooNE and $\mu$ BooNE



- ✓ 2σ regions overlap
  - ➡ poster #667 by Ivan Martinez-Soler and poster #152 by Nick Kamp
- ✓ good sensitivity due to downward fluctuation in  $\nu_e$ , upward fluctuation in  $\nu_\mu$
- ✓ more conservative / robust limits from CLs method
  - ➡ poster #152 by Nick Kamp
- ✓ stay tuned for Hanyu Wei's talk tomorrow

Argüelles Esteban Hostert Kelly  
JK Machado Martinez-Soler Perez-Gonzalez  
arXiv:2111.10359

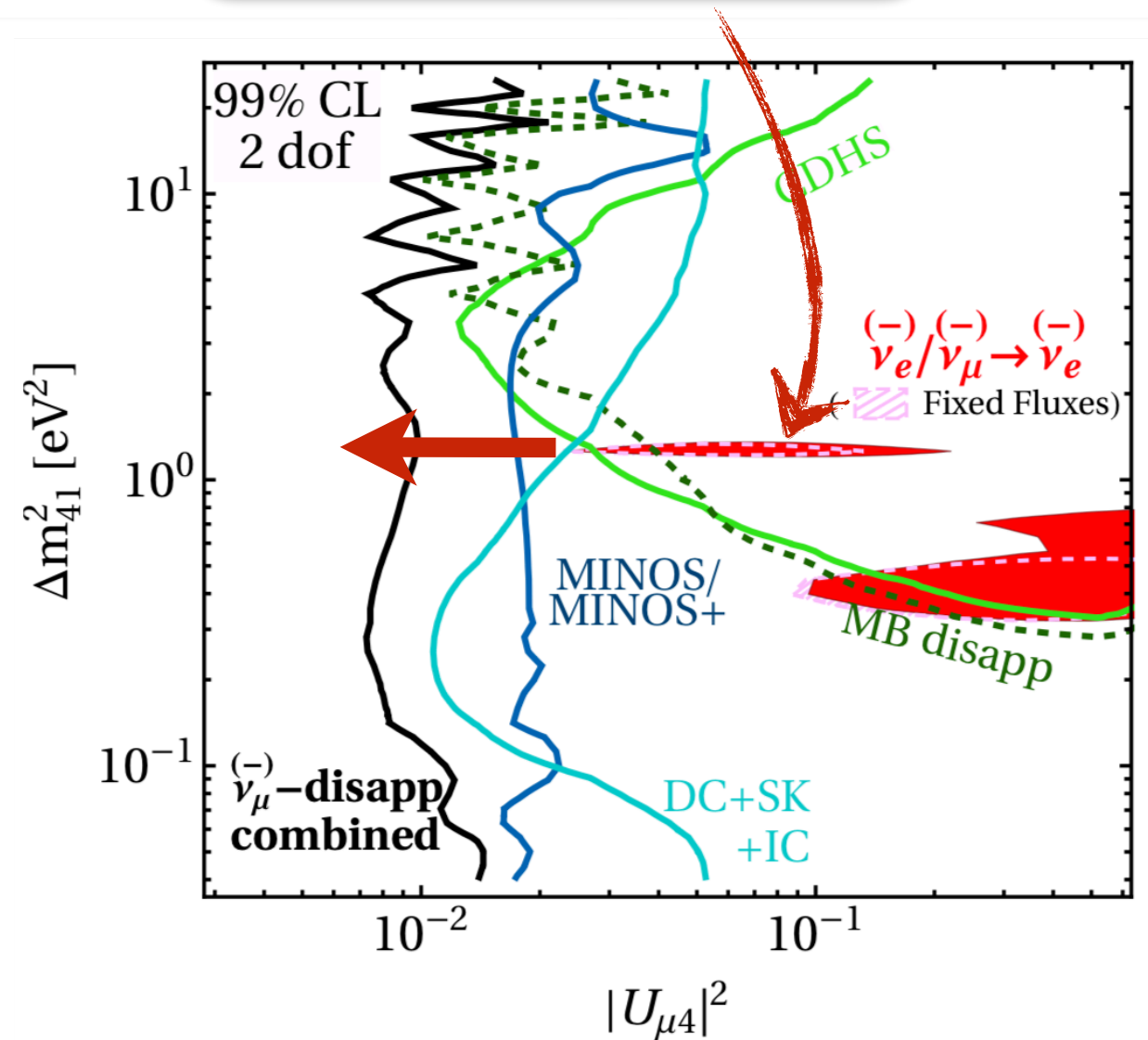
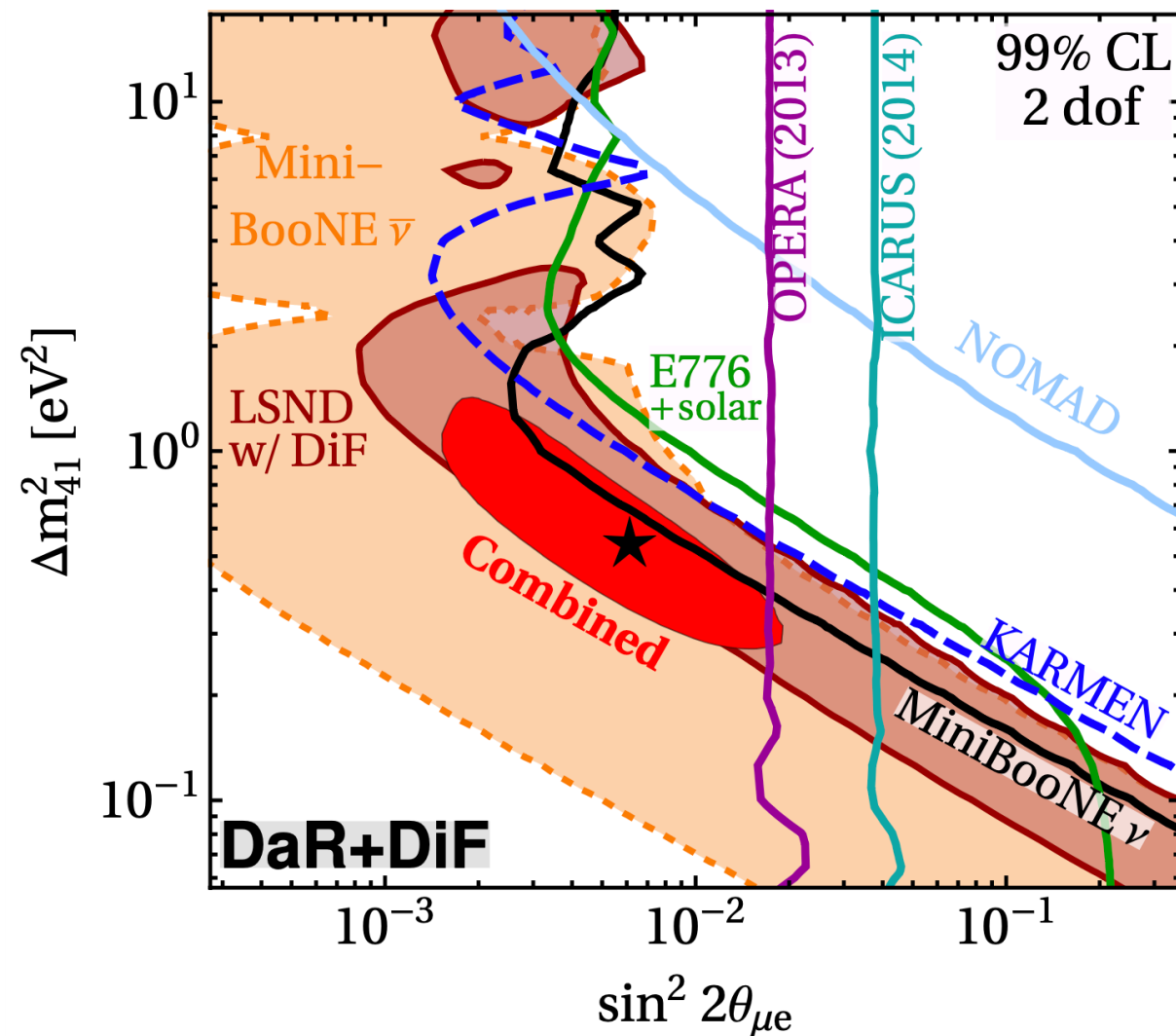
# Recent Updates





