



The 50th Anniversary

Chris Chirg



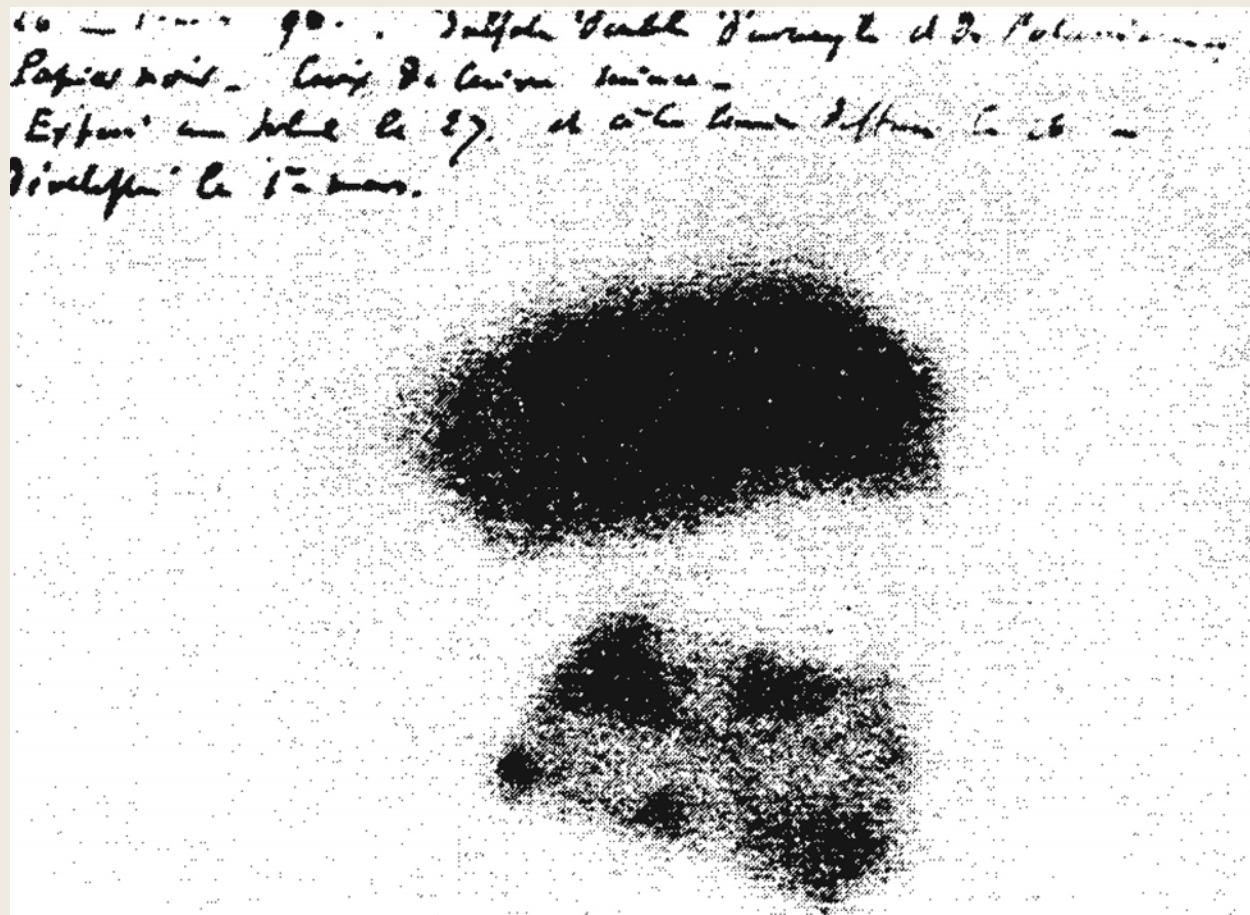


Benjamin Whisoh Lee (1935–1977)

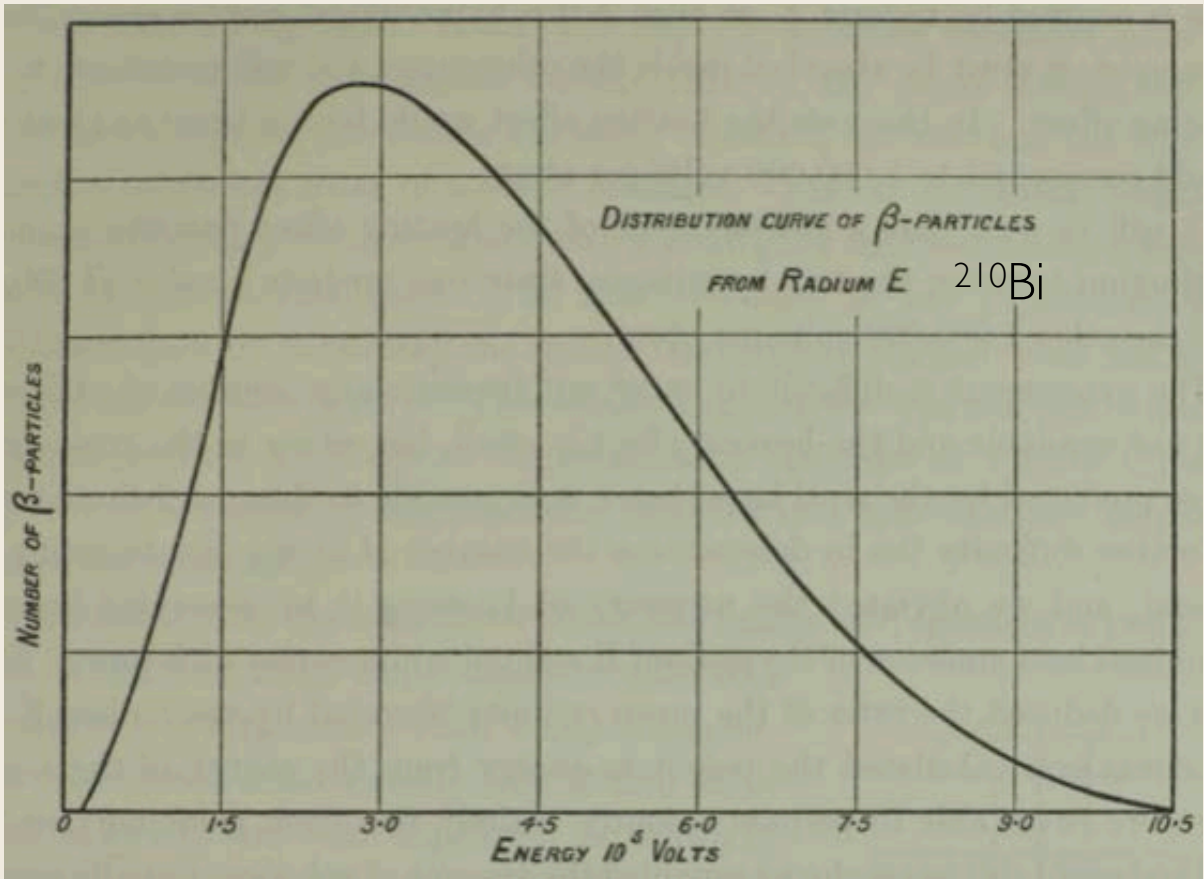
Neutrino '72, Balatonfüred, Hungary



Participants of Neutrino '72 conference. In the front row: T. D. Lee, G. L. Radicati, R. P. Feynman, B. Pontecorvo, G. Marx, V. F. Weisskopf, F. Reines, C. L. Cowan and P. Budini



Becquerel, 1896



Chadwick, 1914;
Ellis-Wooster, 1927

Physikalisches Institut
der Eidg. Technischen Hochschule
Zürich

Zürich, 4. Des. 1930
Olariastrasse

Liebe Radioaktive Damen und Herren,

Wie der Ueberbringer dieser Zeilen, den ich höflichst
anzuhören bitte, Ihnen des näheren auseinandersetzen wird, bin ich
angesichts der "falschen" Statistik der N - und $Li-6$ Kerne, sowie
des kontinuierlichen beta-Spektrums auf einen verweifelten Ausweg
verfallen um den "Wechselsatz" (1) der Statistik und den Energiesatz
zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale
Teilchen, die ich Neutronen nennen will, in den Kernen existieren,
welche den Spin $1/2$ haben und das Ausschliessungsprinzip befolgen und
sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie
nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen
müsste von derselben Grossenordnung wie die Elektronenmasse sein und
jedenfalls nicht grösser als $0,01$ Protonenmasse.- Das kontinuierliche
beta-Spektrum wäre dann verständlich unter der Annahme, dass beim
beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert
wird, derart, dass die Summe der Energien von Neutron und Elektron
konstant ist.

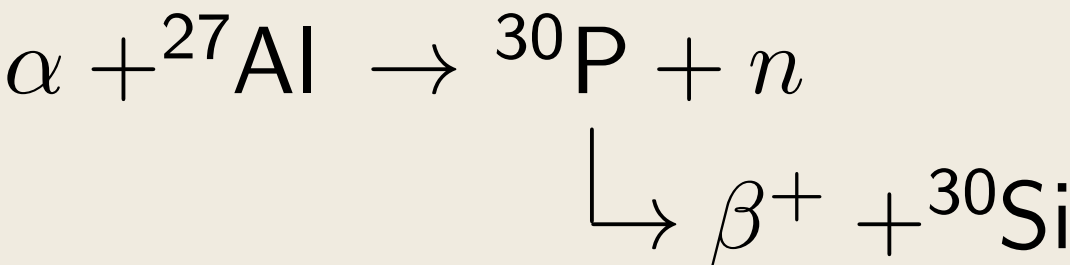
Pauli, 1930

TENTATIVO DI UNA TEORIA DEI RAGGI β

Nota (1) di ENRICO FERMI (1932-4)

Sunto. - Si propone una teoria quantitativa dell'emissione dei raggi β
in cui si ammette l'esistenza del « neutrino » e si tratta l'emissione degli
elettroni e dei neutrini da un nucleo all'atto della disintegrazione β con
un procedimento simile a quello seguito nella teoria dell'irradiazione
per descrivere l'emissione di un quanto di luce da un atomo eccitato.
Vengono dedotte delle formule per la vita media e per la forma dello
spettro continuo dei raggi β , e le si confrontano coi dati sperimentali.

β^+ -emission observed



Curie & Joliot, 1934

Other rare processes foreseen:
Inverse β decay Bethe & Peierls, 1934
 $\nu\nu\beta\beta$ decay Goeppert-Mayer 1935
and, following Majorana 1937,
 $0\nu\beta\beta$ decay Furry, 1939

1932: n (Chadwick), e^+ (Anderson)

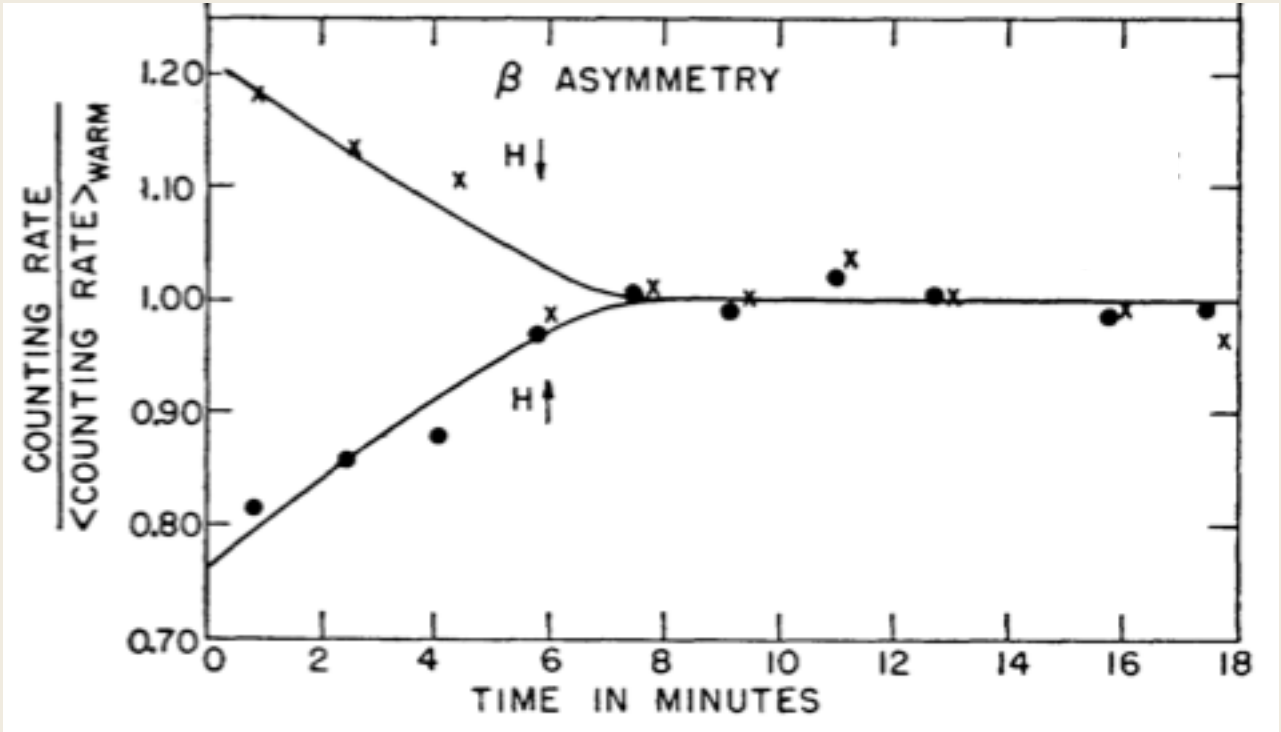
1932-1946: Leptonic nature of μ

Detection of the Free
Neutrino: a Confirmation

C. L. Cowan, Jr., F. Reines, F. B. Harrison,
H. W. Kruse, A. D. McGuire

$\bar{\nu}_p \rightarrow e^+ n$, 1956

P violation in polarized ${}^{60}\text{Co}$ β decay



Wu, Ambler, ..., 1957

P and C violation in $\pi \rightarrow \mu \rightarrow e$ chain

Es ist uns eine bange Nacht,
bekannt im Leben, das unsere
langjähige, liebe Frau die
PARITY
am 19. Januar 1957 nach kurzem
Leben bei ~~starkem~~ experimentellen
Eingriffen sanft entschlafen ist.
Für die Hinterbliebenen
 e, μ, ν .