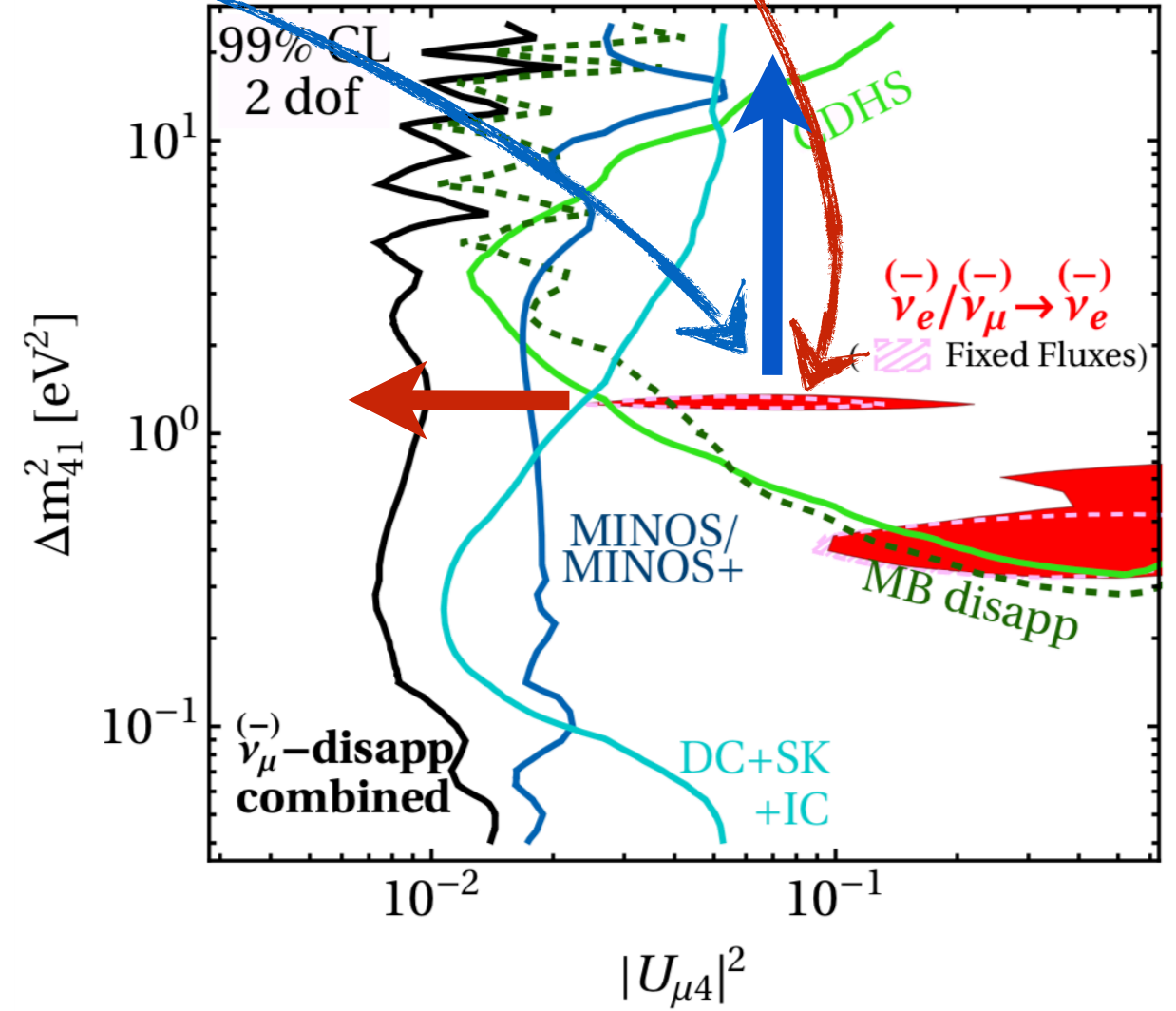
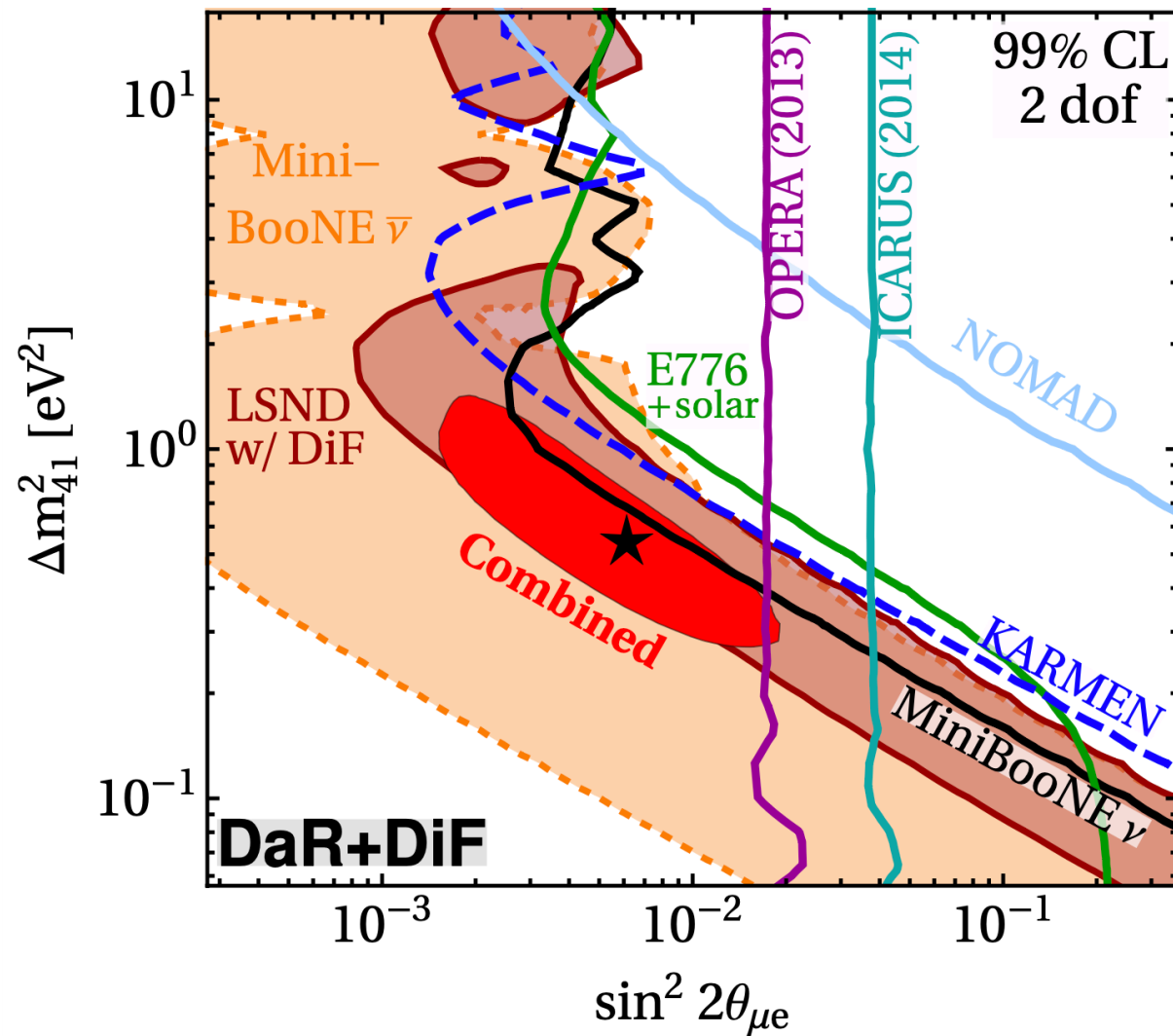
# Recent Updates

**BEST** and  **$\mu$ BooNE**  
push towards lower  $|U_{\mu 4}|^2$



**BEST and Neutrino-4**  
push towards larger  $|\Delta m_{41}|^2$

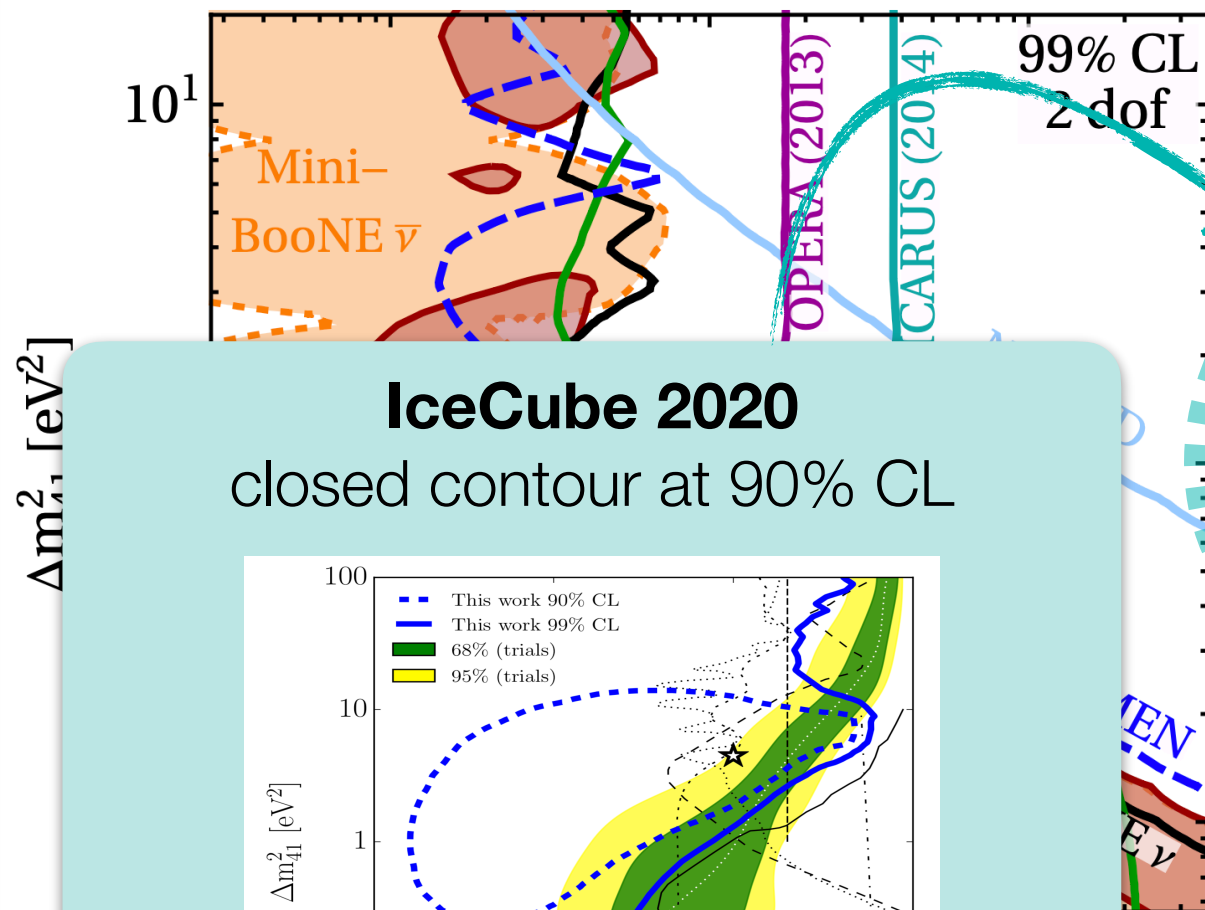
**BEST and  $\mu$ BooNE**  
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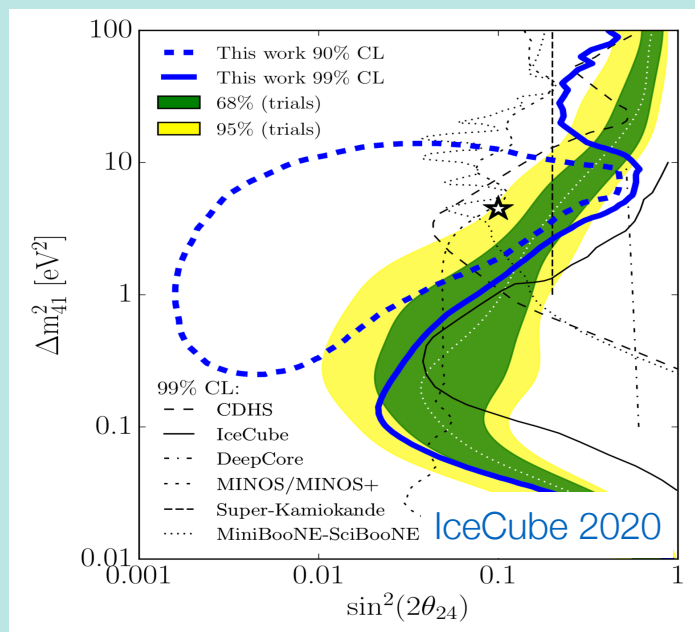
# Recent Up