Pauli to Weisskopf

Garwin, et al., 1957

ν mass too small to measure,
 ν helicity = $-\frac{1}{2}$
 \leadsto Universal Fermi Interaction (V–A)
 with only ν_L
 “two-component neutrino”
 CVC hypothesis

1960s: Golden Age of β decay

High-energy π, K decay $\rightarrow \nu$ beams

$$V_\mu \neq V_e$$

OBSERVATION OF HIGH-ENERGY NEUTRINO REACTIONS AND THE EXISTENCE OF TWO KINDS OF NEUTRINOS*
 G. Danby, J.-M. Gaillard, K. Goulianos, L. M. Lederman, N. Mistry, M. Schwartz,[†] and J. Steinberger[†]
 Columbia University, New York, New York and Brookhaven National Laboratory, Upton, New York
 (Received June 15, 1962)

Cf. Wu & Moszkowski, 1966

Possibility of $\nu_\mu - \nu_e$ mixing
 \leadsto oscillations

$\nu \leftrightarrow \bar{\nu}$ oscillations?

(Pontecorvo, 1957)
 Maki, Nakagawa, Sakata, 1962

$SU(3)_{\text{flavor}}$ symmetry inspires
 Cabibbo universality, 1963

$$J_\lambda^{(+)} = \bar{u} \gamma_\lambda (1 - \gamma_5) d \cos \theta_C + \bar{u} \gamma_\lambda (1 - \gamma_5) s \sin \theta_C$$

\leadsto current algebra,
 Kobayashi–Maskawa 3×3 , 1973

Rising $\sigma(\nu_\mu e \rightarrow \mu \nu_e) \propto E_{\text{cm}}^2$
 \leadsto new physics by $E_{\text{cm}} \approx 300$ GeV
 Second-order needs cutoff

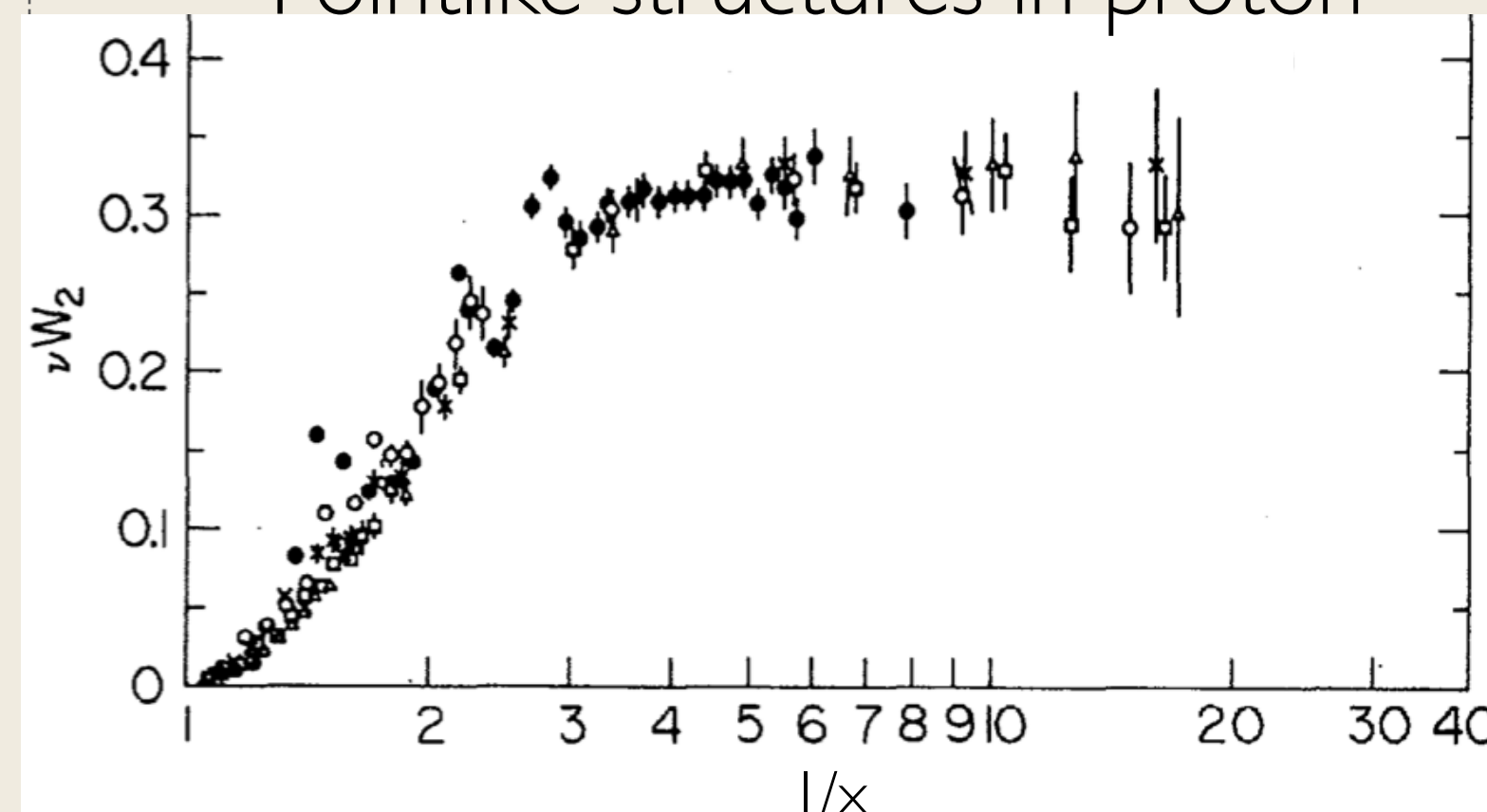
intermediate boson?

CP violation in weak decay

EVIDENCE FOR THE 2π DECAY OF THE K_2^0 MESON*[†]
 J. H. Christenson, J. W. Cronin,[‡] V. L. Fitch,[‡] and R. Turlay[§]
 Princeton University, Princeton, New Jersey
 (Received 10 July 1964)

A new *superweak* interaction?

Pointlike structures in proton



SLAC–MIT, 1969

Bubble chambers for ν physics



$\nu_\mu p \rightarrow \mu^- \pi^+ p$, ANL 12, 1970

Issues in the air approaching v'72*

Status and origin of $\Delta I = 1/2$ rule

Existence of Intermediate Vector Bosons, W^\pm

Testing lepton universality:

$$\sigma(\nu_e e) \leq 40\sigma_{V-A}; \sigma(\bar{\nu}_e e) \leq 40\sigma_{V-A}$$

Existence, properties of Neutral Currents

$K_L \rightarrow \mu^+ \mu^-$ Puzzle

CP Violation!

Search for second-class currents

Implications of Bjorken scaling, partons: CERN propane BC, $\sigma \propto E_\nu$

High-energy ν beams coming at Fermilab and CERN SPS

* Sam Treiman (1971)

From v'72 Summary Talks

Bruno Pontecorvo

“ambitious and difficult investigations in which somebody tries very hard to find and measure something, but does not see anything.”

1. $K_L \rightarrow \mu\mu$ puzzle (inequalities)
2. Solar neutrinos (upper limit)
3. Lepton charge conservation (upper limits)
4. “Stable heavy leptons” (“negative” results)
5. $\bar{\nu}e$ scattering, reactors (upper limits)
6. Neutral currents (upper limits)

“danger if you believe really in extraordinary things even before you are forced [to] by hard facts”

Viki Weisskopf

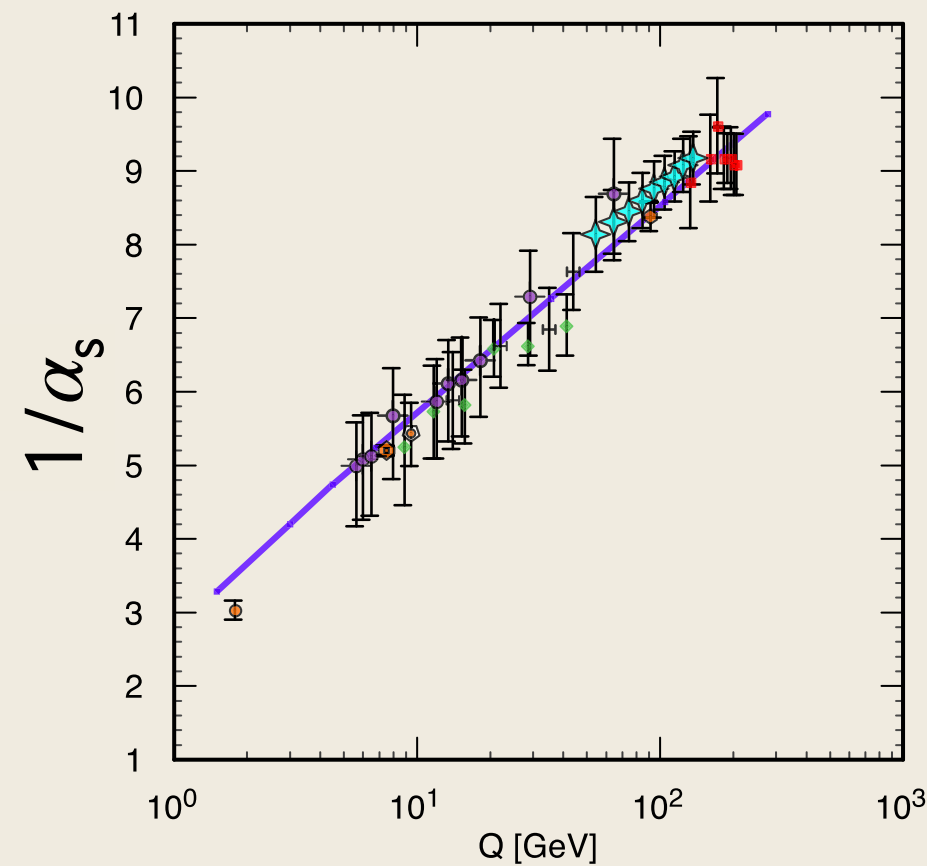
“You cannot talk about weak interactions alone.”