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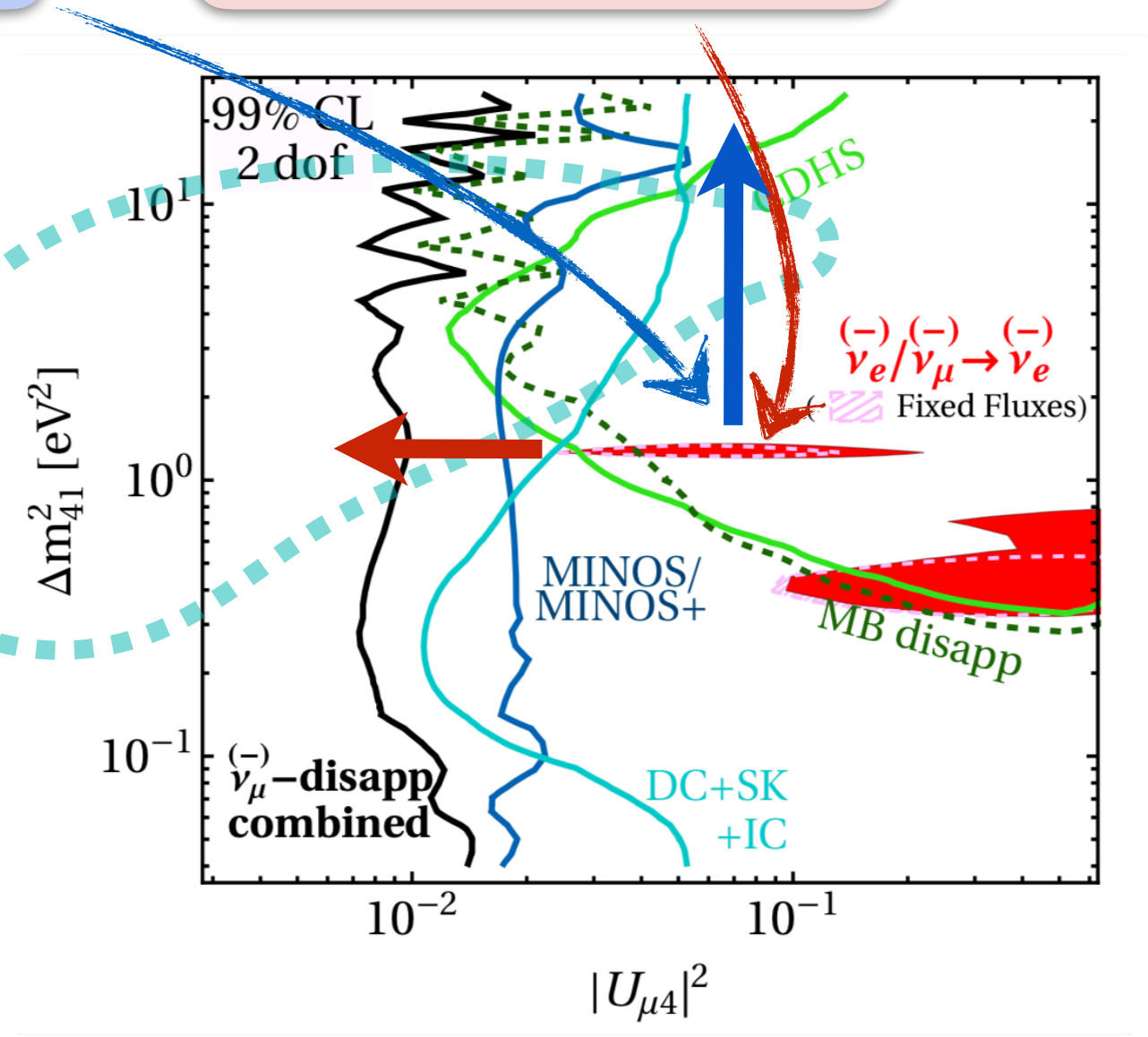
**BEST and  $\mu$ BooNE**  
push towards lower  $|U_{\mu 4}|^2$