1. Troubles with V–A “Fermi” interaction:
unitarity, nonrenormalizability, **CP** violation
2. Weinberg’s Theory of Leptons
spontaneously broken $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_{EM}$
↪ neutral lepton currents
gauge cancellation in $e^+e^- \rightarrow W^+W^-$
“whole approach disregards **CP** violation”
3. Weak & EM: tools to investigate hadrons
Field theory vs. partons; sum rules as diagnostics;
quarks seemed to have Bose statistics

$\nu_{\mu}e \rightarrow \nu_{\mu}e$ specimen



Gargamelle, 1973



ASF \hookrightarrow QCD, 1973

Unified Theories \hookrightarrow Proton Decay

$$SU(5) \supset SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$$

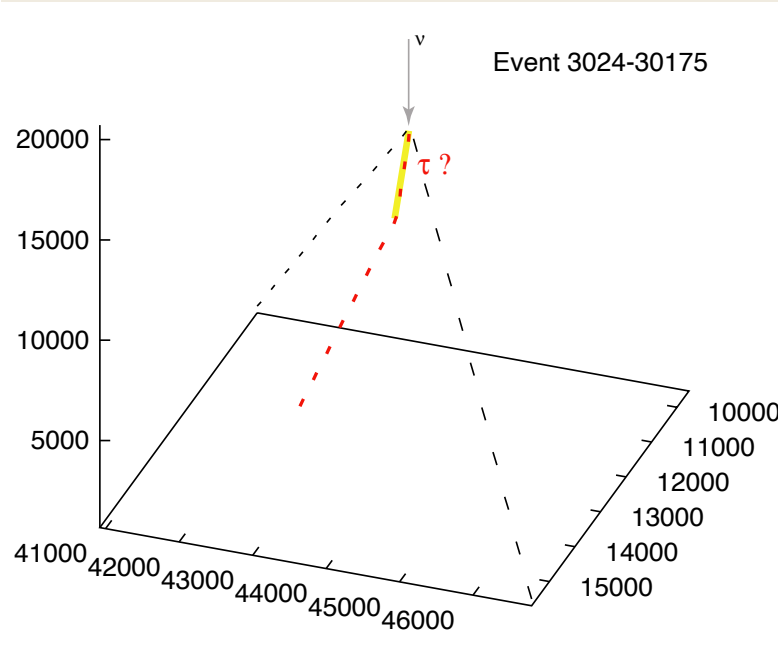
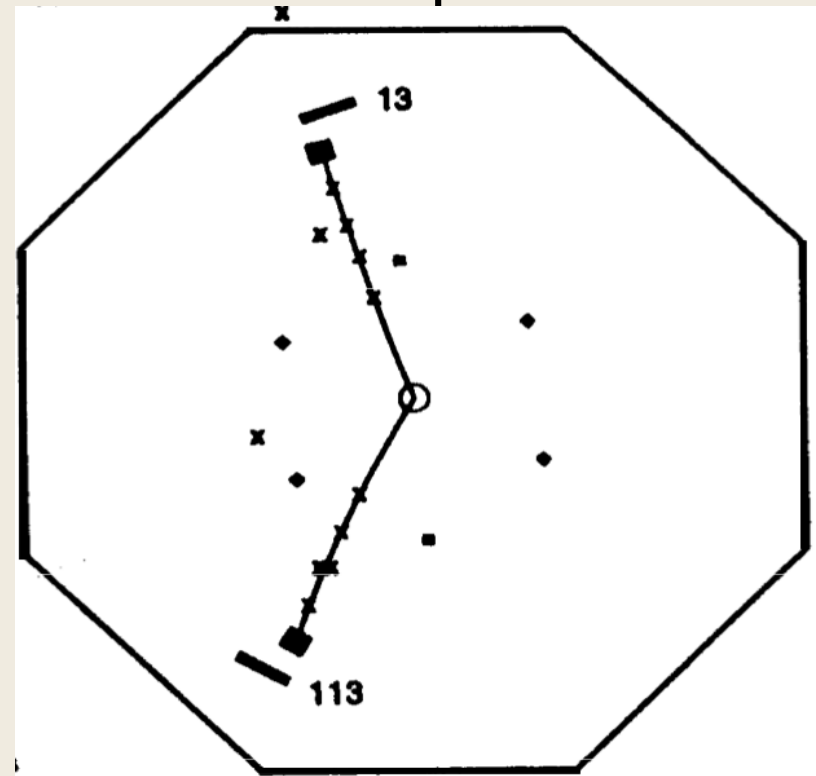
$$SO(10) \supset SU(4)_c \otimes SU(2)_L \otimes SU(2)_R$$

\Rightarrow Large-volume detectors

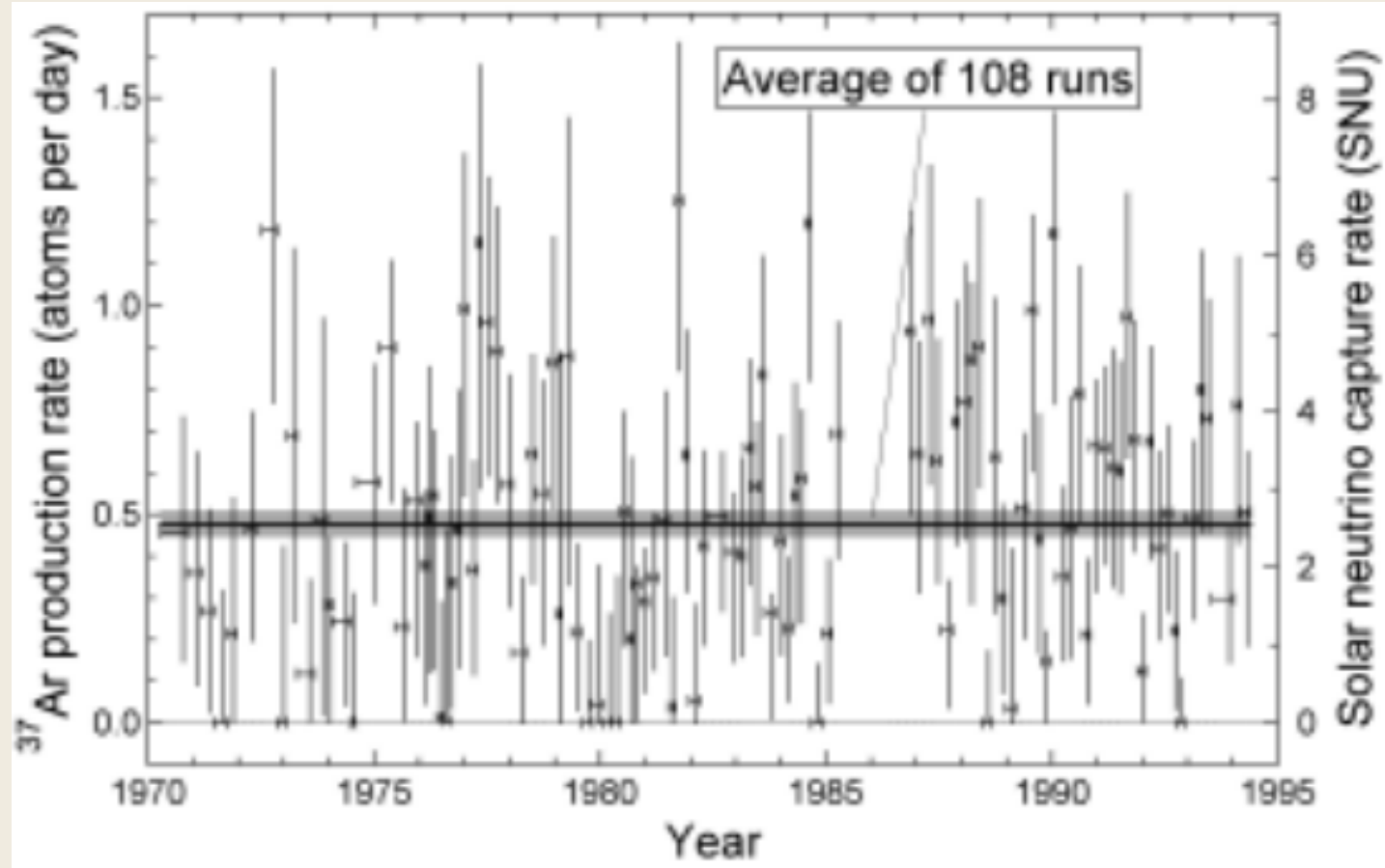
seesaw mechanism for light ν

1973–5

SPEAR $e\mu$ events $\rightarrow \tau$ lepton $\hookrightarrow \nu_\tau$

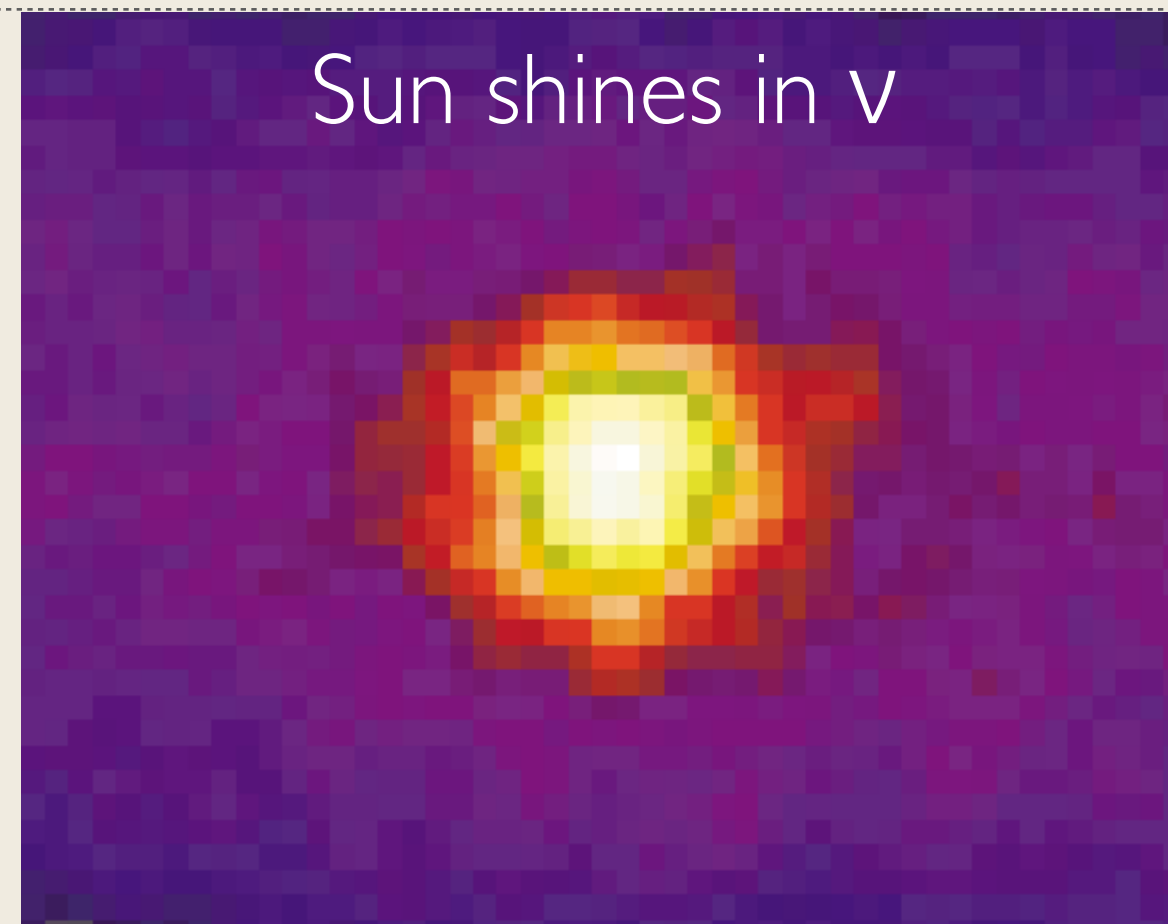


DONuT, 2000



Homestake

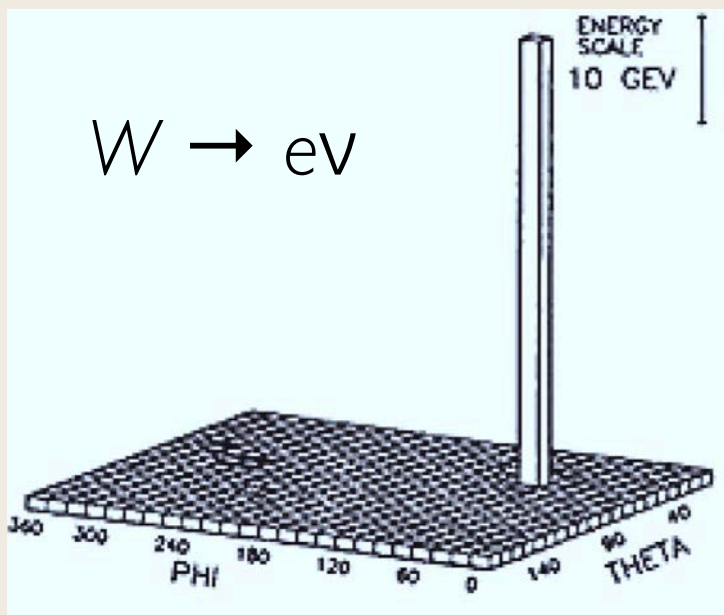
Sun shines in ν



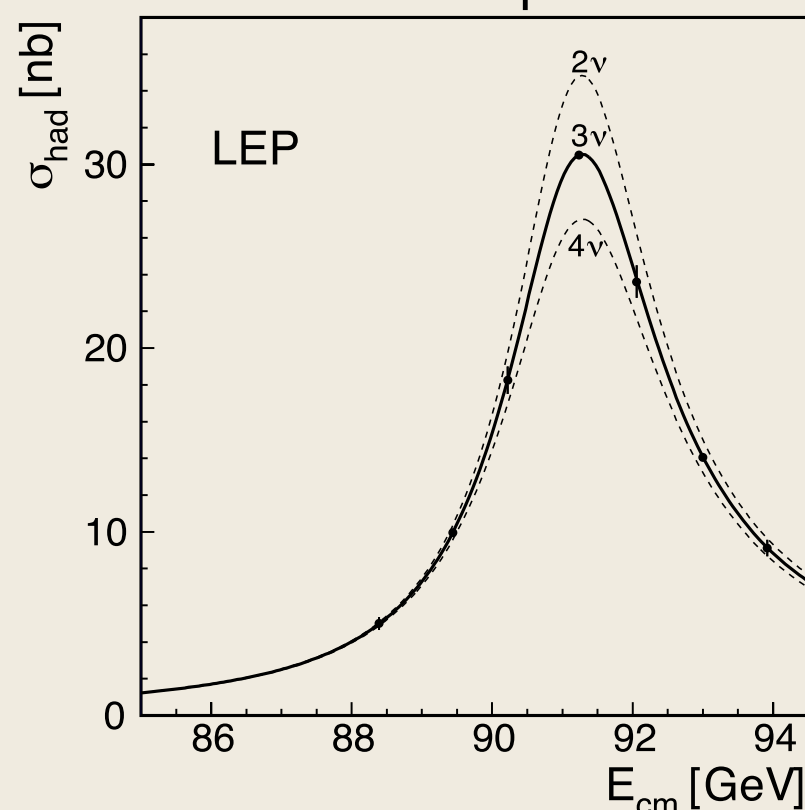
Super-Kamiokande

W, Z discovery

$\rightarrow 3 \nu$ species