**IceCube 2020**  
closed contour at 90% CL



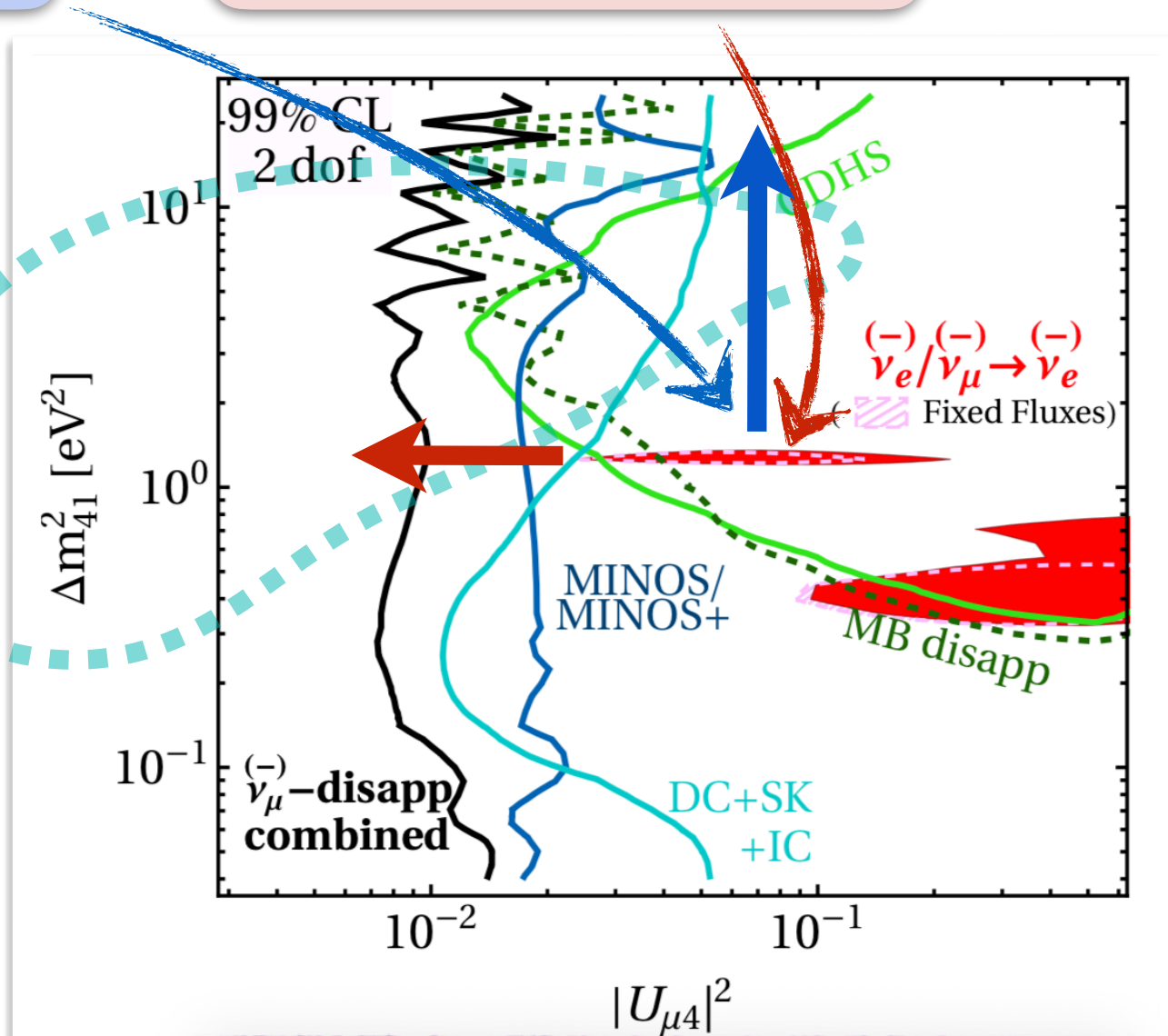
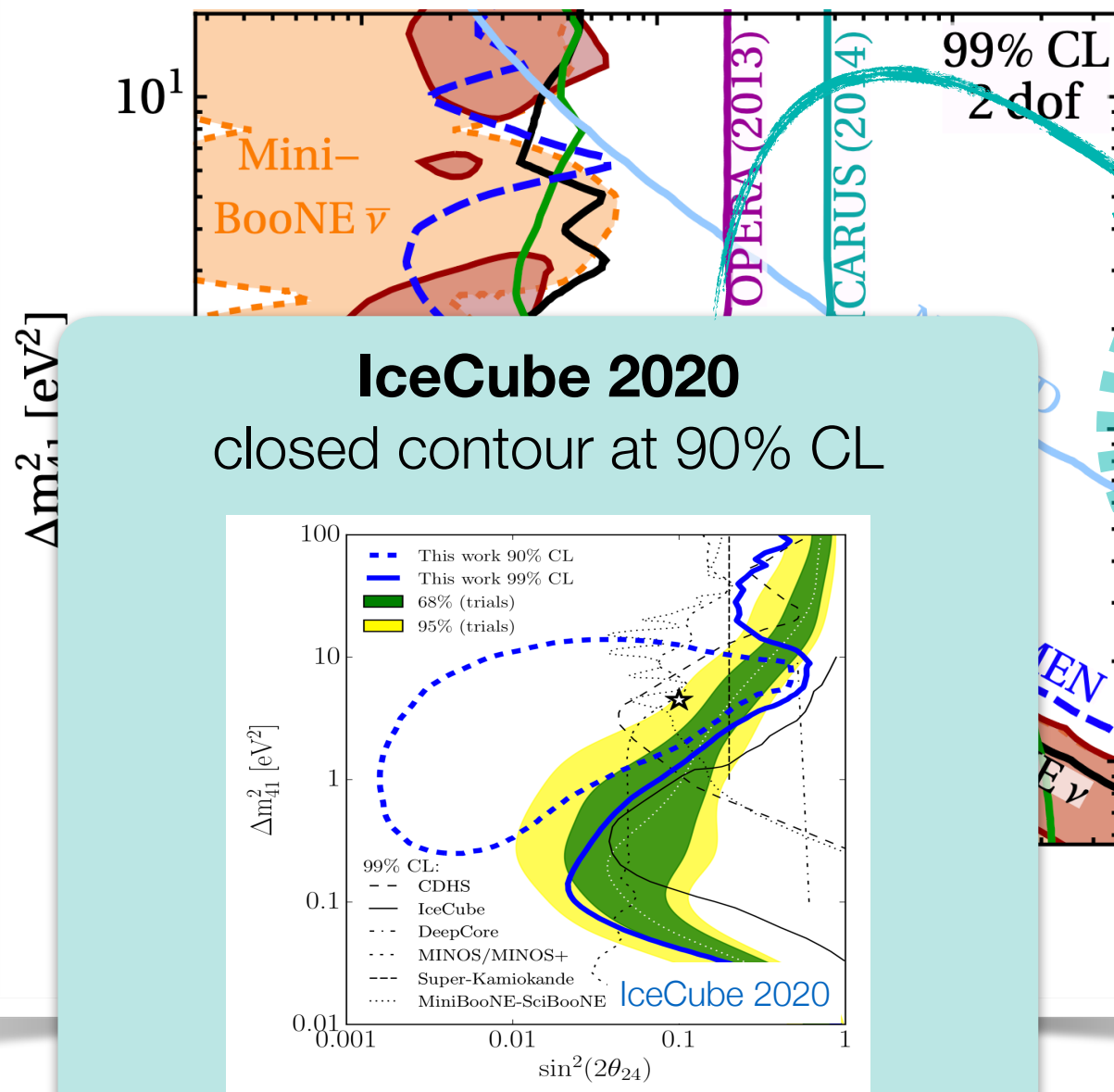
posters #105 (Bejnamin Smithers)  
& #279 (Alfonso Garcia-Soto)



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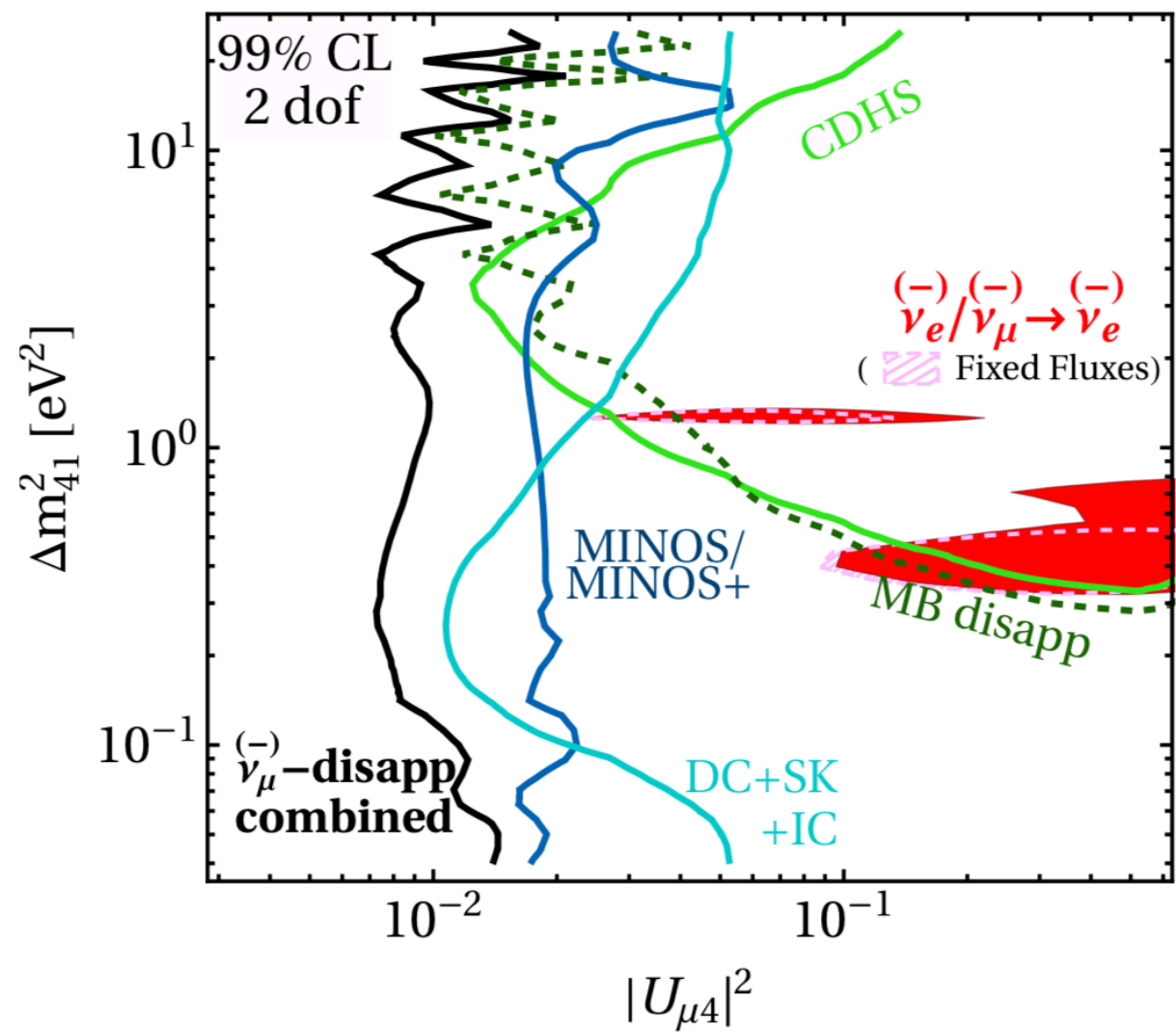
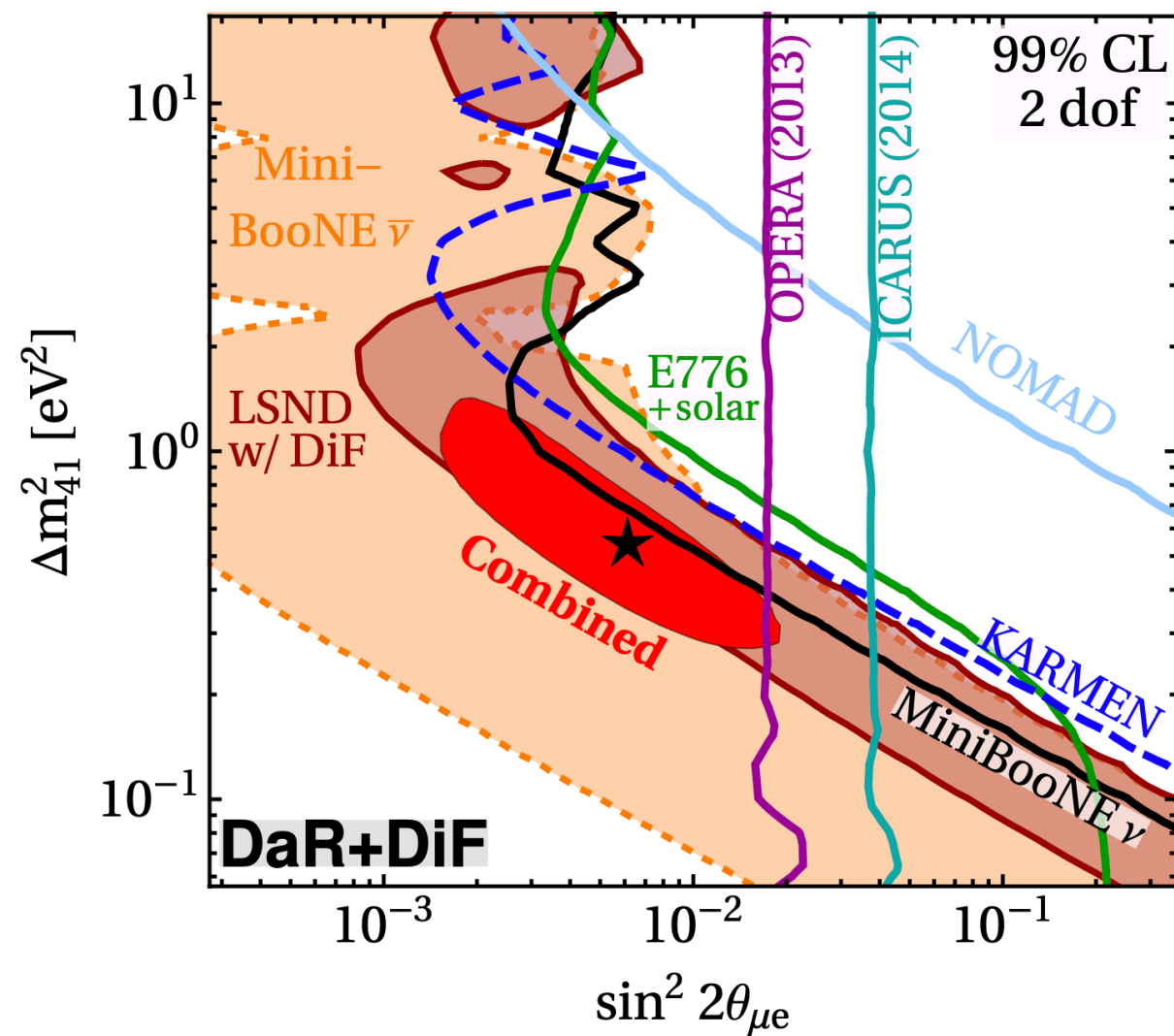
► posters #105 (Bejnamin Smithers)  
& #279 (Alfonso Garcia-Soto)