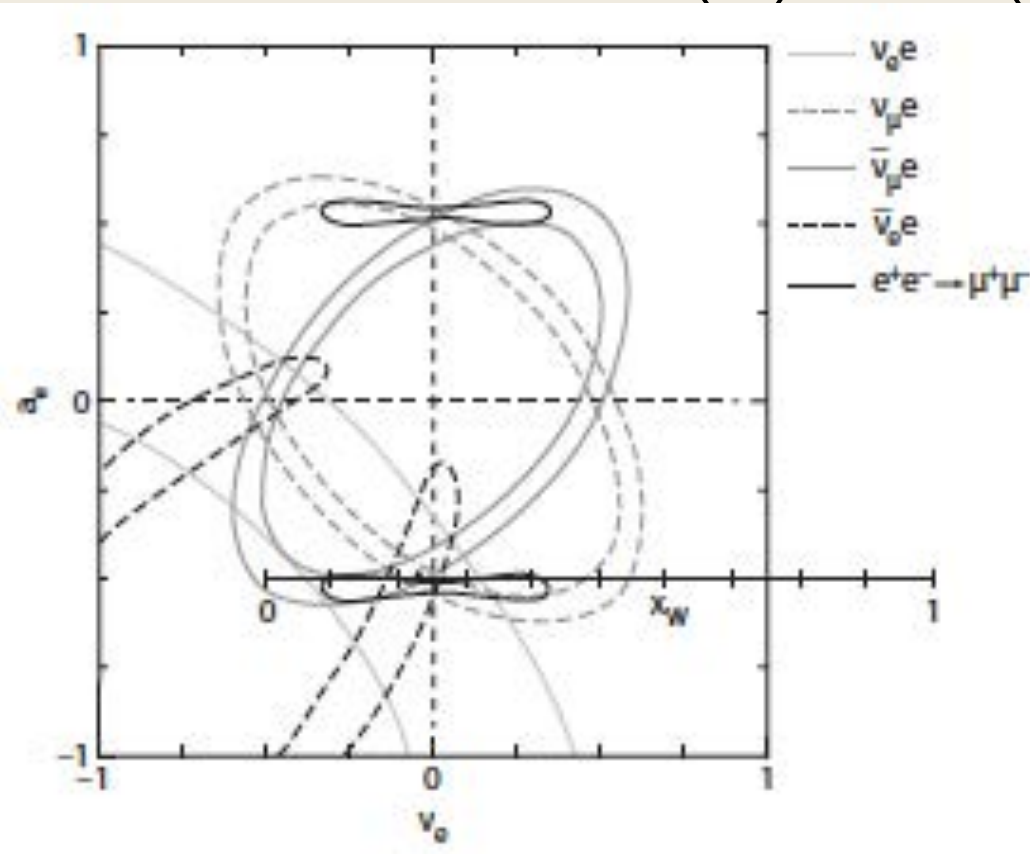


UA1/UA2, 1982–3



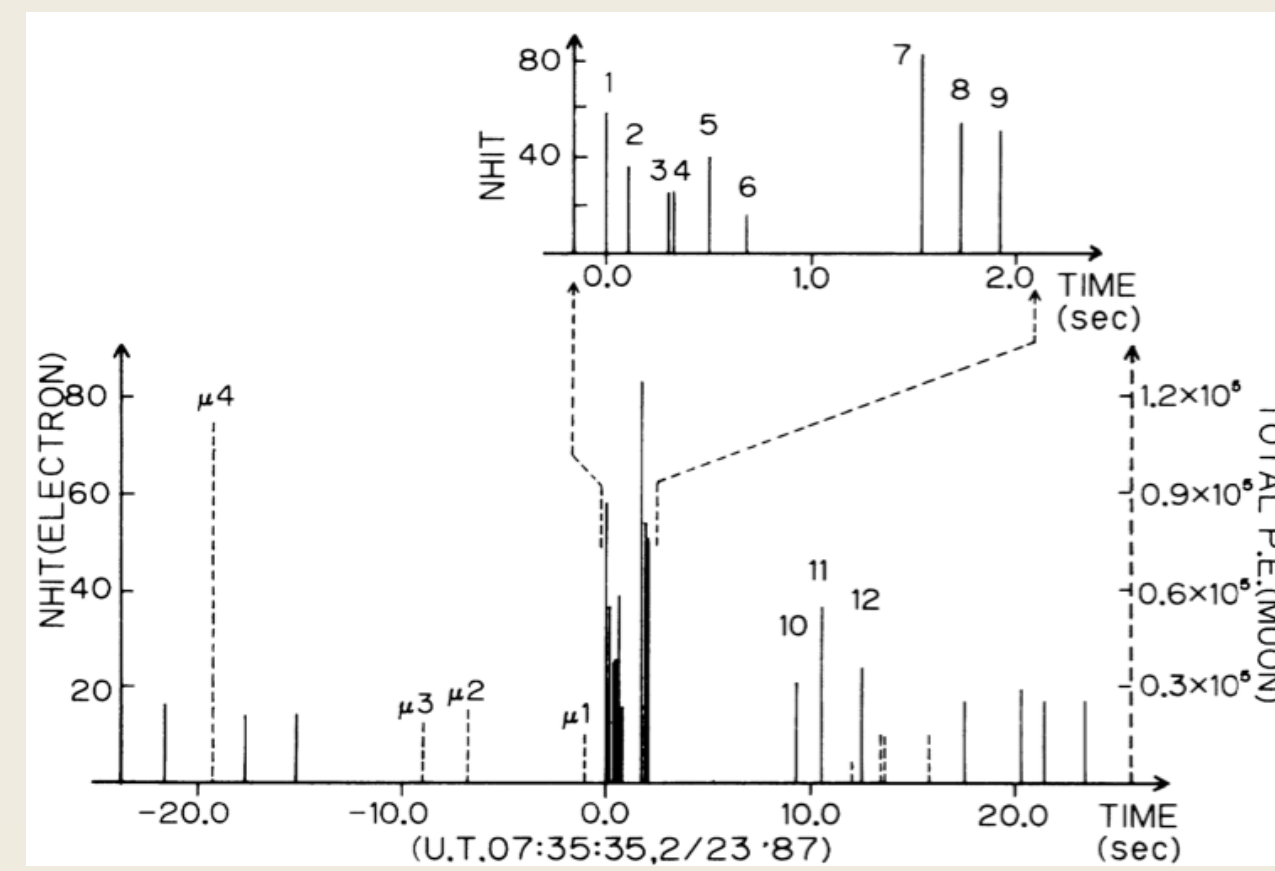
LEP, 1989–

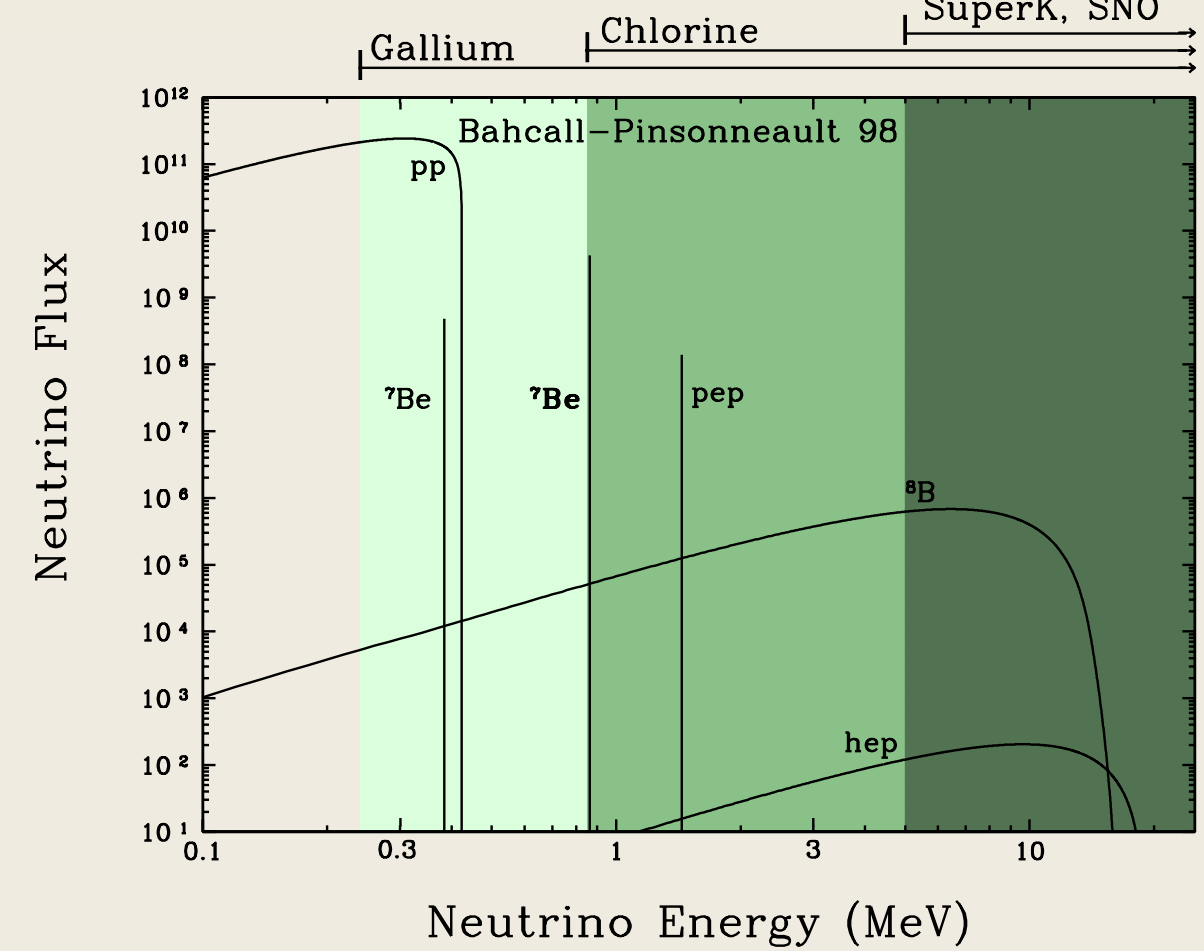
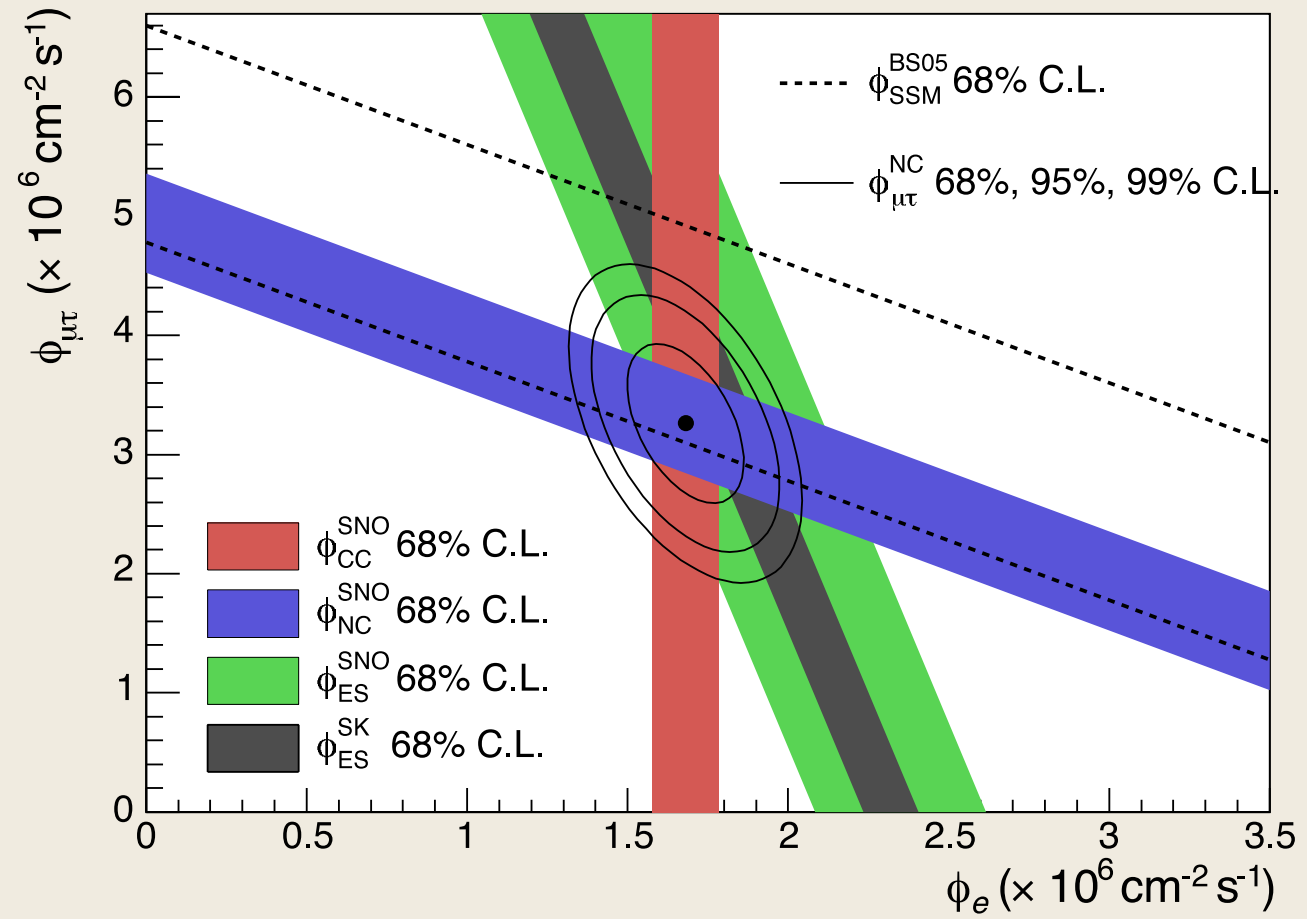
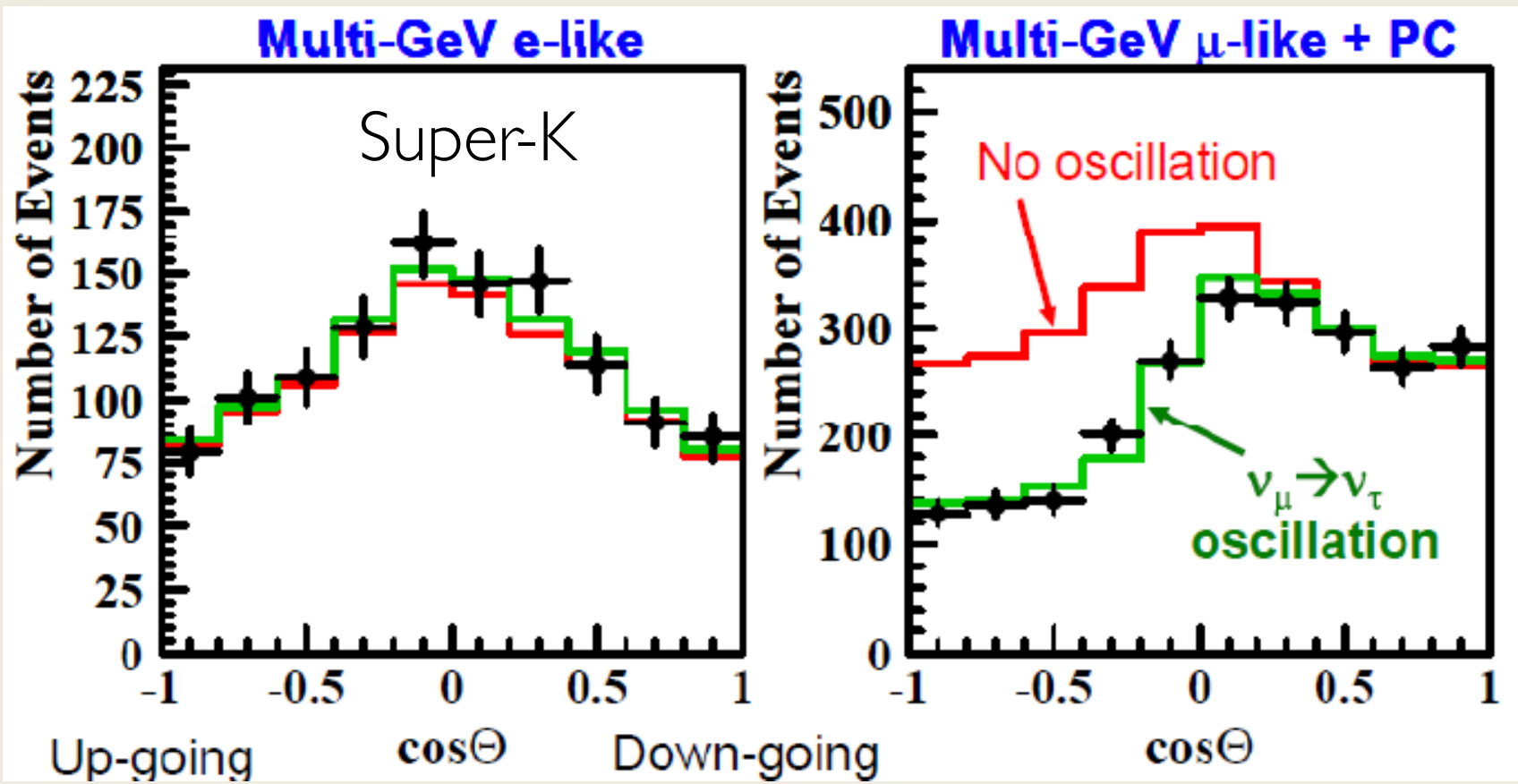
Leptonic σ favor $SU(2)_L \otimes U(1)_Y$



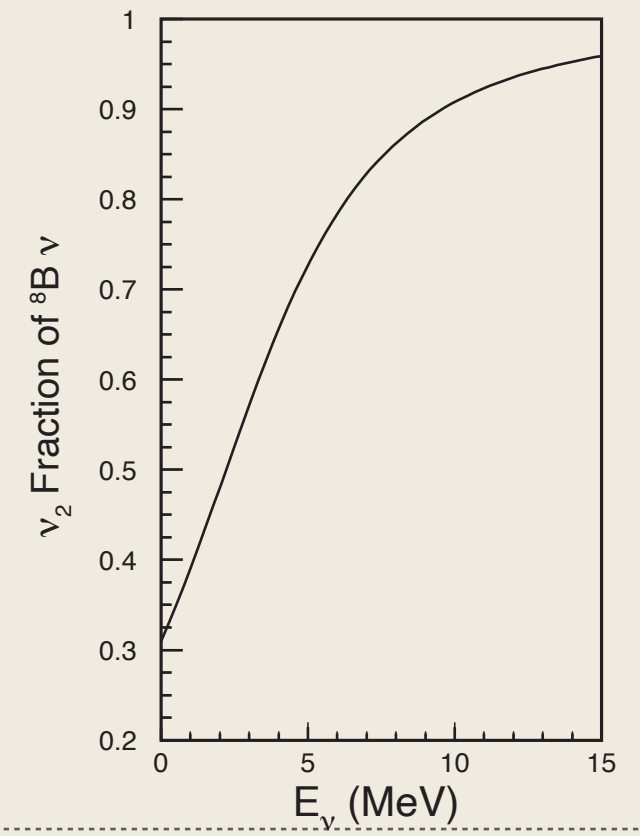
ca. 1987

SN1987a ν burst



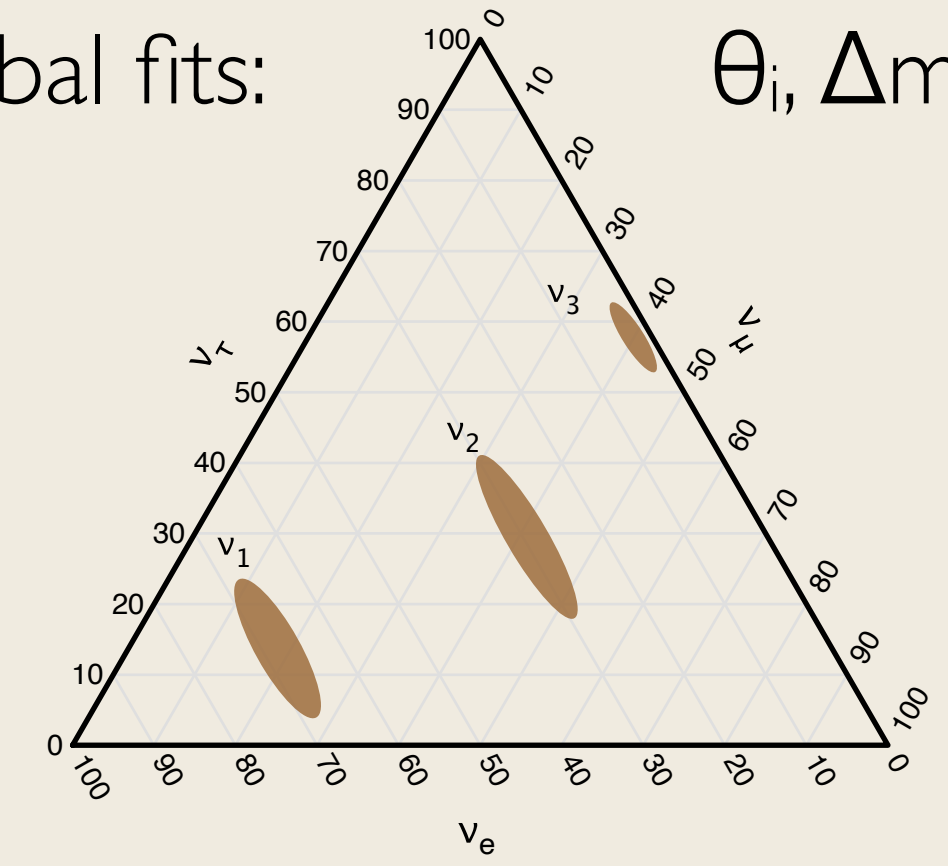


Matter-modulated oscillation (MSW)

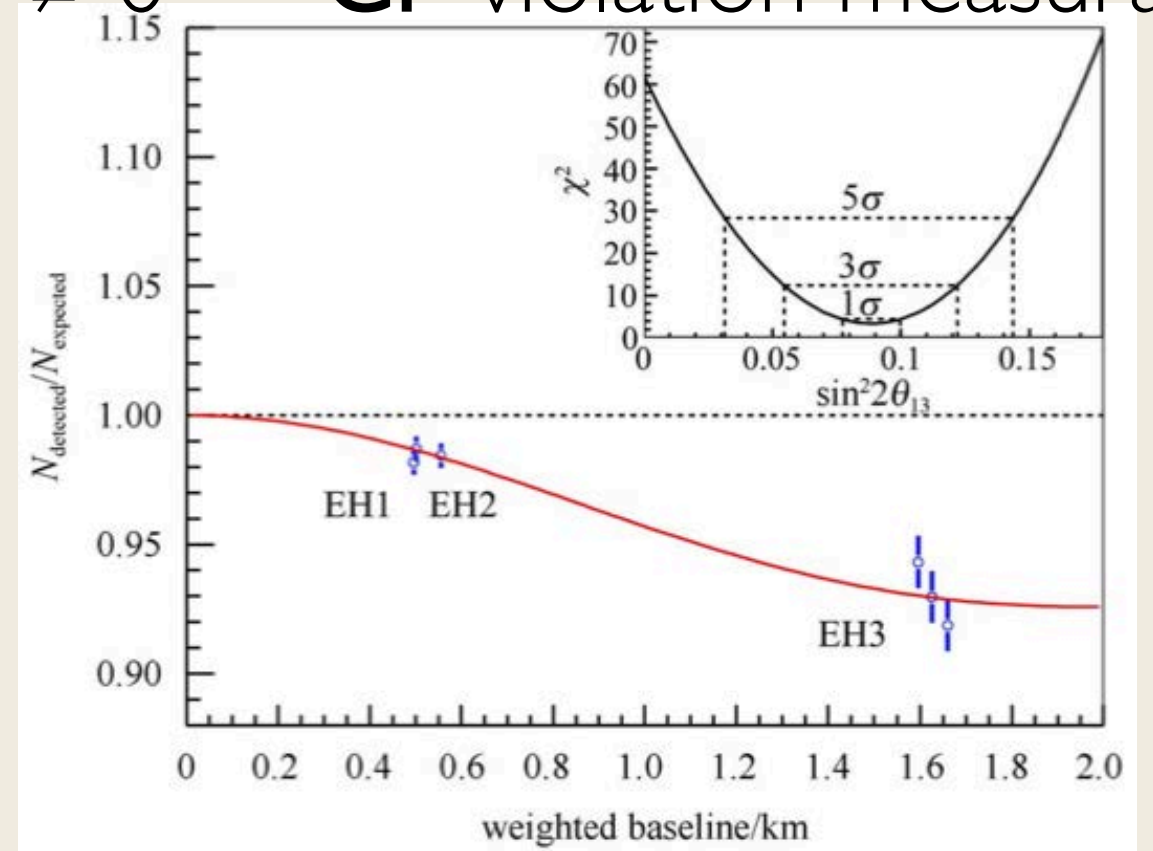


$P(\nu_e:\nu_2)$

Global fits: $\theta_i, \Delta m_{ij}^2, \delta_{CP}$



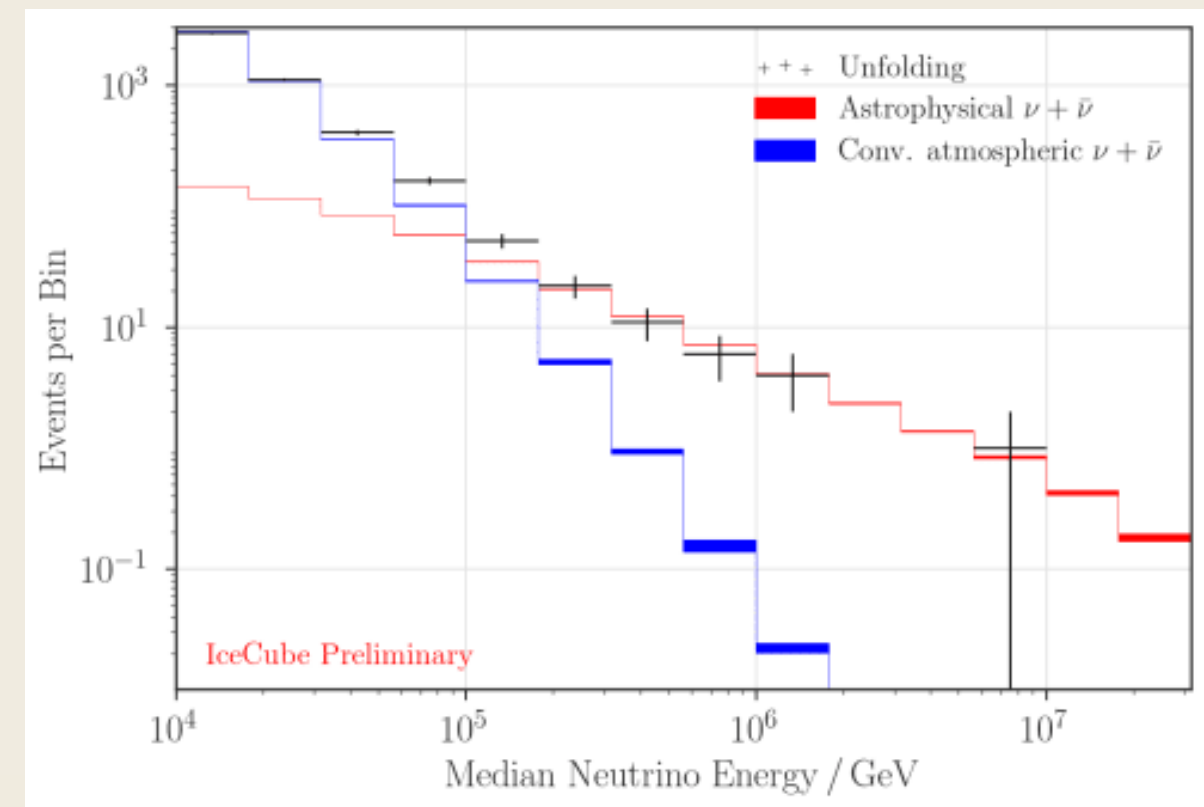
$\theta_{13} \neq 0 \hookrightarrow$ CP violation measurable



Ultra-rare decays

1987: $\nu\nu\beta\beta$ $^{82}\text{Se}, T_{1/2} \approx 10^{20} \text{ y}$
 2022: $(2\nu\text{ECEC})$ $^{124}\text{Xe}, T_{1/2} \approx 10^{22} \text{ y}$
 $0\nu\beta\beta$ decay not yet observed
 at level of $T_{1/2} \approx 10^{25} \text{ y}$
 $[\tau_p > 10^{31}-10^{33} \text{ y}]$

Extraterrestrial ν



IceCube

ν mass constraints

Cosmological arguments:
 $\Sigma_i m_{\nu_i} \lesssim (0.10-0.26) \text{ eV}$
 $^3\text{H} \beta$ decay:
 $m_{\nu_\beta} < 0.8 \text{ eV, 90\% CL}$

Daya Bay, RENO, ...

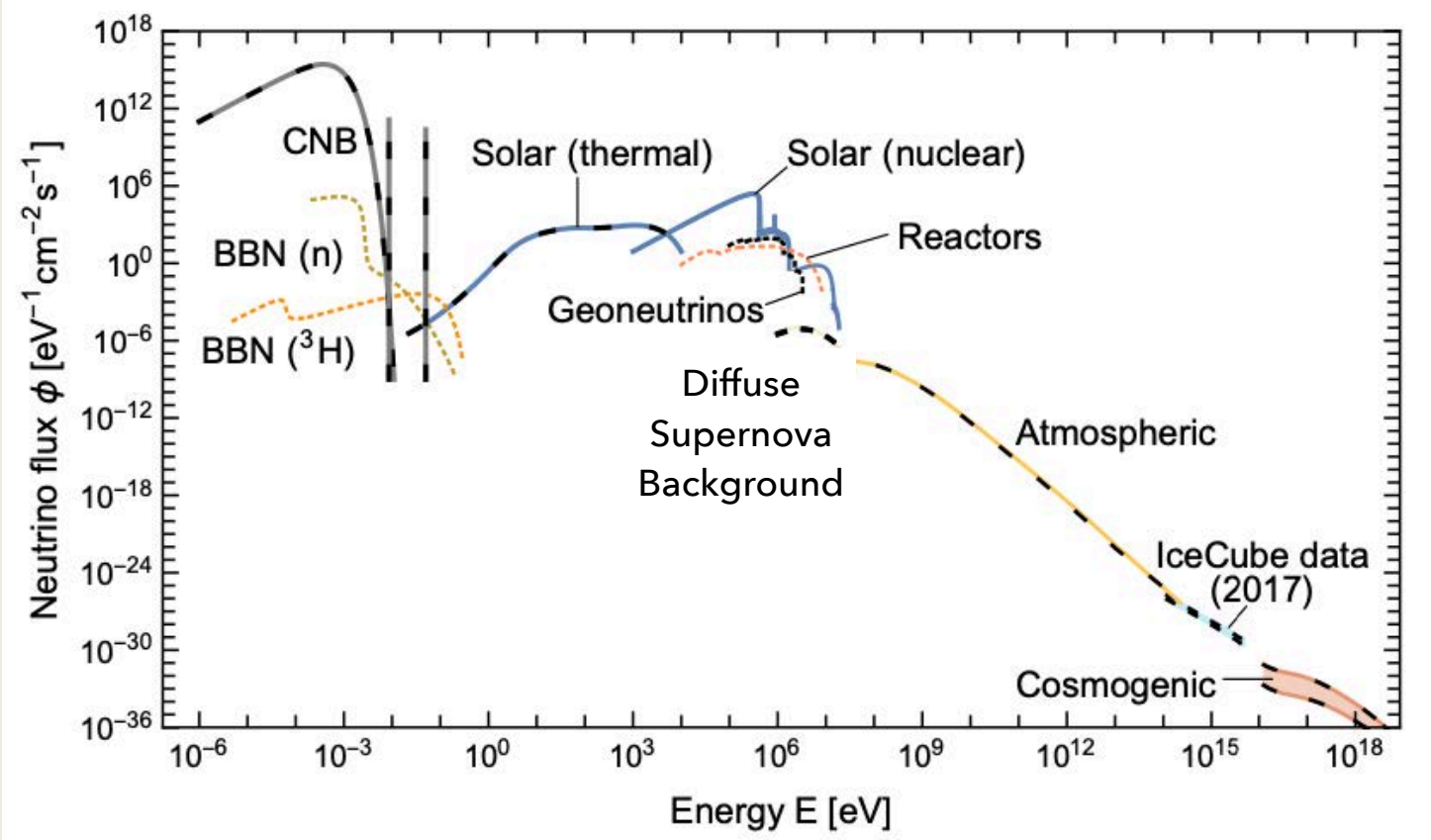
KATRIN, ...

Neutrinos through cosmological history

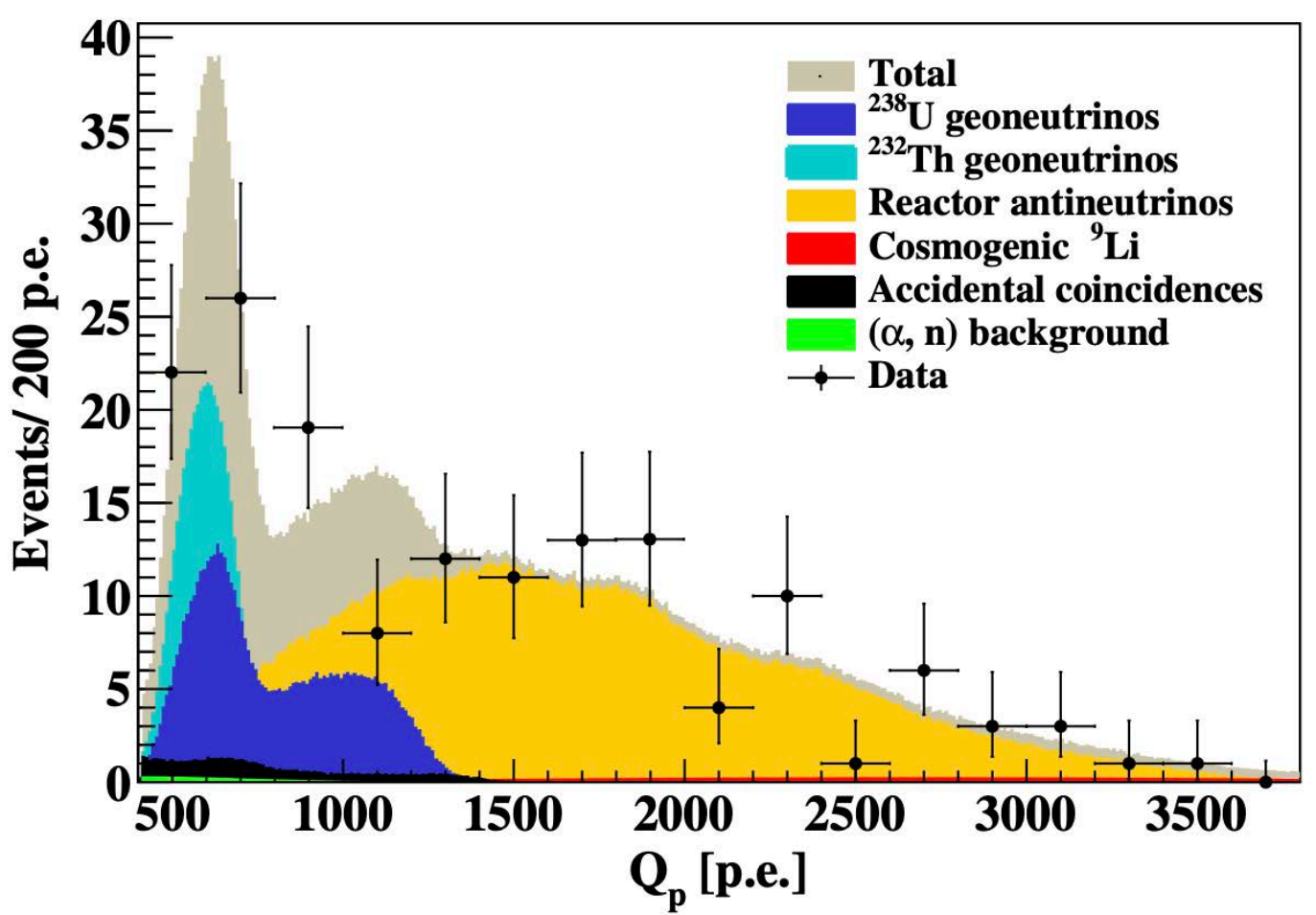
Relic neutrinos present at
Big-bang nucleosynthesis (few minutes),
Decoupling era (380 ky),
Large-scale structure formation (few %)

WMAP, Planck, ...