Expect tension to *increase*

- reactor fluxes vs. BEST
- MiniBooNE vs.  $\mu$ BooNE



# Recent Updates

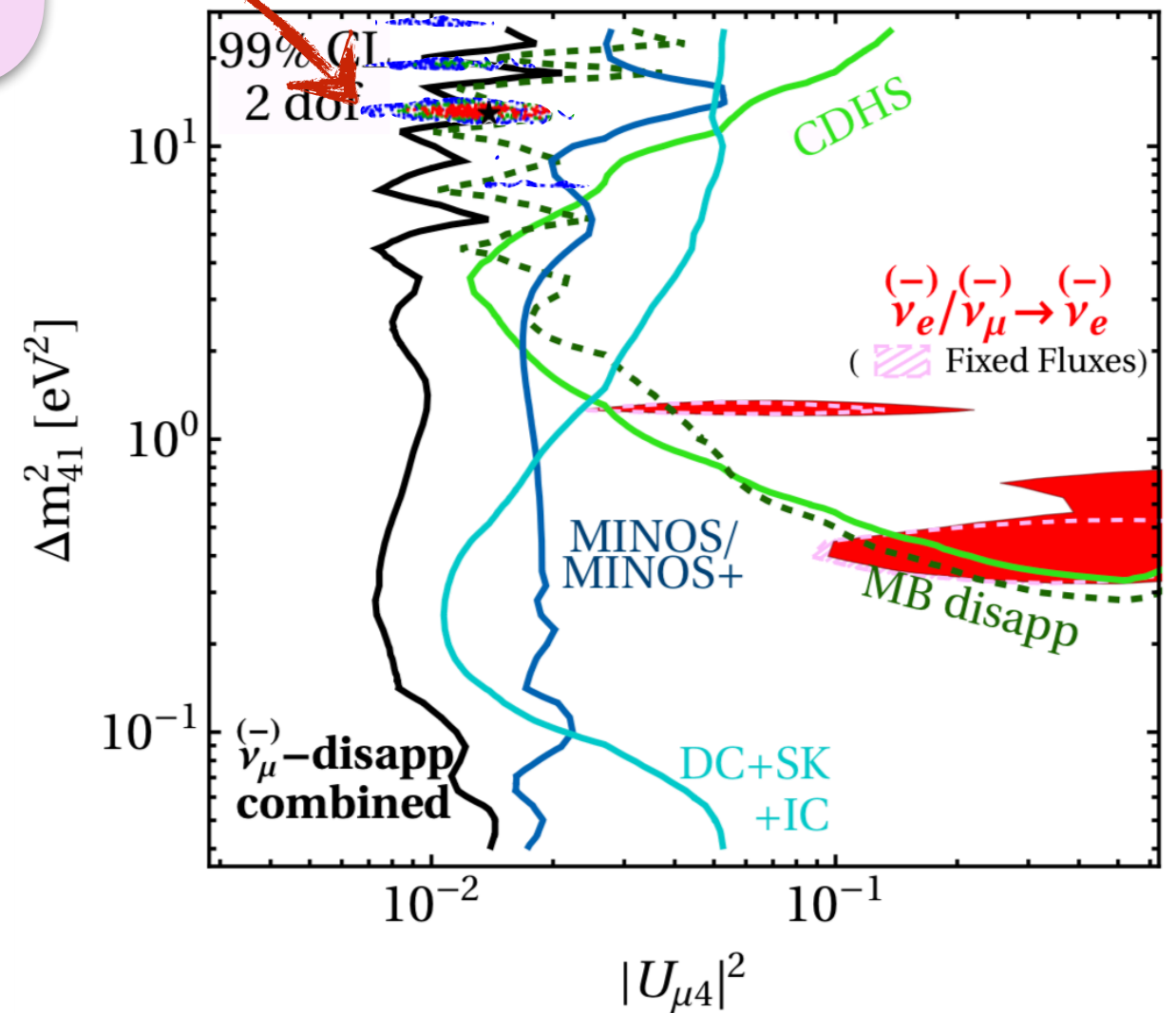
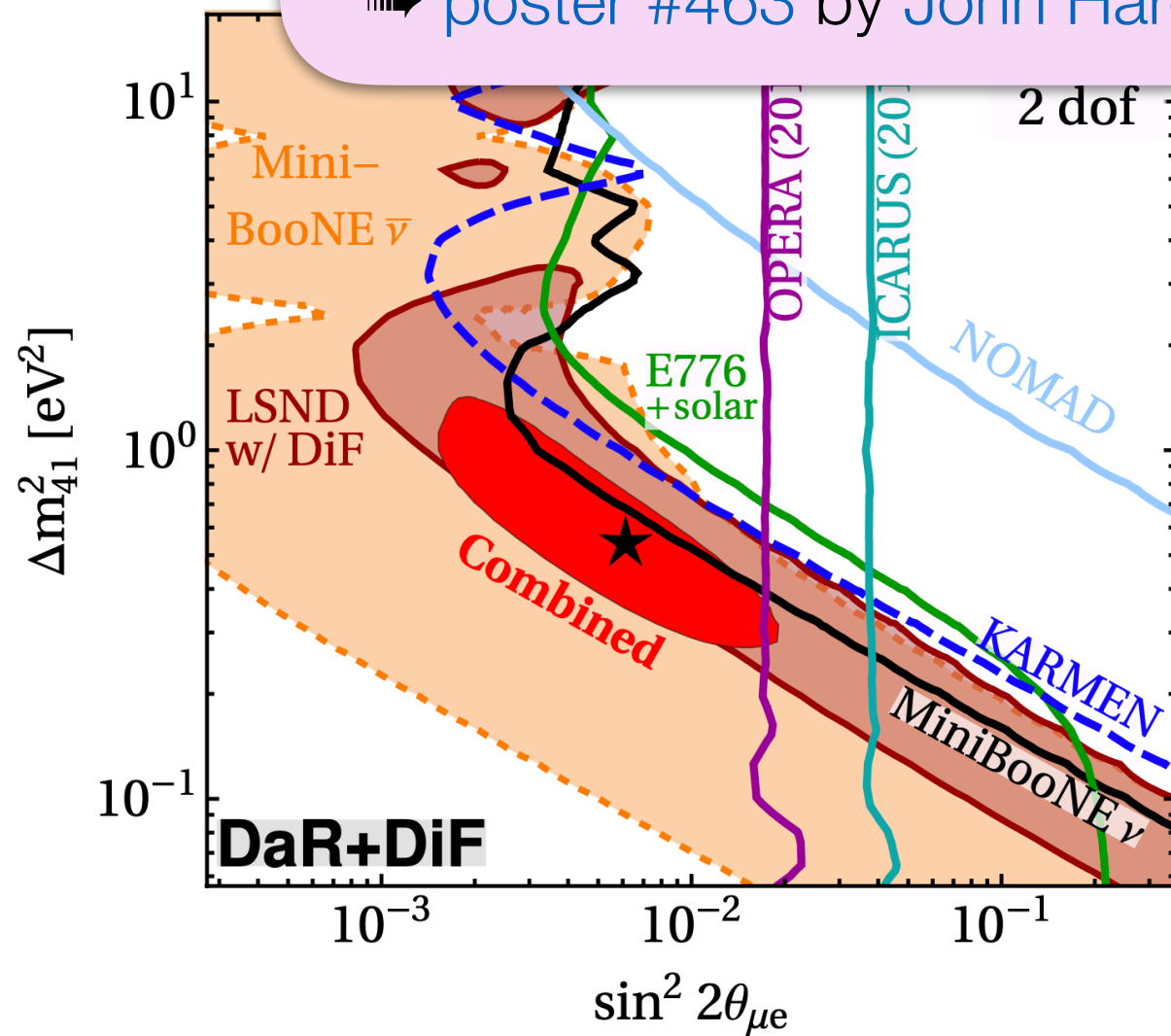


# Recent Updates

## new global fit

by the Columbia–Harvard–MIT group

➡ poster #463 by John Hardin



# Sterile Neutrinos in Cosmology

An extra neutrino species is in **severe tension with cosmology**.

Standard picture:  $\nu_s$  production via oscillation at  $T \gtrsim \text{MeV}$

$$N_{\text{eff}} \approx 3.16 \quad \text{⚡}$$

$$\Sigma m_\nu \approx 0.12 \text{ eV} \quad \text{⚡}$$

Planck TT,TE,EE+lowE+lensing+BAO 2018

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measure for the

**energy density in relativistic particles**

extra neutrino species would imply  $N_{\text{eff}} \sim 4$

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**sum of neutrino masses**

affects structure formation

sterile neutrino compatible with anomalies

would imply  $\Sigma m_\nu \sim 1 \text{ eV}$

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would imply  $\Sigma m_\nu \sim 1 \text{ eV}$

**potentially solvable in extended models**

➡ talk by [Olga Mena](#) on Saturday

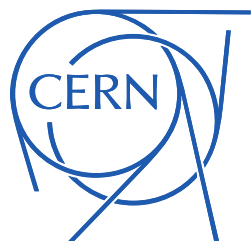
Back to the Future

# The Future

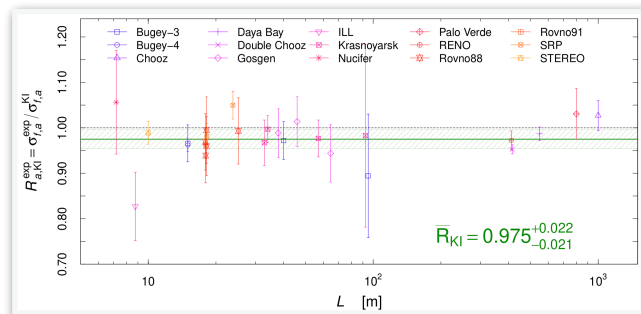
- ☑ Look for **more exotic signatures** (no known SM candidates)
  - single- $\gamma$  production unrelated to the  $\Delta(1232)$ ?
  - single electrons unrelated to  $\nu_e$ ?
  - boosted  $e^+e^-$  pairs?

⇒ talk by **Carlos Argüelles** tomorrow
- ☑ More data incoming:
  - $\mu$ BooNE / SBND / ICARUS at Fermilab
    - ⇒ talks by **Hanyu Wei** and **Anne Schukraft** tomorrow
  - SBL reactor + gallium experiments
    - ⇒ **Matthieu Licciardi**, **Jinyu Kim** today, **Steve Elliott** tomorrow
  - JSNS-2 ⇒ talk by **Jungsic Park** tomorrow
  - Long-baseline Experiments, Neutrino Telescopes, ...
  - Future Proposals (IsoDAR, ...) ⇒ **poster session**

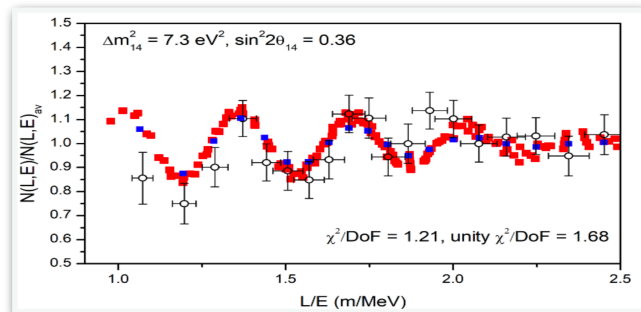
# Summary



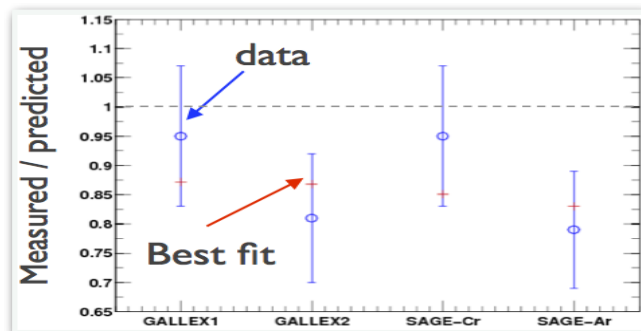
# Summary



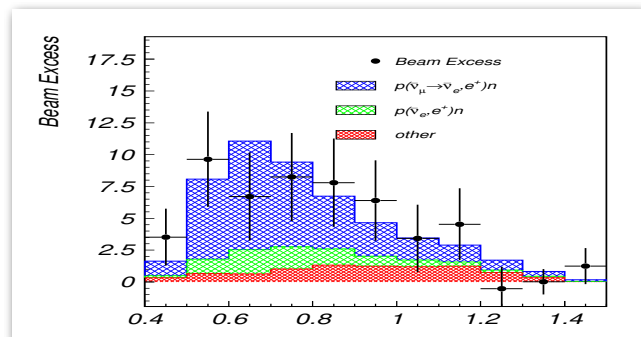
reactor flux anomaly  
resolved with new input data  
to flux calculation



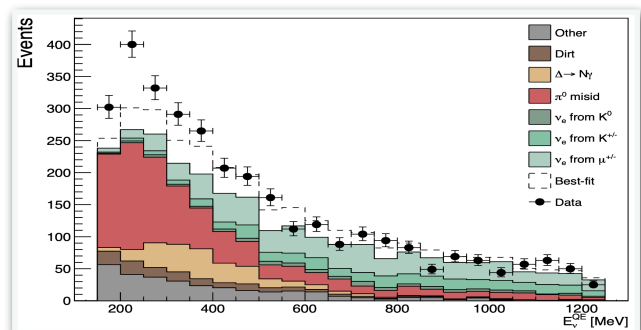
reactor spectra  
is there really an anomaly?



gallium anomaly  
unresolved, recently reinforced



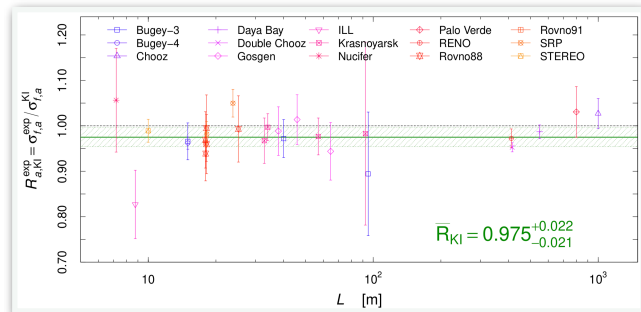
LSND  
unresolved



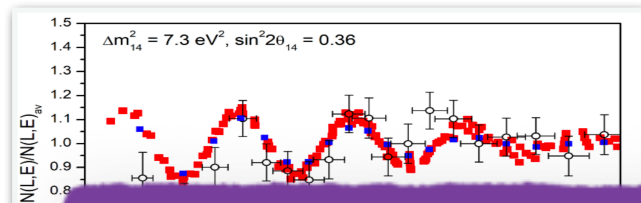
MiniBooNE  
unresolved  
resolvable by next-gen. SBL experiments



# Summary



reactor flux anomaly  
resolved with new input data  
to flux calculation



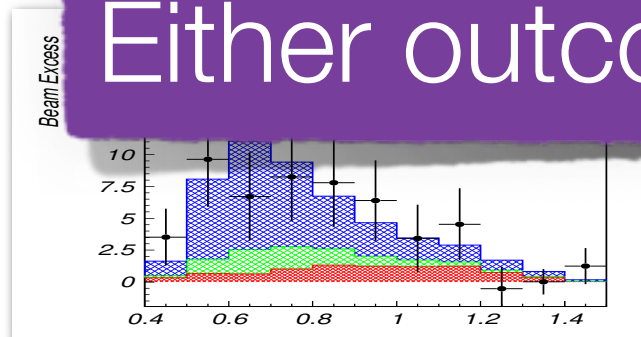
reactor spectra



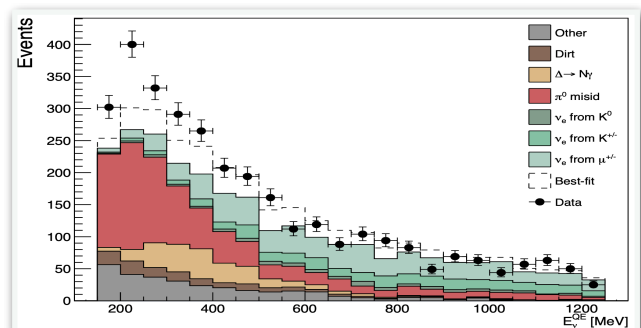
★ scrutinize anomalies for unknown systematics  
(need 4 independent effects!)

★ scrutinize also null results?

Either outcome **will teach us crucial lessons.**



unresolved



MiniBooNE

unresolved

resolvable by next-gen. SBL experiments



# Thank You!