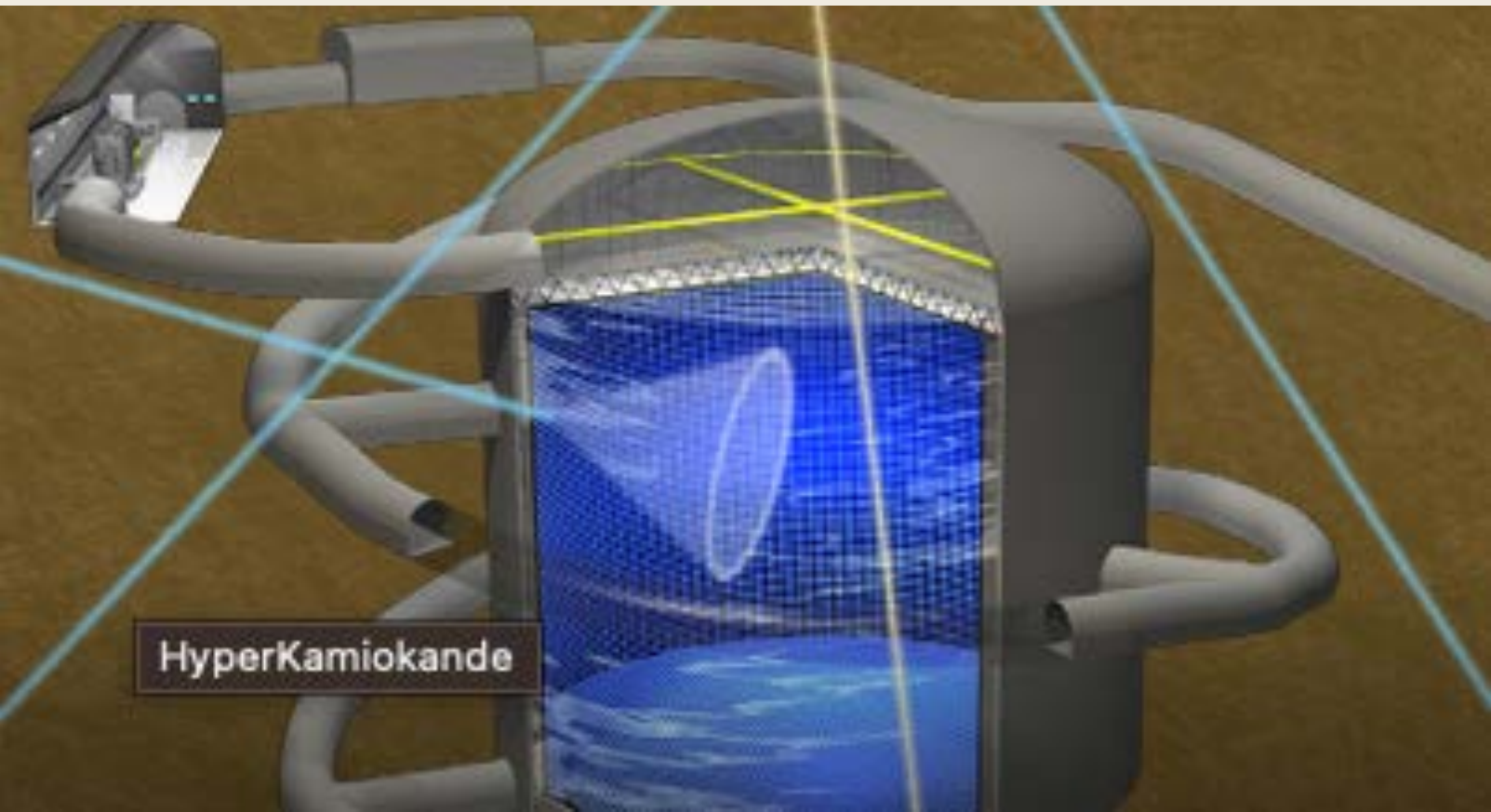
“Grand Unified v Spectrum”



Vitagliano +, RMP 92, 45006



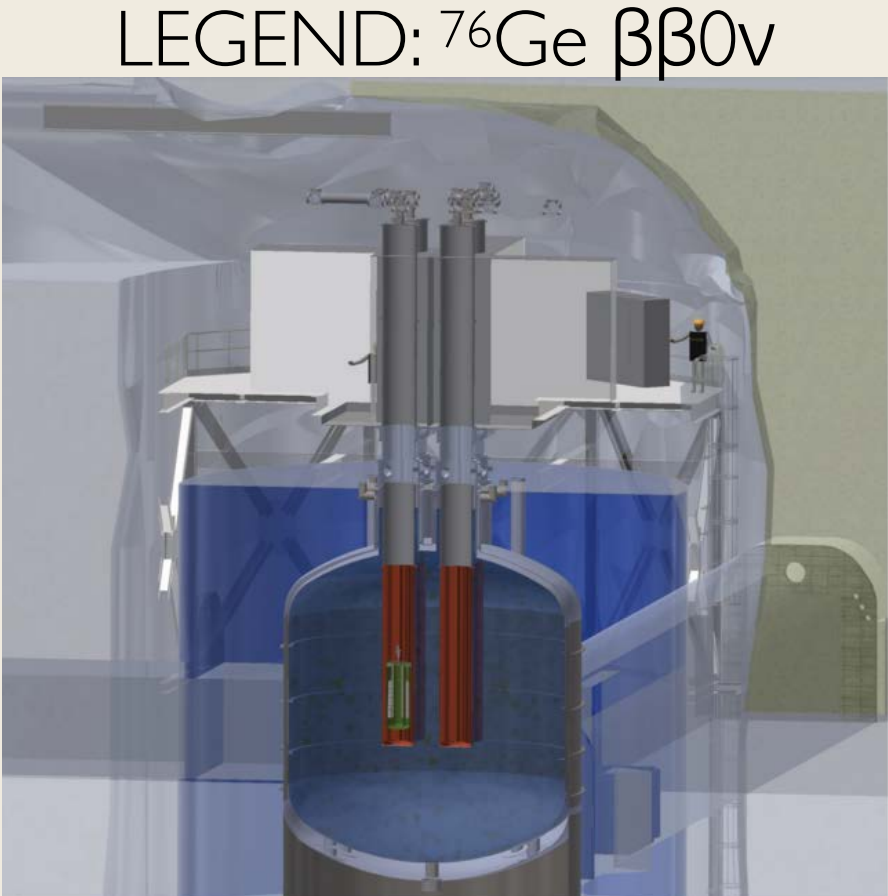
Borexino



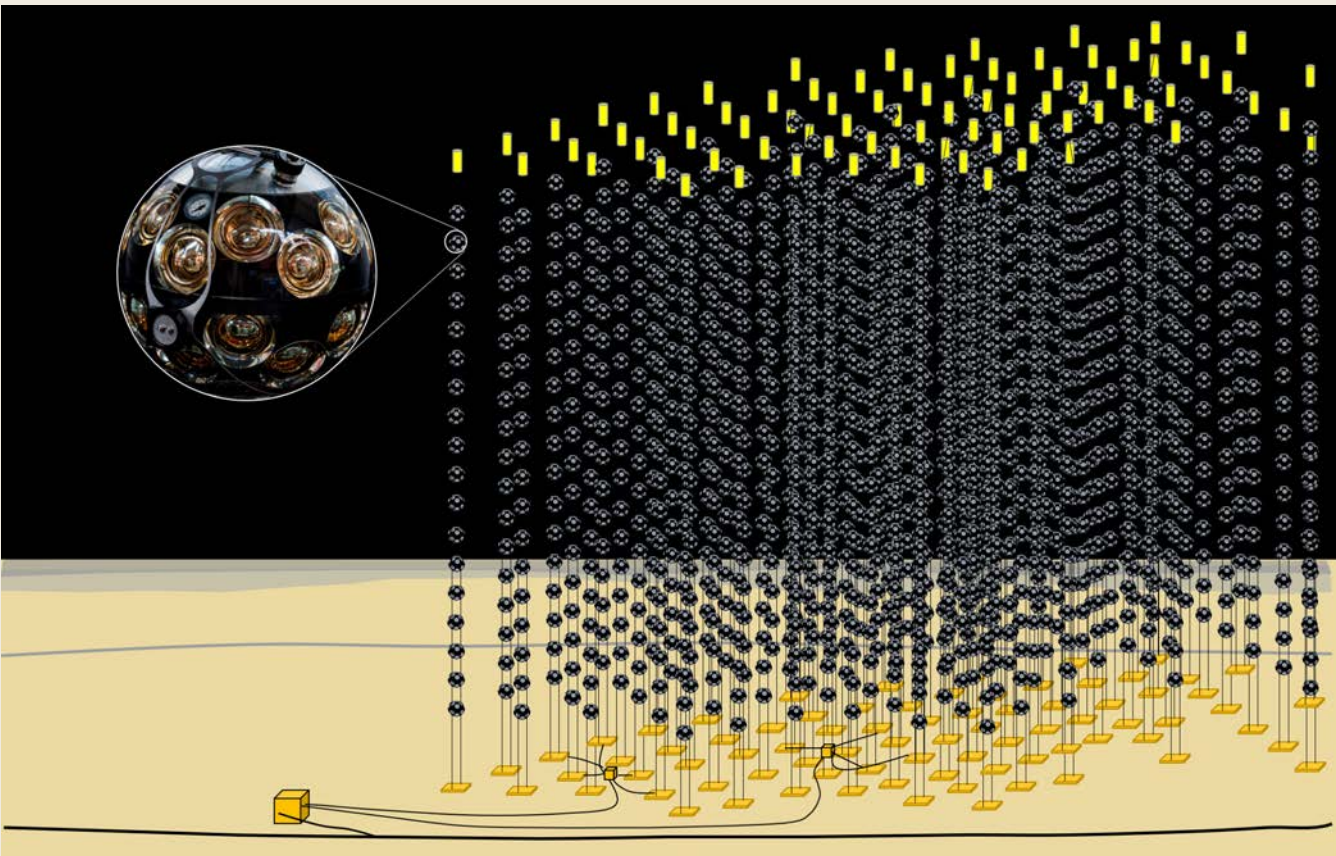
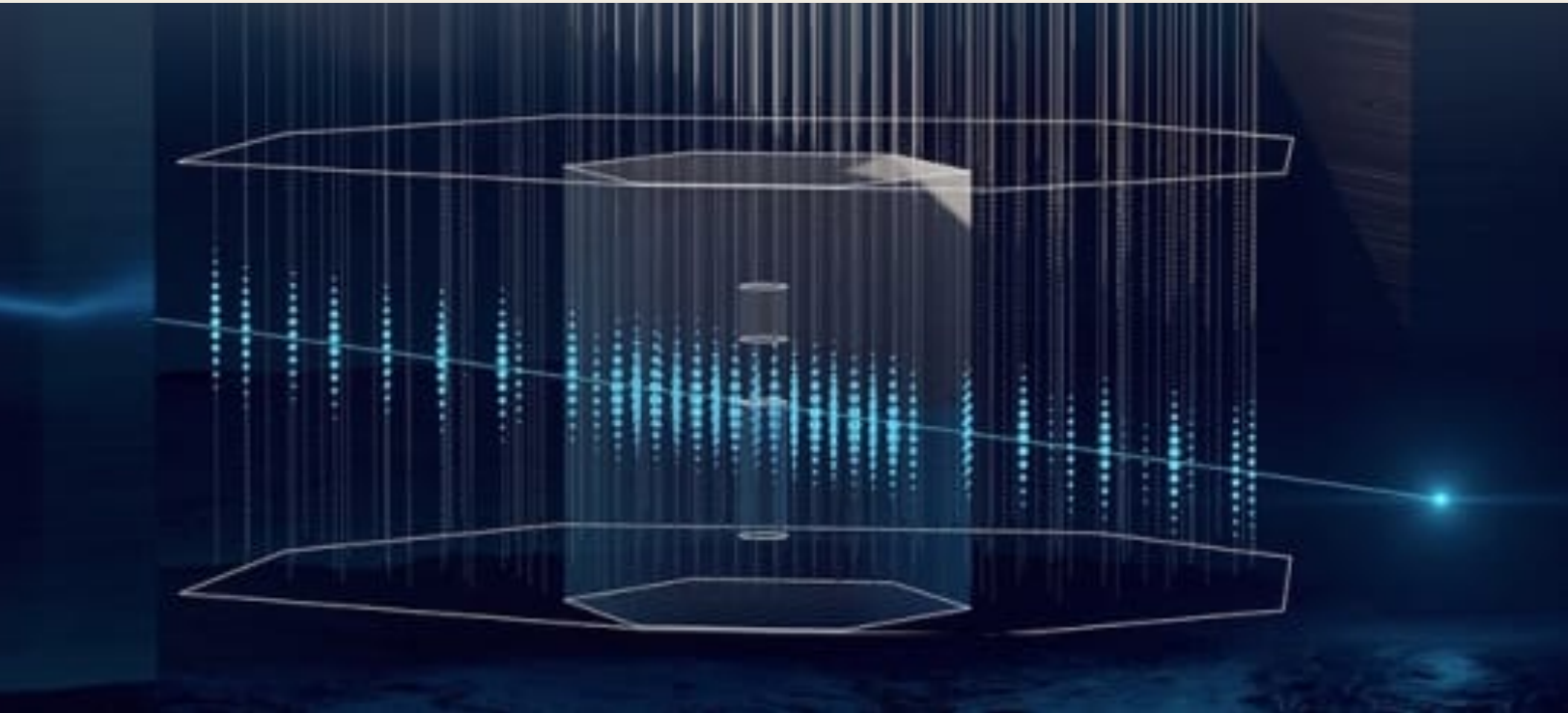
Ice-Cube-Gen2



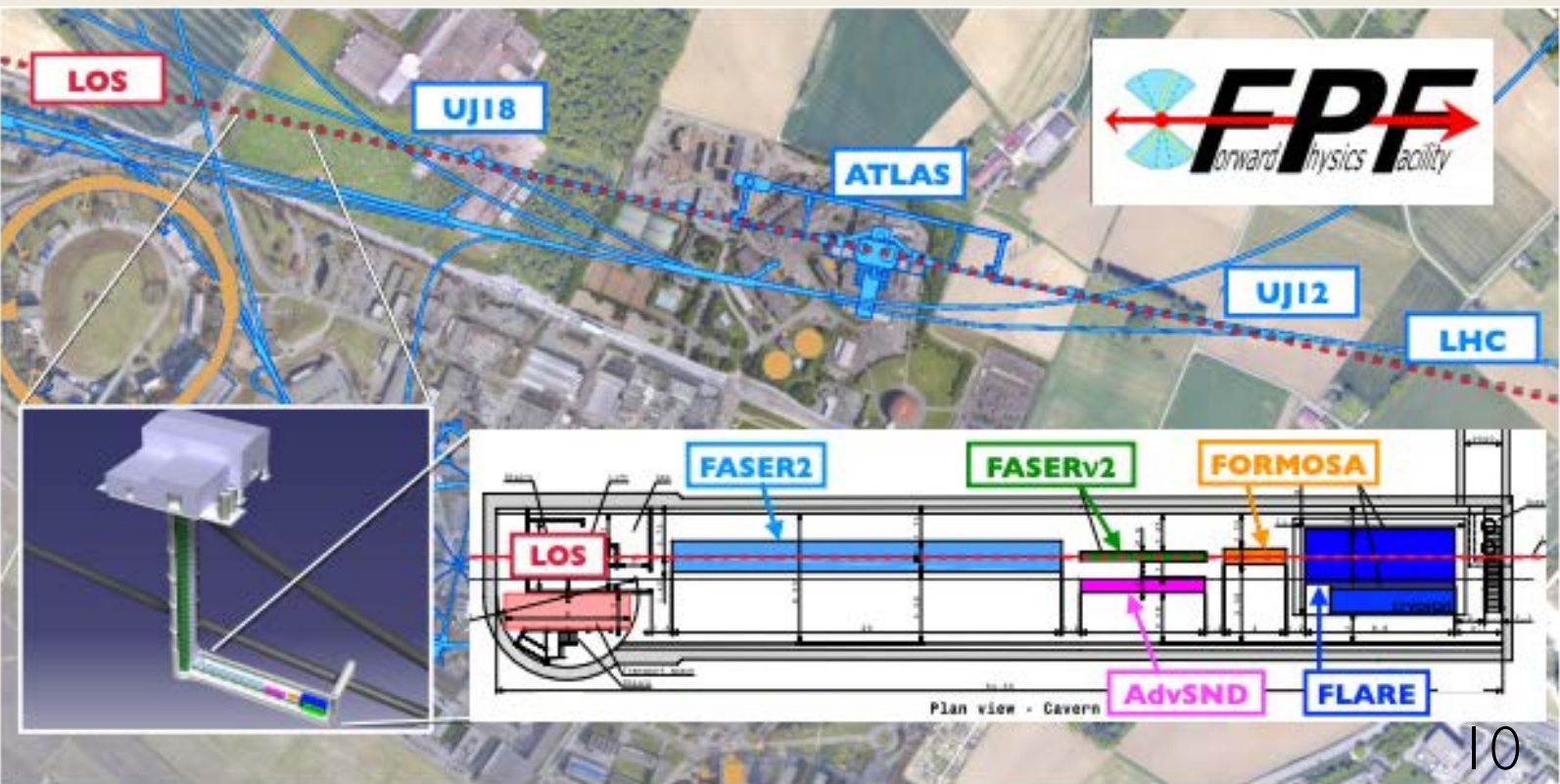
DUNE



LEGEND: ⁷⁶Ge ββ_{0ν}



KM3NeT



Forward Physics Facility for LHC

Some issues approaching v2022

Why does the neutrino weigh? At what scale are m_ν set?

What is the order of the levels ν_1, ν_2, ν_3 ? Absolute scale?

How is $m_\nu \neq 0$ a sign of BSM physics? Is Higgs field implicated?

How much (how little) do ν contribute to the dark matter?

Can we find a link between ν and dark matter?

Do neutrinos have nonstandard interactions, not mediated by W and Z ?

Can we observe electromagnetic properties of ν ?

Can we detect right-handed charged-current interactions?

What is the nature of right-handed ν ? Are there light sterile ν ? **noble**

Do 3 light (LH) ν suffice?

Some issues approaching $\nu 2022^{\text{bis}}$

Are neutrinos Majorana particles?

Are ν_1, ν_2, ν_3 stable on cosmological time scales?
How can we detect the cosmic neutrino background?

Is **CP** violated in ν oscillations? How does it arise?
Will ν yield insight into the matter excess in the universe?

What can we learn from the next supernova?

Some issues approaching v2022^{ter}

Are the interactions of ν_e, ν_μ, ν_τ universal?

What about the interactions of charged leptons?

Can we detect charged-lepton flavor violation?

What is its relation to neutrino mixing?

Need to resolve outstanding anomalies:
neutron lifetime, LSND/MiniBooNE, reactor flux, W mass,
LHCb hints of departures from lepton universality.

What have we overlooked? What do we know that is not true?

A Preview: Highlights of v2072

Long-baseline experiments will establish oscillation parameters with admirable precision, but real predictions are still lacking.

The study of leptonic **CP** violation will take on a life of its own; the connection to leptogenesis remains psychological.

The answer to the matter asymmetry will be “none of the above.”

ν observatories will locate point sources, including transients.

Flavor identification will characterize flavor mix at steady sources.

A neutrino factory will enable precise ν_e / ν_μ comparisons and revolutionize nucleon femtoscopy using polarized and active targets.

A meta-analysis will elucidate LSND/Mini-BooNE as a cocktail.

Large-scale experiments will hint that some neutrinos decay.

A Preview: Highlights of $\nu 2072^{\text{bis}}$

Systematics of $2\nu\beta\beta$ will provide new insights into nuclear dynamics.

Intense high-energy beams of ν_τ will suggest subtle differences with ν_e , ν_μ .

$0\nu\beta\beta$ will be observed for several isotopes.

Someone in today's audience will connect quark and neutrino mixing, but debate persists: are particle parameters deeply meaningful or environmental?

Neutrino cosmology will precisely measure light degrees of freedom and the sum of neutrino masses; ^3H β -decay will yield a larger m_ν , calling into question the canonical history of the universe.

Relic neutrinos will be detected after prodigious effort, and “soon” will be studied in undergraduate labs.

Detection of cosmogenic (GZK) neutrinos will fix birth of first stars.

A Preview: Highlights of $\nu 2072^{\text{ter}}$

Neutrino tomography of Earth's interior and widespread detection of geoneutrinos (+ lunar counterpart) draws planetary scientists to $\nu 2072$.

Extensive studies of CNO ν will inform solar physics and test understanding of matter-modulated flavor change.

Collider experiments will find multi-TeV neutral leptons.

Neutrino observatories will complement multi-messenger astronomy, enriching understanding of the Sun, supernovae, and cataclysmic mergers.

RH charged-current interactions, W -bosons suggest new paths to unification.

We will learn to use undersea fiber-optic networks and ubiquitous environmental surveillance infrastructure to study ν .

Proton decay will be observed in several modes.

Thanks to v'22 Organizers & Participants,
and to my v Collaborators, Advisors, Teachers

Carl Albright, Gabriela Barenboim, John Beacom,
Marcela Carena, Debajyoti Choudhury, Gene Commins,
Raj Gandhi, Dave Jackson, Joachim Kopp, Ben Lee, Magda Lola,
Pedro Machado, Olga Mena, Irina Mocioiu, Stephen Parke,
Mary Hall Reno, Ina Sarcevic, Robert Shrock, Jack Smith, Terry Walker,
and many inspiring experimental colleagues.

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NEUTRINO '72

EUROPHYSICS CONFERENCE

BALATONFÜRED, HUNGARY,

11-17 JUNE 1972

3x2
organized by

THE HUNGARIAN PHYSICAL SOCIETY

PROCEEDINGS VOLUME I.

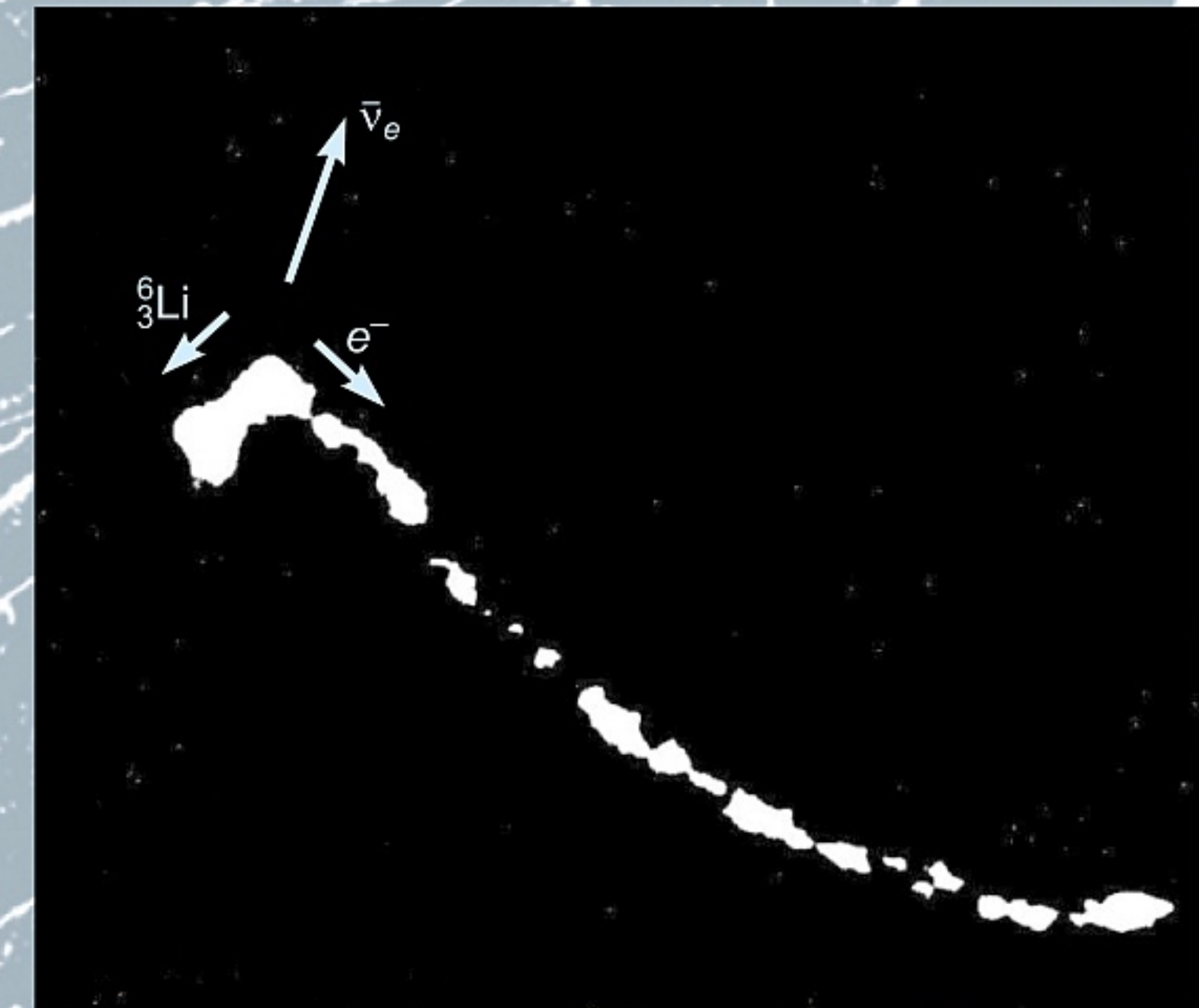
A. FRENKEL
G. MARX Editors

Also see J. Schneps, "Brief history of 'Neutrino', the International Conference on Neutrino Physics and Astrophysics," in *Neutrino 2014*.

OMKDK-TECHNOINFORM

fizikai szemle

HUNGARIAN PHYSICAL REVIEW



2022 KÜLÖNSZÁM
SPECIAL ISSUE

nka

Special Issue to commemorate the 50th anniversary of v '72